

Time-Dependent Density Functional Theory for the Inner Crust of Neutron Stars

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Hierarchy of Scales in the Universe



Hierarchy of Scales in the Universe

Neutron stars, NS merger, nucleosynthesis, GW, ...



Macroscopic

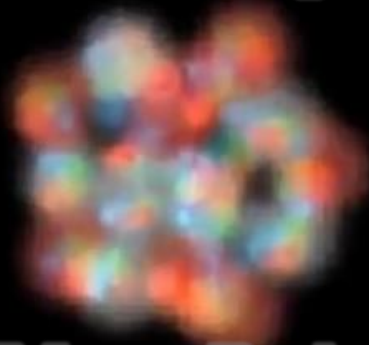


(Nuclear)Astrophysics

Nuclear structure, Equation of State (EoS)
Superfluidity & Superconductivity
Reaction rates, Fission fragments, ...

Neutron-star structure, Star quakes, GW
Pulsar glitches, Cooling
Stellar evolution, Nucleosynthesis, ...

Microscopic



Nuclear Many-Body Problem



10 fm

Hierarchy of Scales in the Universe

Neutron stars, NS merger, nucleosynthesis, GW, ...



Macroscopic



(Nuclear)Astrophysics

Our Mission:

Nuclear structure, Fermi
Superfluid
Reaction rates

To establish a concrete microscopic foundation of macroscopic models

Neutron stars, Star quakes, GW
Cooling
Nucleosynthesis, ...



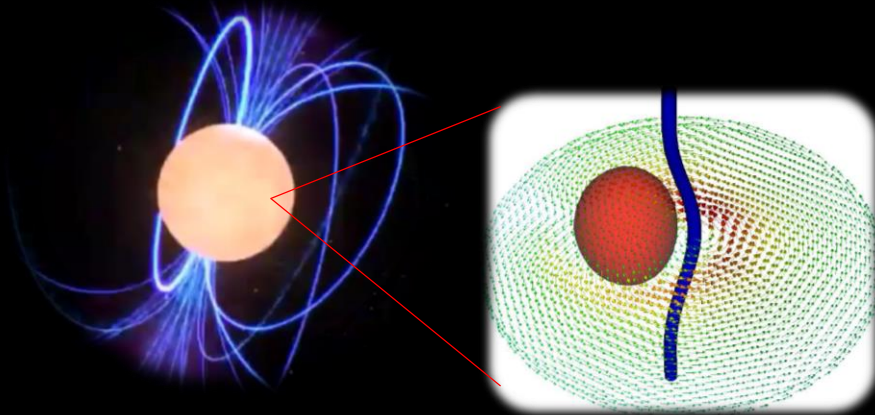
Nuclear Many-Body Problem



10 fm

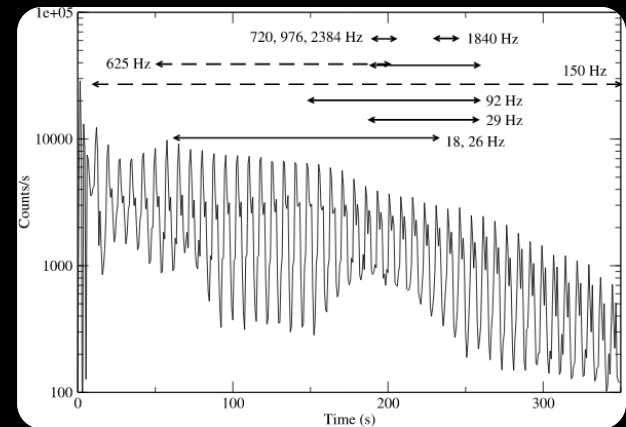
Today, I will talk about:

Dynamics of quantum vortices of superfluid neutrons



↔ Neutron-star glitches

Time-dependent band theory for the inner crust of neutron stars



↔ Quasi-periodic oscillations

■ (TD)DFT in a tiny nutshell



A theory which gives us access to the *exact* solution

Equivalent!
(for a special EDF)

$$\hat{H}\Psi(\mathbf{r}_1, \dots, \mathbf{r}_N) = E\Psi(\mathbf{r}_1, \dots, \mathbf{r}_N)$$

Kohn-Sham equation

$$\left[-\frac{\hbar^2}{2m} \nabla^2 + v_{\text{KS}}[\rho(\mathbf{r})] \right] \phi_i(\mathbf{r}) = \varepsilon_i \phi_i(\mathbf{r})$$

$$v_{\text{KS}}[\rho(\mathbf{r})] = \frac{\delta \mathcal{E}[\rho]}{\delta \rho} \quad \rho(\mathbf{r}) = \sum_{i=1}^N |\phi_i(\mathbf{r})|^2$$

EDF

This is the key!

Quantum Many-Body Problem



Energy can also be written as a functional of density

$$E[\rho] = \langle \Psi[\rho] | \hat{H} | \Psi[\rho] \rangle$$

w.f. is a functional of density

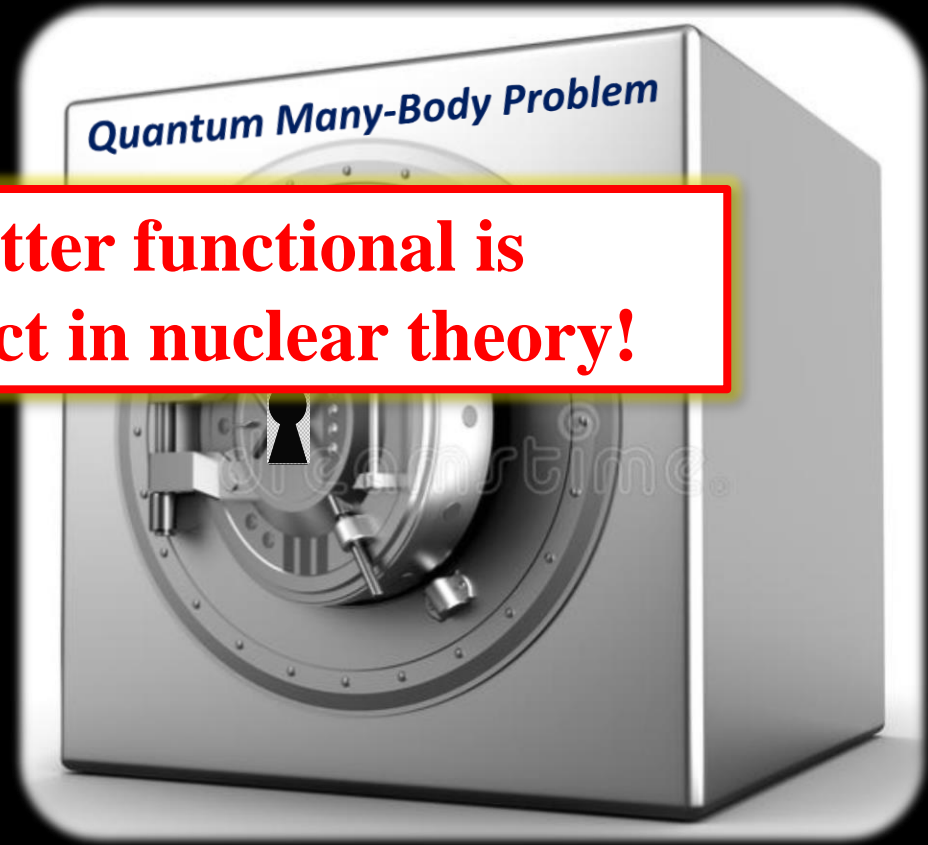
P. Hohenberg and W. Kohn, Phys. Rev. B **136**, 864 (1964)

CAUTION!

The existence was proven, but its shape is unknown..

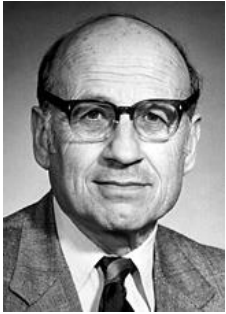
“Inverse Kohn-Sham”

**Developing a better functional is
an important subject in nuclear theory!**

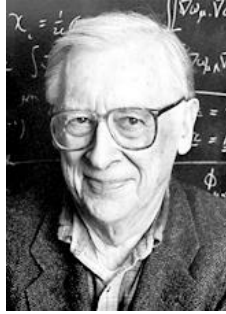


Great Success of the Density Functional Theory

The Nobel Prize in Chemistry 1998



Walter Kohn

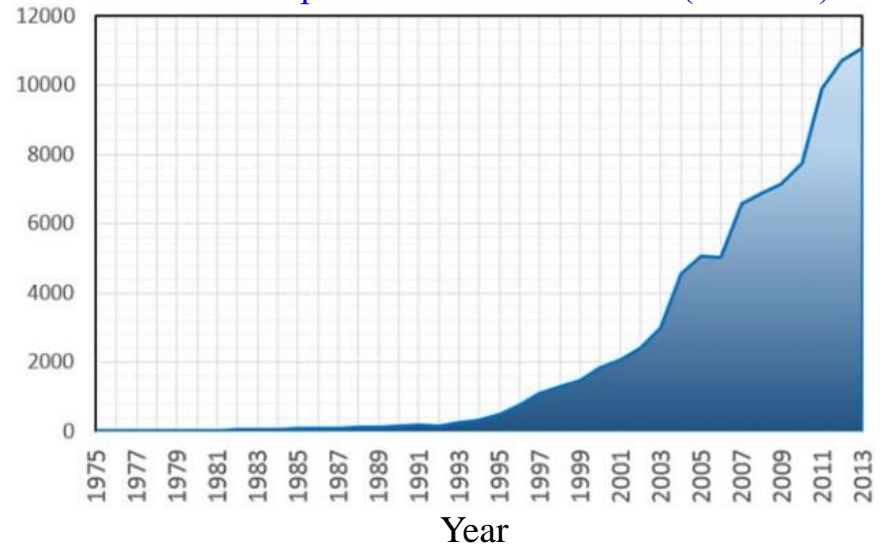


John Pople



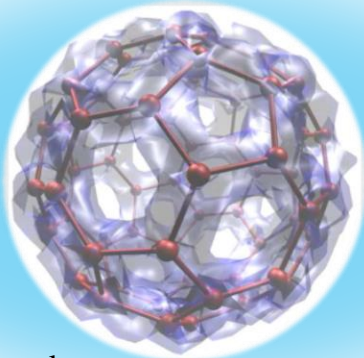
©<https://www.nobelprize.org>

Number of publications with “DFT” (till 2013)



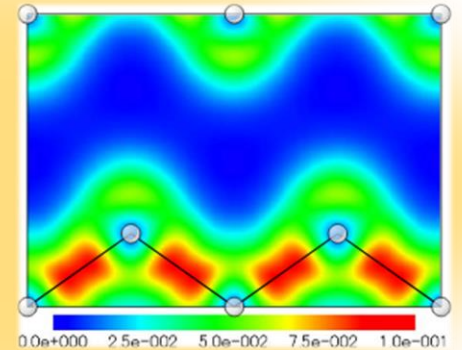
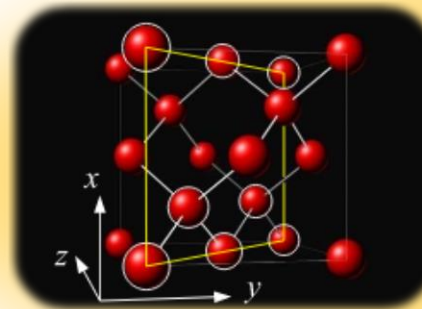
A. Galano and J.R. Alvarez-Idadoy, *J. Compt. Chem.* **35**, 2019 (2014)

Fullerene: C₆₀



C-Z. Gao et al.,
J. Phys. B: At. Mol. Opt. Phys. **48**, 105102 (2015)

Si crystal

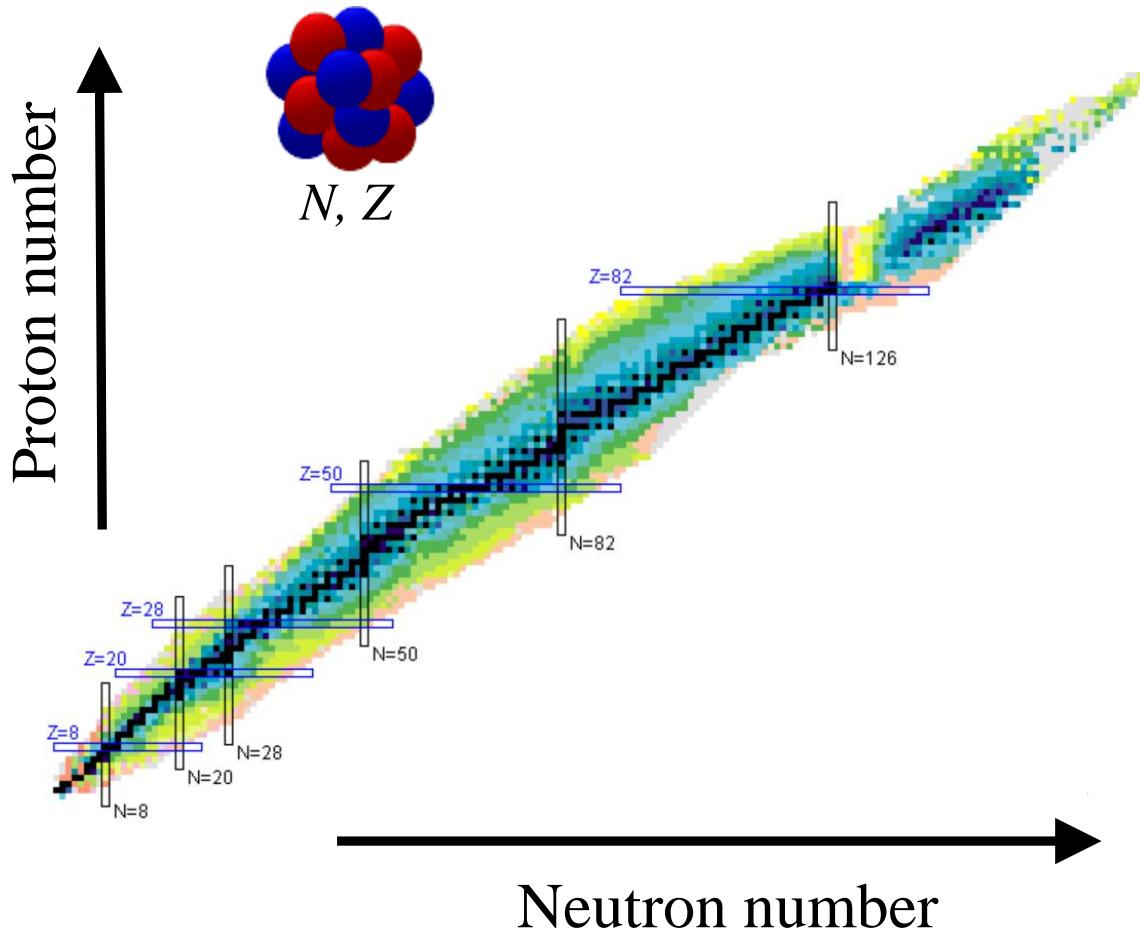


Y. Shinohara, K. Yabana, Y. Kawashita, J.-I. Iwata, T. Otobe, and G. F. Bertsch,
Phys. Rev. B **82**, 155110 (2010)

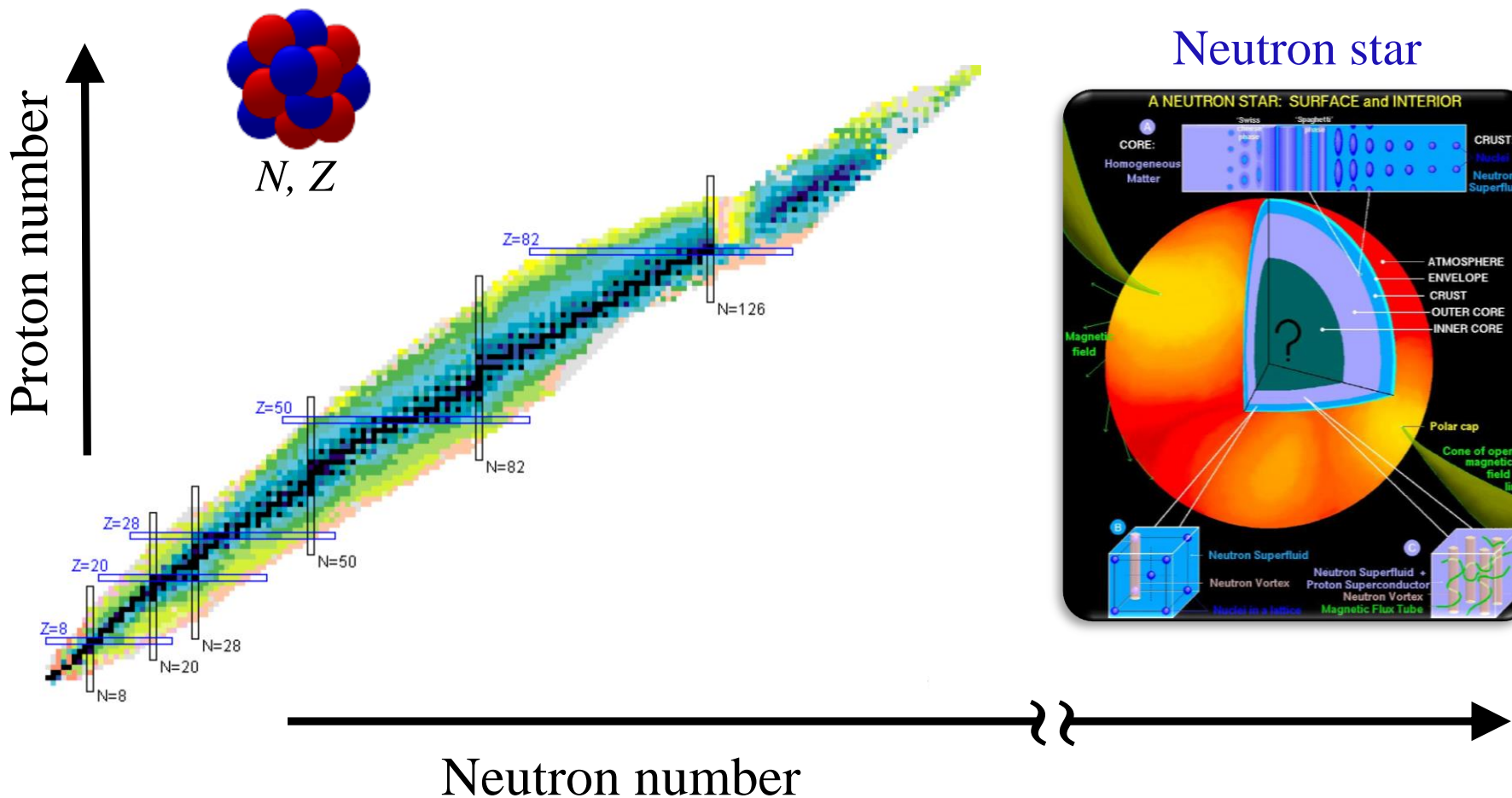
The seminal papers on DFT

- P. Hohenberg and W. Kohn, *Phys. Rev.* **136**, B864 (1964) ➔ **19,015 citations!**
- W. Kohn and L.J. Sham, *Phys. Rev.* **140**, A1133 (1965) ➔ **24,384 citations!**

All nuclei can be described with a *single* EDF

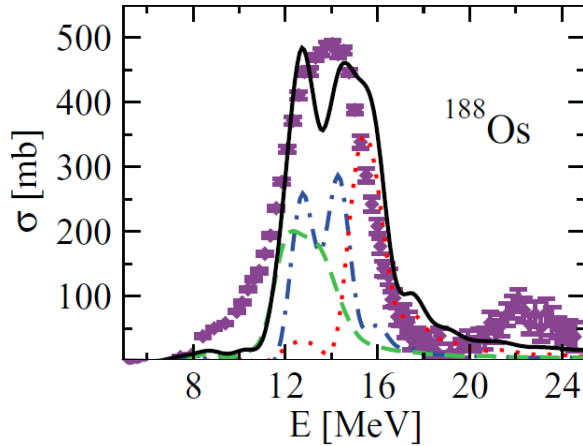


All nuclei can be described with a *single* EDF

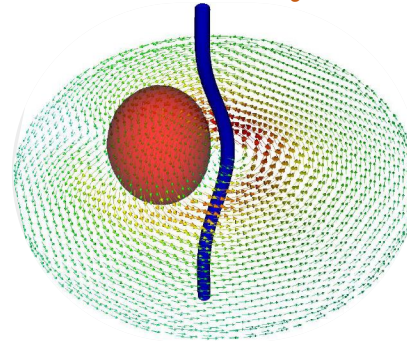


TDDFT is a versatile tool!!

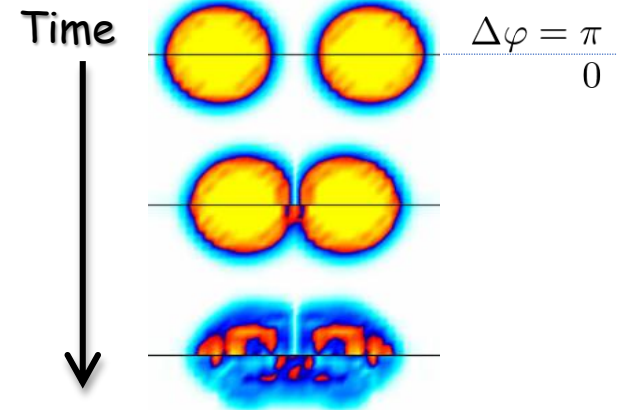
IVGDR



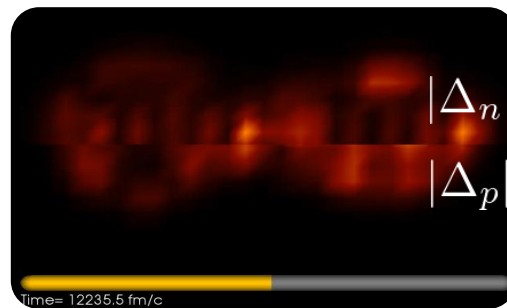
Vortex-nucleus dynamics



Low-energy heavy-ion reactions



Induced fission of ²⁴⁰Pu



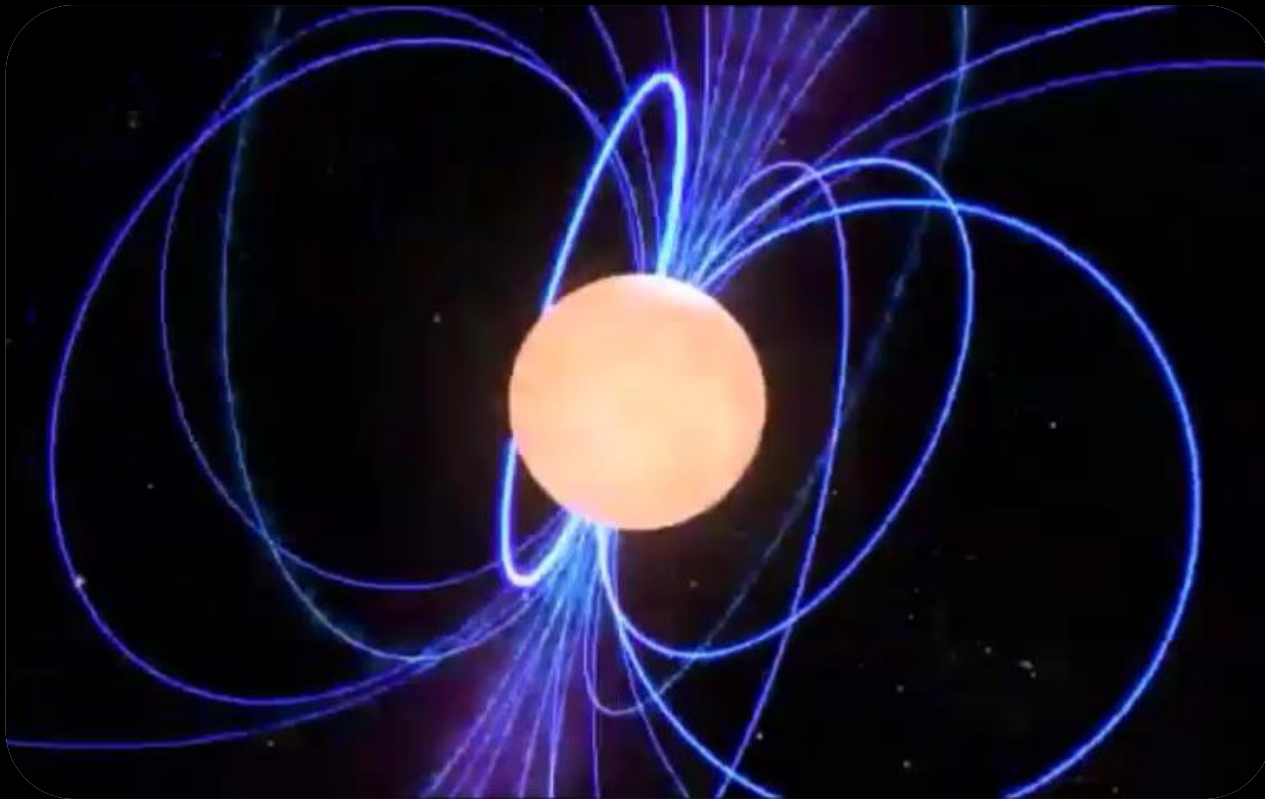
Neutron-star “glitch”



What is the glitch?

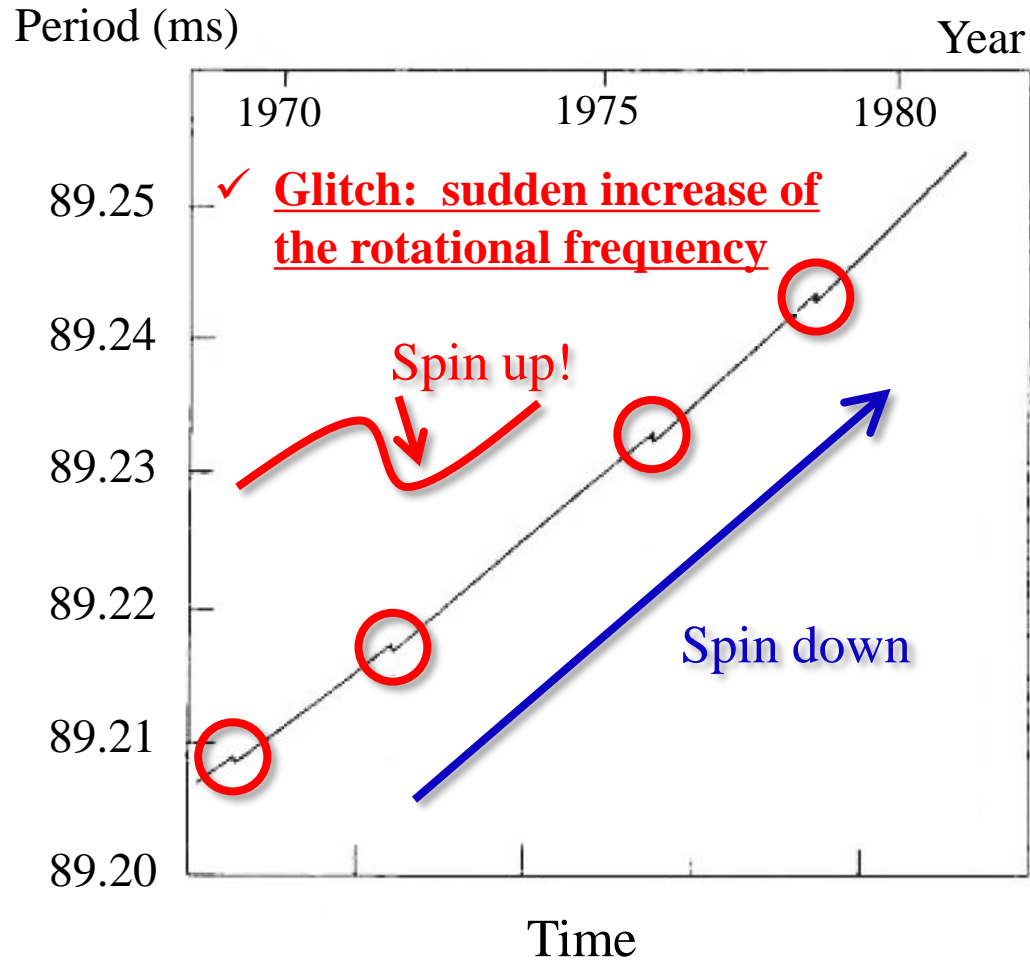
Pulsar - a rotating neutron star

- ✓ First discovery in August 1967 → “Little Green Man” LGM-1 → PSR B1919+21
- ✓ Since then, more than 2650 pulsars have been observed
- ✓ It gradually spins down due to the EM radiation



Typical example: the Vela pulsar

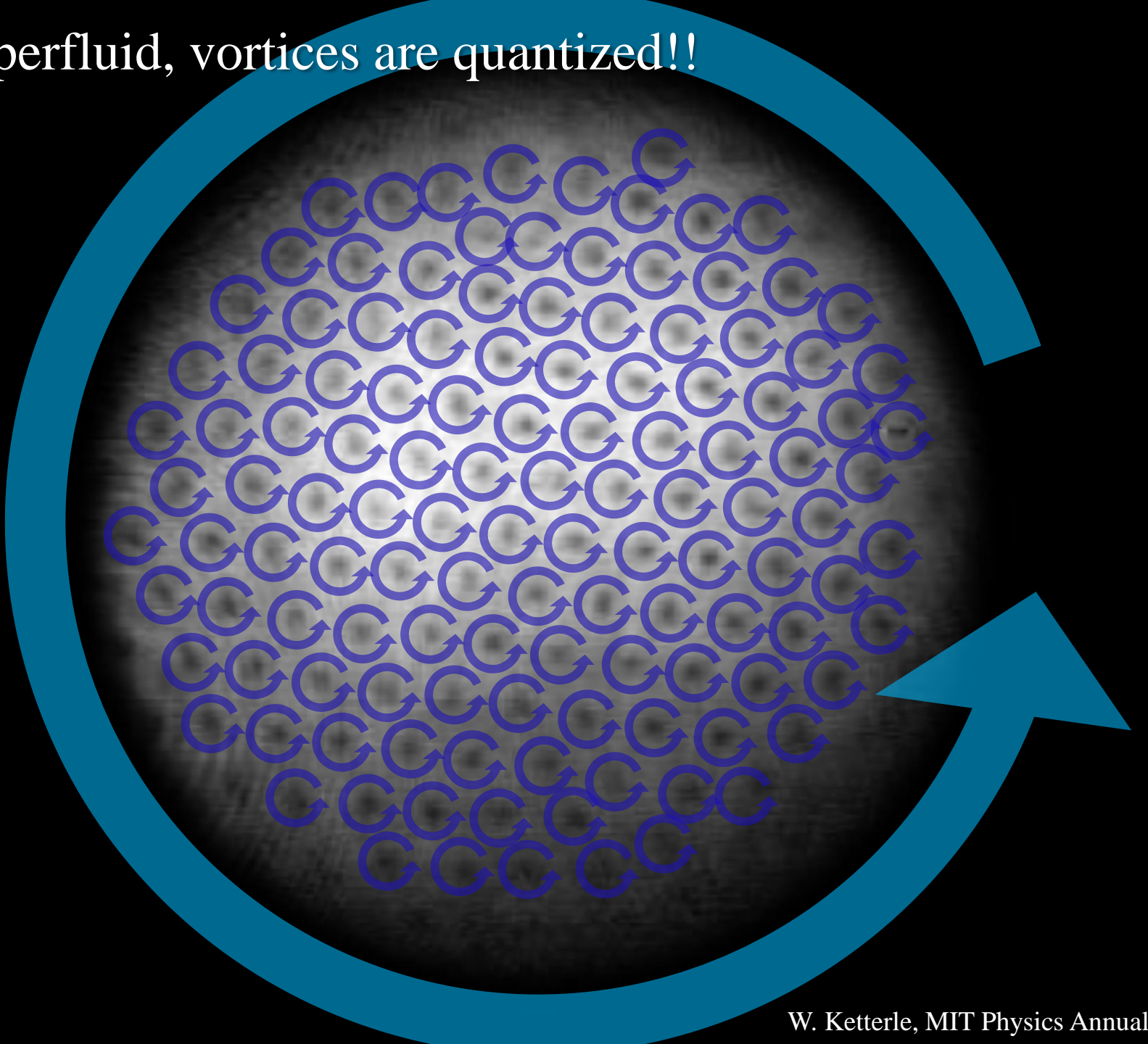
- *Irregularity* has been observed from continuous monitoring of the pulsation period



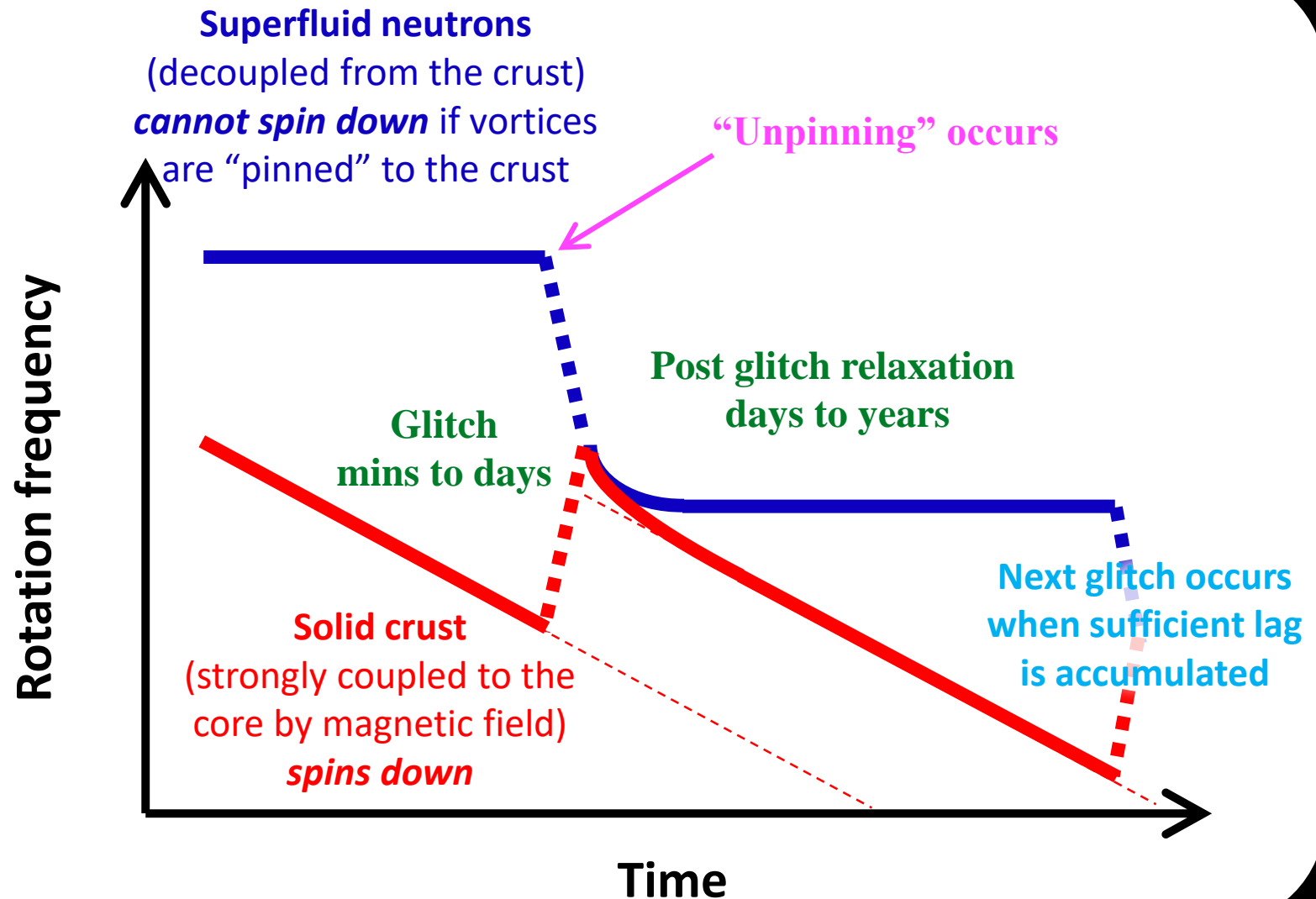
In daily life, a vortex is continuous..



In superfluid, vortices are quantized!!



The vortex mediated glitch: Naive picture



To fully understand the glitches, we need to clarify:

Glitch dynamics

and, of course,
details of NS matter..

How do vortices move?

Pinning mechanism

How are vortices pinned?

Trigger mechanism

How are vortices unpinned?

We attacked this problem using
the state-of-the-art microscopic nuclear theory

We attack this problem with HPC on GPU supercomputers
with TDDFT for superfluid systems, TDSLDA!

TDSLDA: TDDFT with local treatment of pairing

Kohn-Sham scheme is extended for non-interacting quasiparticles

➤ TDSLDA equations (formally equivalent to TDHFB or TD-BdG equations)

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_{k,\uparrow}(\mathbf{r}, t) \\ u_{k,\downarrow}(\mathbf{r}, t) \\ v_{k,\uparrow}(\mathbf{r}, t) \\ v_{k,\downarrow}(\mathbf{r}, t) \end{pmatrix} = \begin{pmatrix} h_{\uparrow\uparrow}(\mathbf{r}, t) & h_{\uparrow\downarrow}(\mathbf{r}, t) & 0 & \Delta(\mathbf{r}, t) \\ h_{\downarrow\uparrow}(\mathbf{r}, t) & h_{\downarrow\downarrow}(\mathbf{r}, t) & -\Delta(\mathbf{r}, t) & 0 \\ 0 & -\Delta^*(\mathbf{r}, t) & -h_{\uparrow\uparrow}^*(\mathbf{r}, t) & -h_{\uparrow\downarrow}^*(\mathbf{r}, t) \\ \Delta^*(\mathbf{r}, t) & 0 & -h_{\downarrow\uparrow}^*(\mathbf{r}, t) & -h_{\downarrow\downarrow}^*(\mathbf{r}, t) \end{pmatrix} \begin{pmatrix} u_{k,\uparrow}(\mathbf{r}, t) \\ u_{k,\downarrow}(\mathbf{r}, t) \\ v_{k,\uparrow}(\mathbf{r}, t) \\ v_{k,\downarrow}(\mathbf{r}, t) \end{pmatrix}$$

$$h_{\sigma} = \frac{\delta E}{\delta n_{\sigma}} \quad : \text{ s.p. Hamiltonian}$$

$$\Delta = -\frac{\delta E}{\delta \nu^*} \quad : \text{ pairing field}$$

$$n_{\sigma}(\mathbf{r}, t) = \sum_{E_k < E_c} |v_{k,\sigma}(\mathbf{r}, t)|^2 \quad : \text{ number density}$$

$$\nu(\mathbf{r}, t) = \sum_{E_k < E_c} u_{k,\uparrow}(\mathbf{r}, t) v_{k,\downarrow}^*(\mathbf{r}, t) \quad : \text{ anomalous density}$$

$$\mathbf{j}_{\sigma}(\mathbf{r}, t) = \hbar \sum_{E_k < E_c} \text{Im}[v_{k,\sigma}^*(\mathbf{r}, t) \nabla v_{k,\sigma}(\mathbf{r}, t)] \quad : \text{ current}$$

A large number (10^4 - 10^6) of 3D coupled non-linear PDEs have to be solved!!

of qp orbitals ~ # of grid points

TDSLDA: TDDFT with local treatment of pairing

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Supercomputing!!

$$h_{\sigma} = \frac{\delta E}{\delta n_{\sigma}} : \text{s.p. Hamiltonian}$$

$$n_{\sigma}(\mathbf{r}, t) = \sum_{E_k < E_c} |v_{k,\sigma}(\mathbf{r}, t)|^2 : \text{number density}$$

$$\nu(\mathbf{r}, t) = \sum_{E_k < E_c} u_{k,\uparrow}(\mathbf{r}, t) v_{k,\downarrow}^*(\mathbf{r}, t) : \text{anomalous density}$$

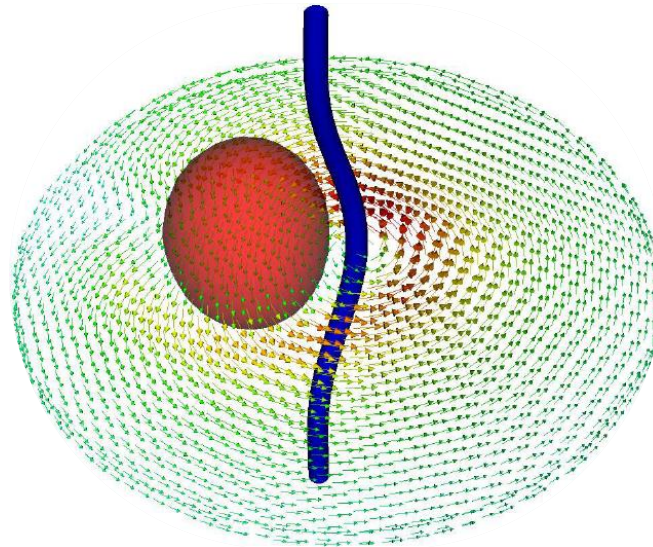
$$\Delta = -\frac{\delta E}{\delta \nu^*} : \text{pairing field}$$

$$\mathbf{j}_{\sigma}(\mathbf{r}, t) = \hbar \sum_{E_k < E_c} \text{Im}[v_{k,\sigma}^*(\mathbf{r}, t) \nabla v_{k,\sigma}(\mathbf{r}, t)] : \text{current}$$

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Vortex-nucleus dynamics within TDSLDA



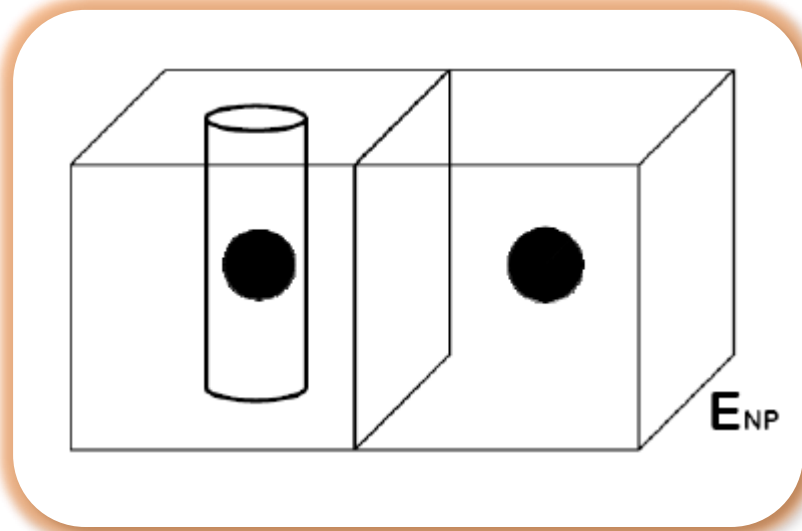
A key to understand the glitches is:
Vortex pinning mechanism in the inner crust of neutron stars

Q. Is the vortex-nucleus interaction

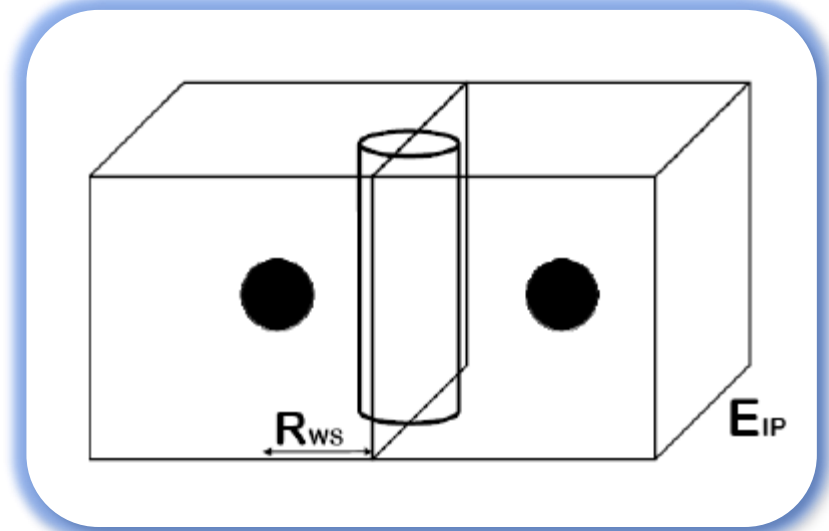
Attractive?

or

Repulsive?



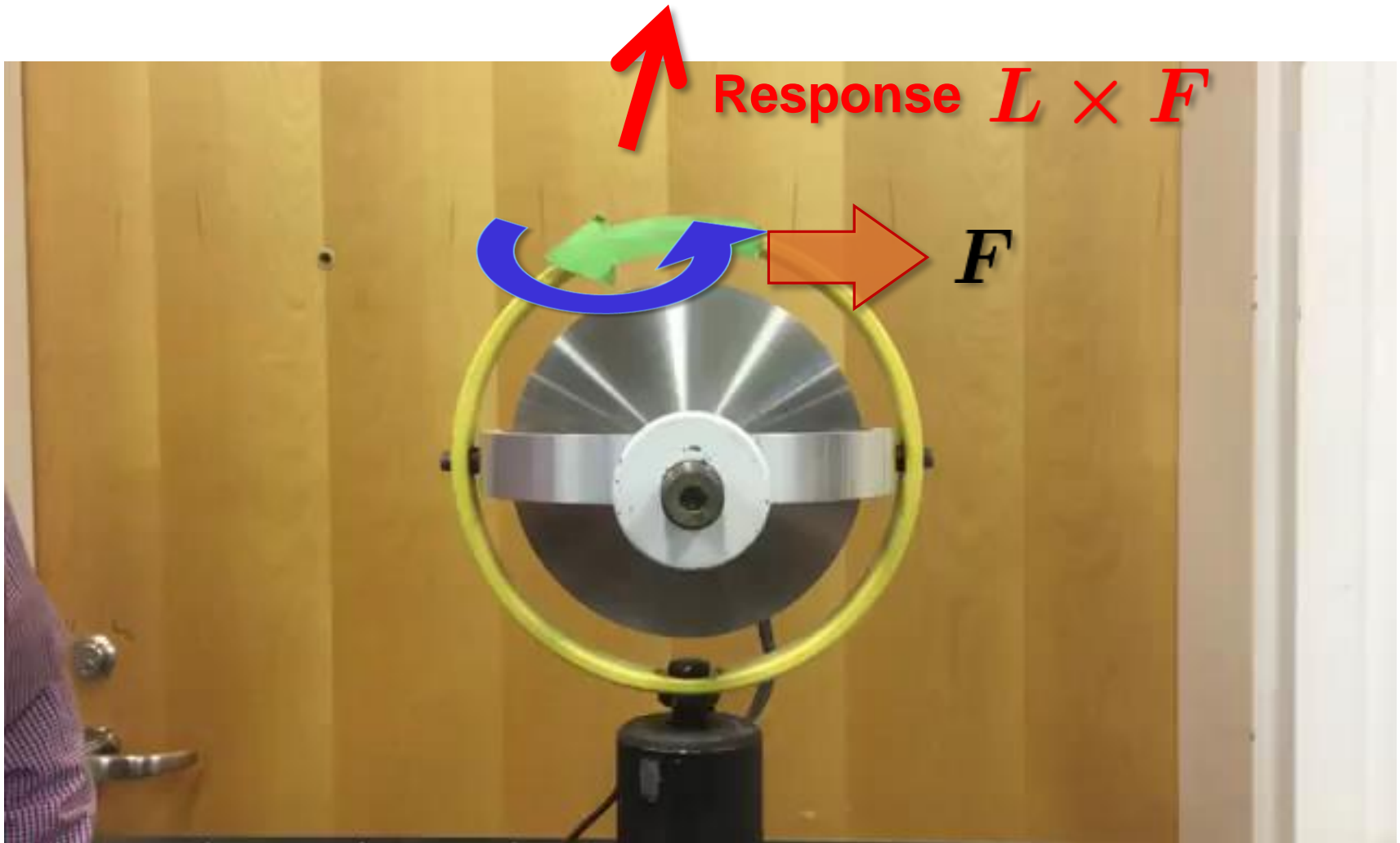
“Nuclear pinning”



“Interstitial pinning”

What we investigated - Vortex-nucleus dynamics

Response of a spinning gyroscope when pushed



We performed 3D, dynamical simulations by TDDFT with superfluidity

□ TDSLDA equations (or TDHFB, TD-BdG)

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix} = \begin{pmatrix} h(\mathbf{r}) & \Delta(\mathbf{r}) \\ \Delta^*(\mathbf{r}) & -h(\mathbf{r}) \end{pmatrix} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix}$$

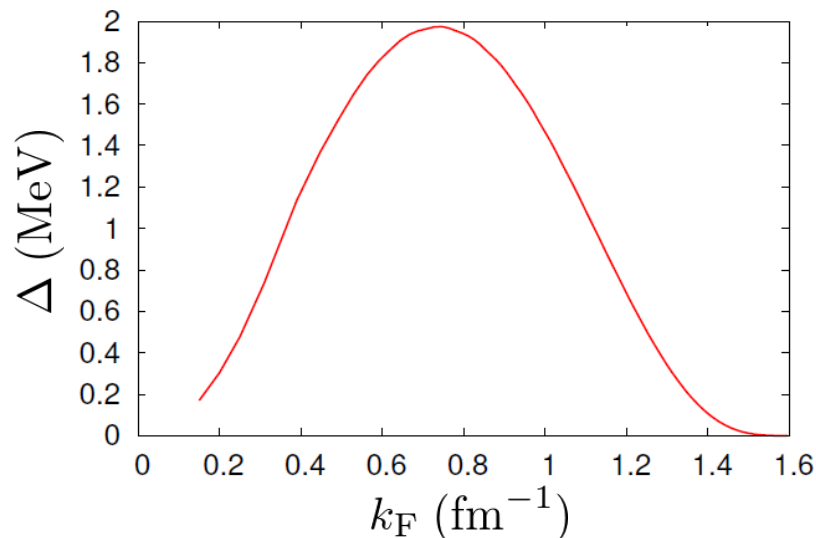
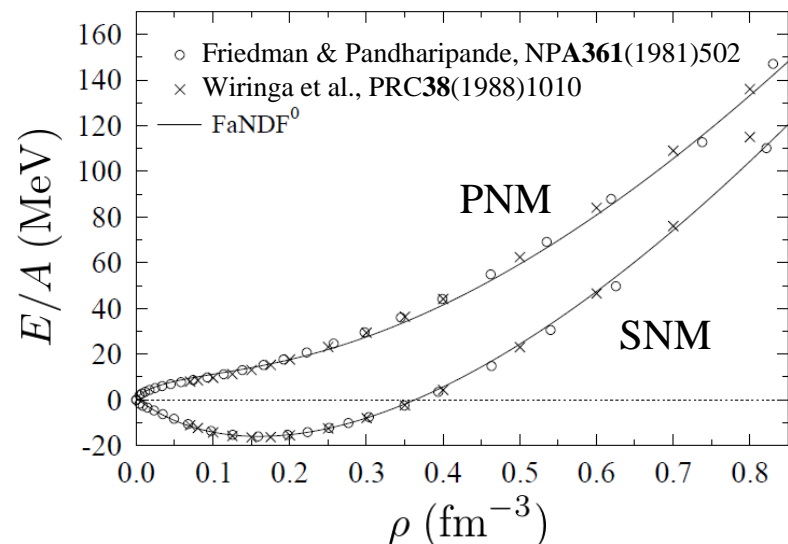
□ Energy density functional (EDF)

$$\mathcal{E}(\mathbf{r}) = \mathcal{E}_0(\mathbf{r}) + \mathcal{E}_{\text{pair}}(\mathbf{r})$$

$\mathcal{E}_0(\mathbf{r})$: Fayans EDF (FaNDF⁰) w/o LS

$$\mathcal{E}(\mathbf{r}) = \sum_{q=n,p} g[\rho_q(\mathbf{r})] |\nu_q(\mathbf{r})|^2$$

S.A. Fayans, JETP Lett. **68**, 169 (1998)



We performed 3D, dynamical simulations by TDDFT with superfluidity

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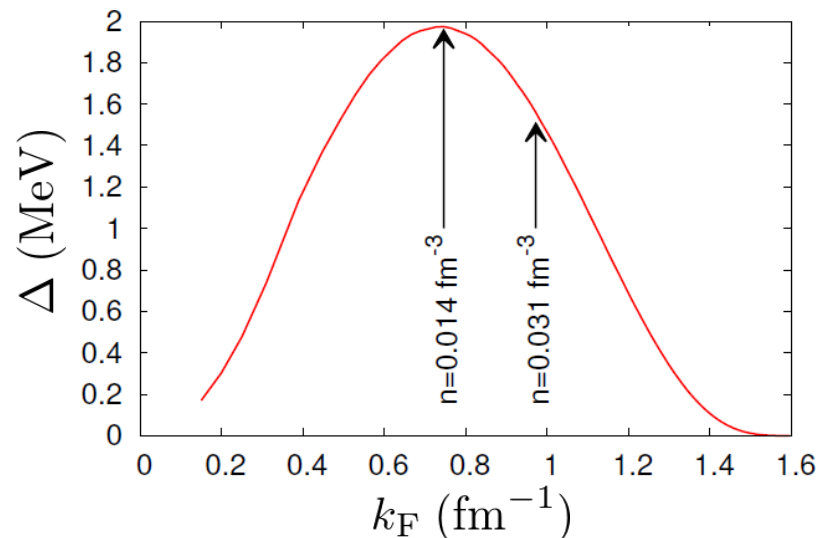
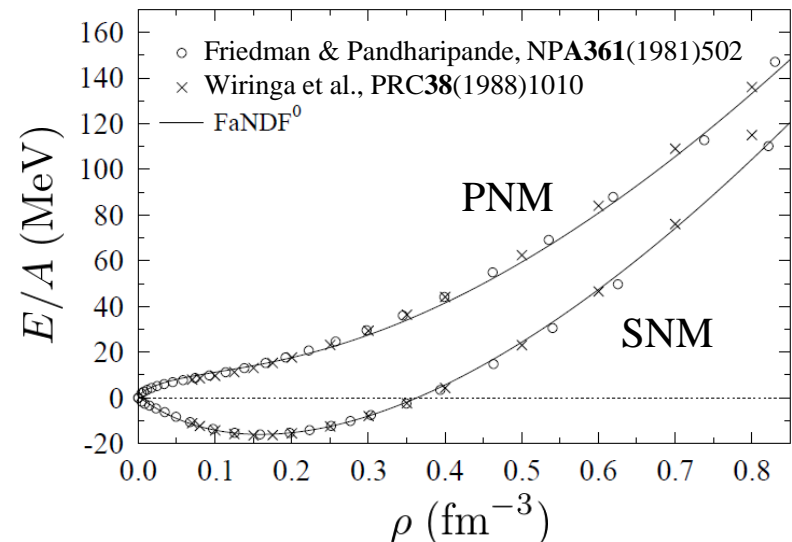
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▣ Computational details

75 fm × 75 fm × 60 fm

(50 × 50 × 40, Δx = 1.5 fm)

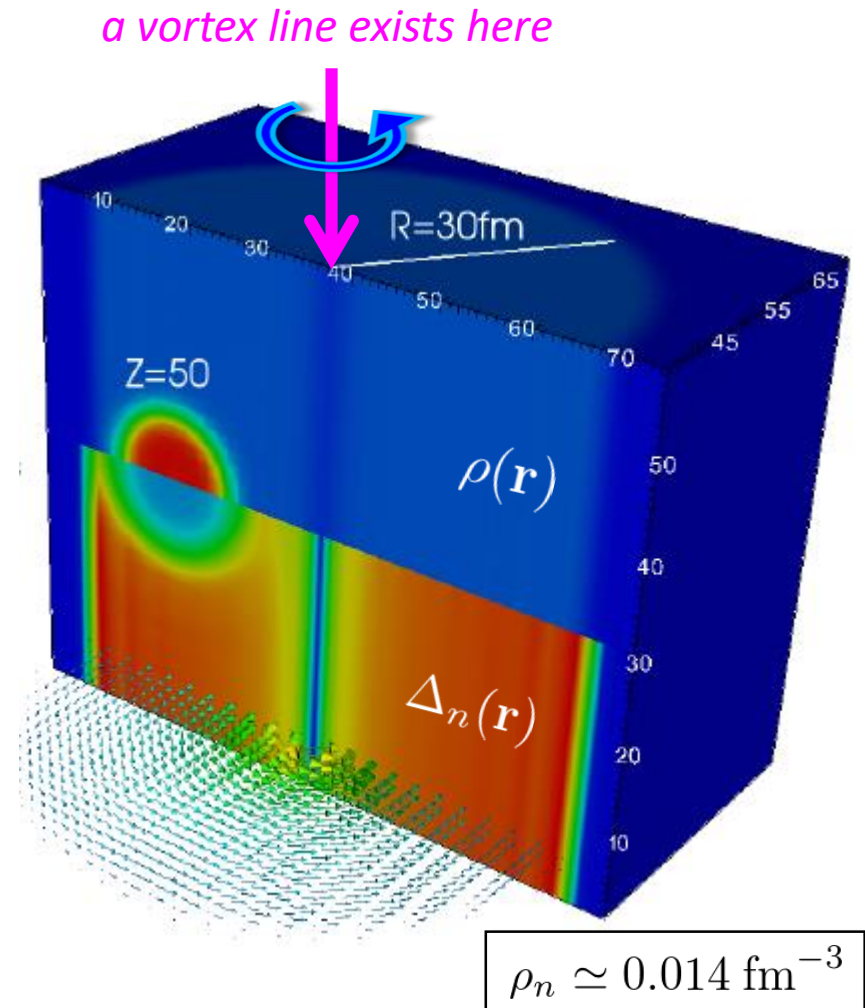
$k_c = \pi/\Delta x > k_F$ $k_F = (3\pi^2\rho_n)^{1/3}$

Nuclear impurity: Z = 50

$\rho_n \simeq 0.014 \text{ fm}^{-3}$ (N ≈ 2,530)

$\rho_n \simeq 0.031 \text{ fm}^{-3}$ (N ≈ 5,714)

of quasi-particle w.f. ≈ 100,000



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of quasi-particle w.f. $\approx 100,000$



TITAN, Oak Ridge



NERSC Edison, Berkeley

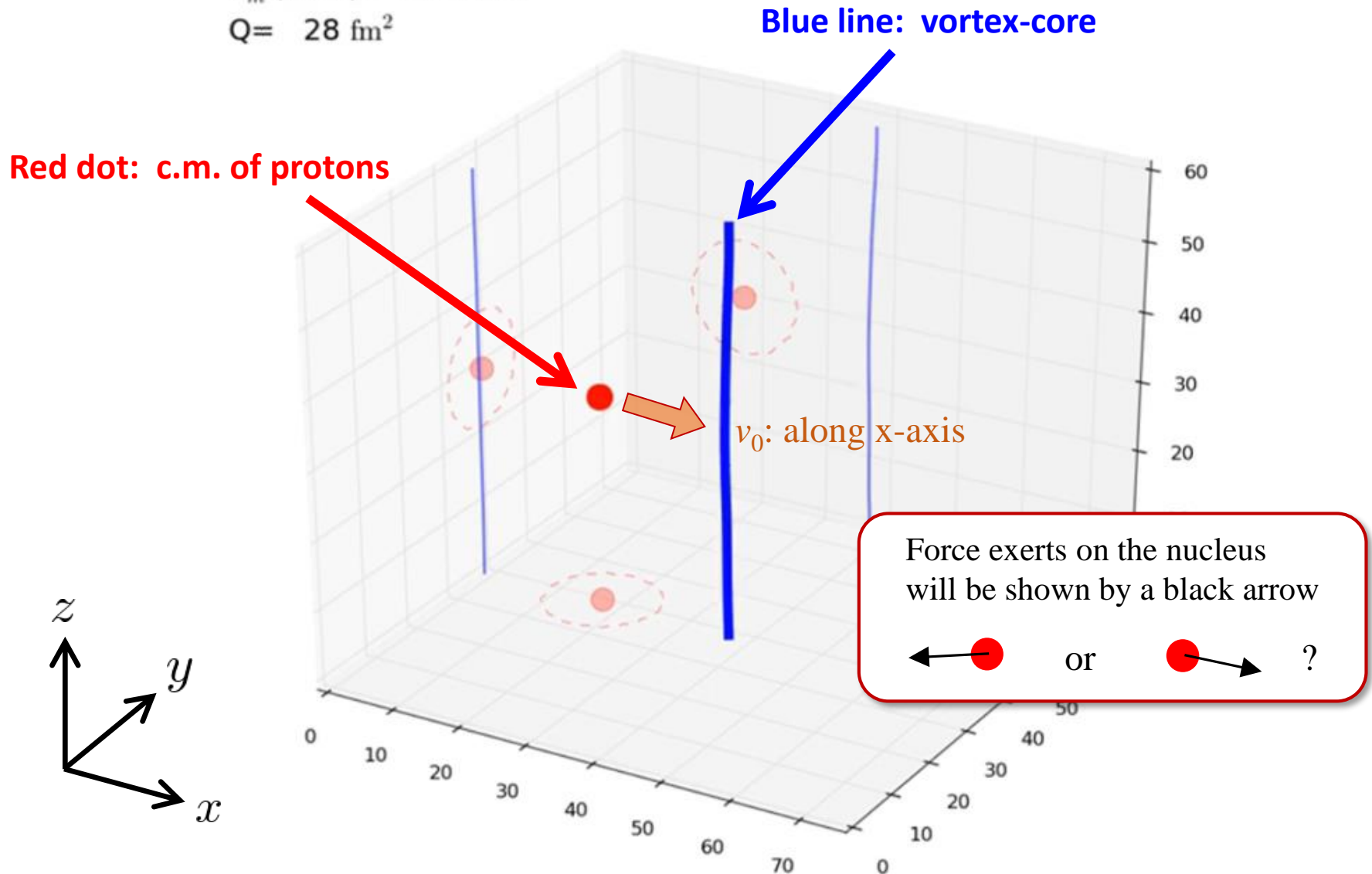


HA-PACS, Tsukuba

MPI+GPU
→ 48h w/ 200GPUs
for 10,000 fm/c

Results of TDSLDA calculation: $\rho_n \simeq 0.014 \text{ fm}^{-3}$

time= 0 fm/c
 $F_m(19.1)$ = unknown
 $Q = 28 \text{ fm}^2$



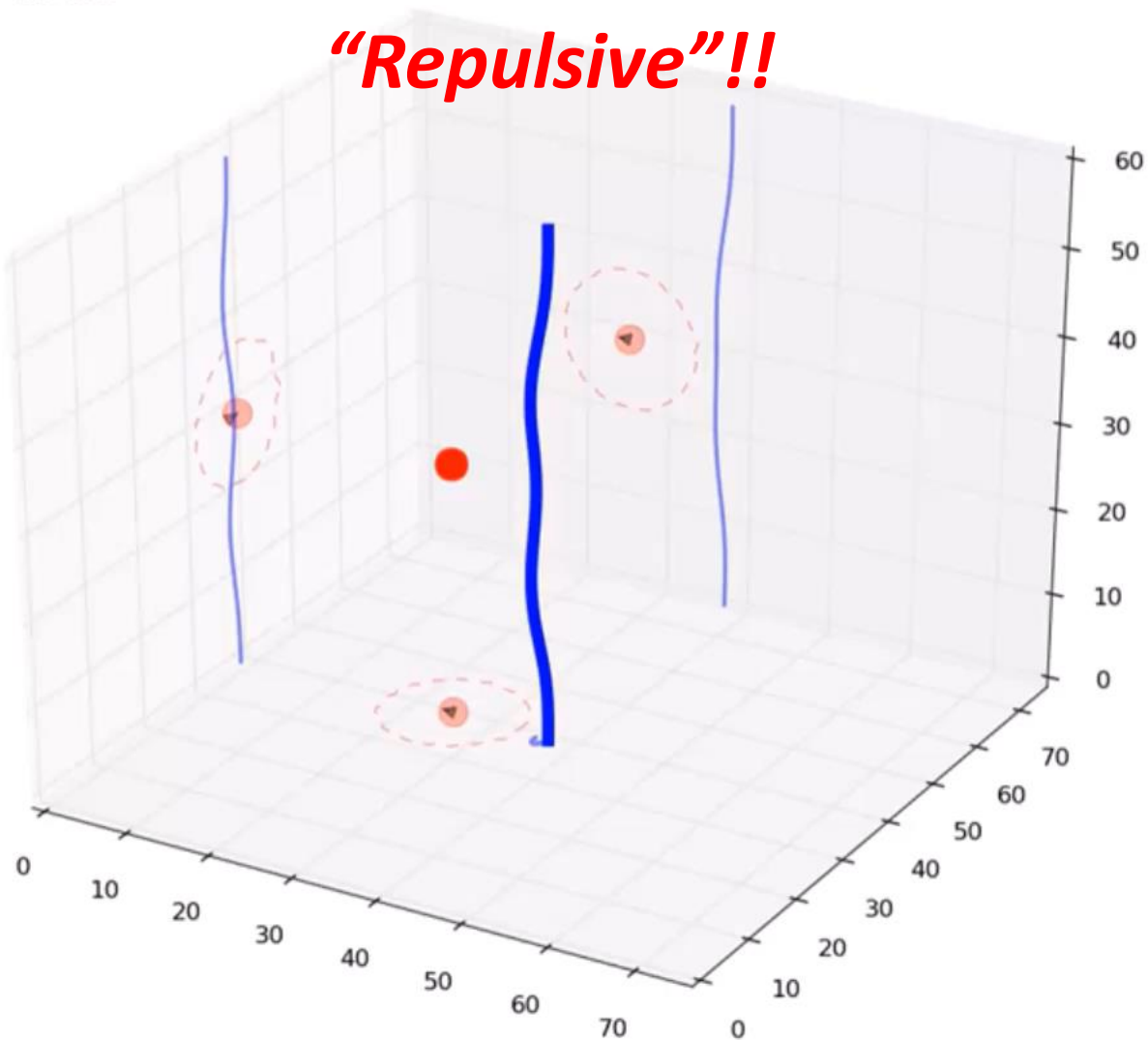
Results of TDSLDA calculation: $\rho_n \simeq 0.014 \text{ fm}^{-3}$

time= 8032 fm/c

$F_m(10.6) = 0.17 \text{ MeV/fm}$

$Q = 13 \text{ fm}^2$

“Repulsive”!!

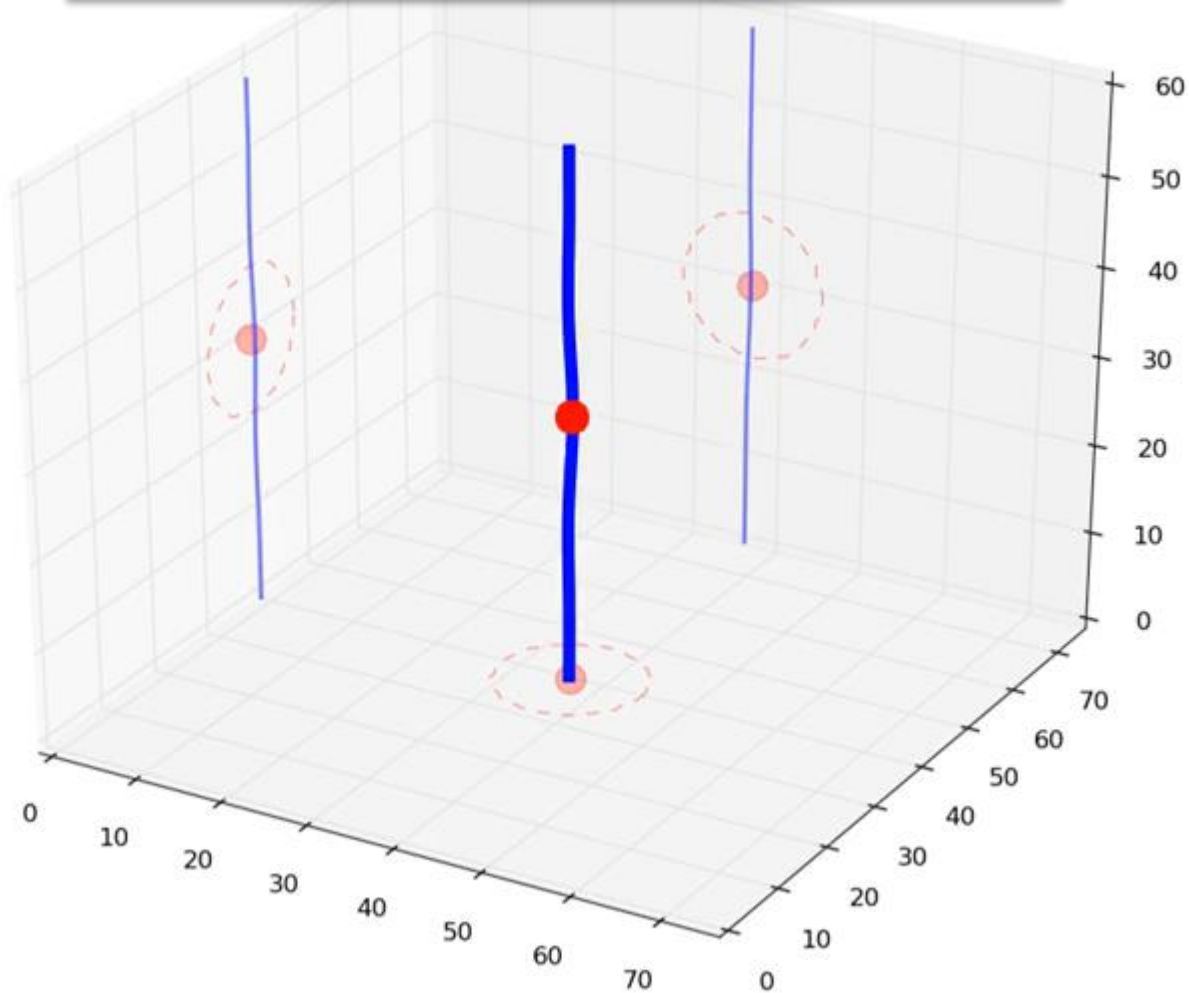


Results of TDSLDA calculation: $\rho_n \simeq 0.014 \text{ fm}^{-3}$

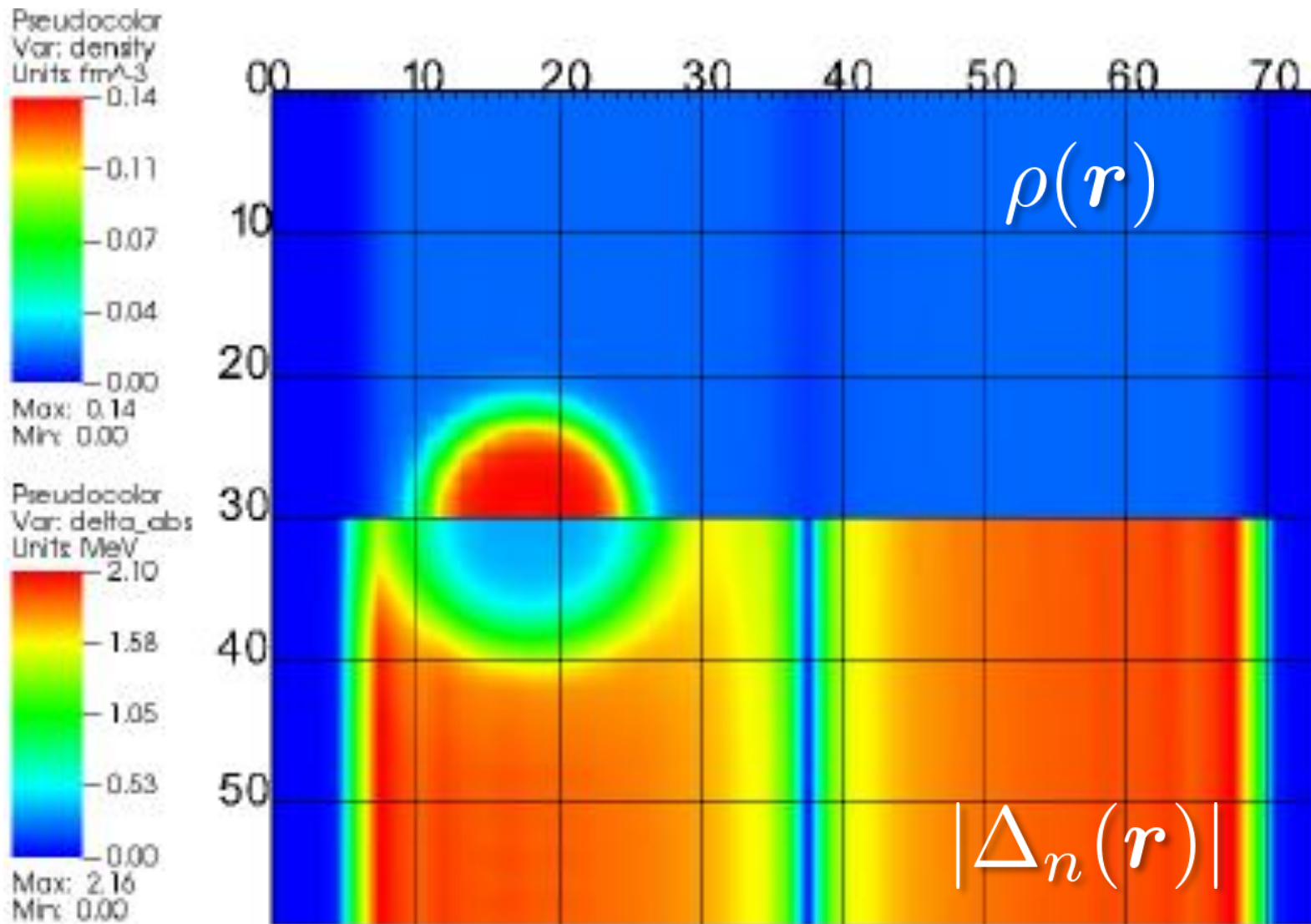
time= 0 fm/c

Q= -11 fm²

Pinned configuration is dynamically unstable

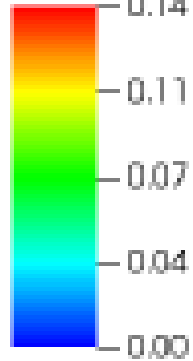


“Unpinned configuration”



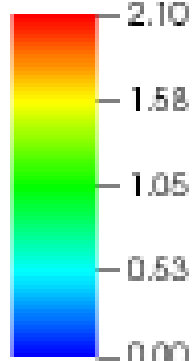
“Pinned configuration”

Pseudocolor
Var: density
Units: fm⁻³

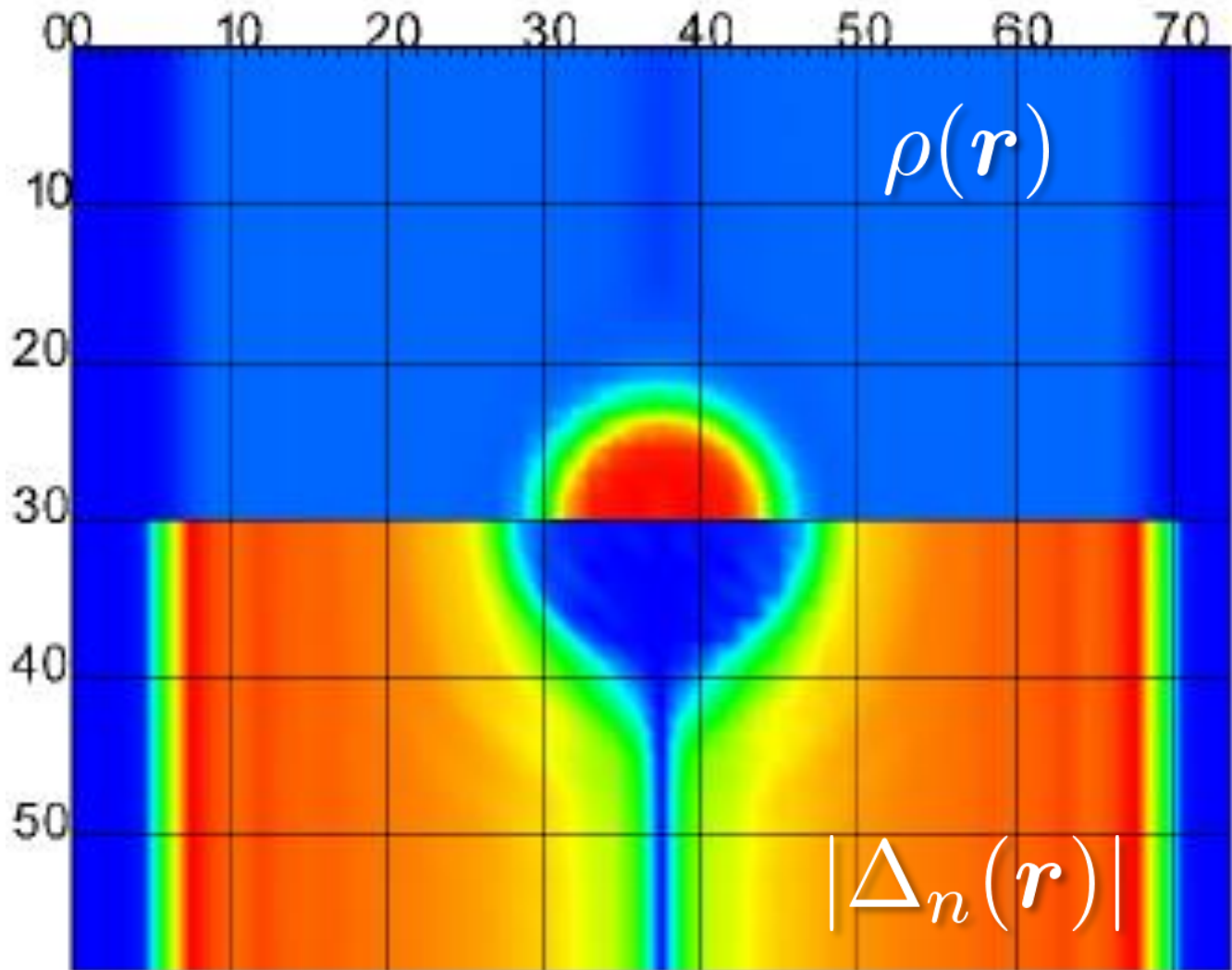


Max: 0.14
Min: 0.00

Pseudocolor
Var: delta_at
Units: MeV



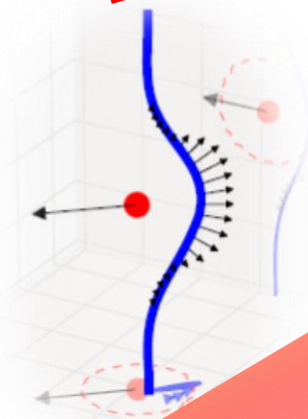
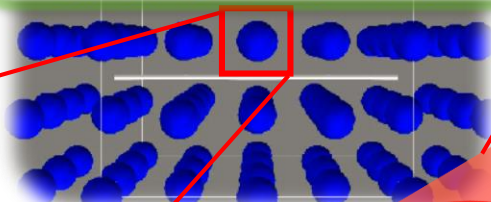
Max: 2.16
Min: 0.00



Our goal and strategy

Goal: Unveil the mechanism of glitches

New collaboration started:
Nicolaus Copernicus Astronomical Centre
B. Haskell et al.



10^{-15} - 10^{-13} m

$\sim 10^{-10}$ m

Mesoscopic

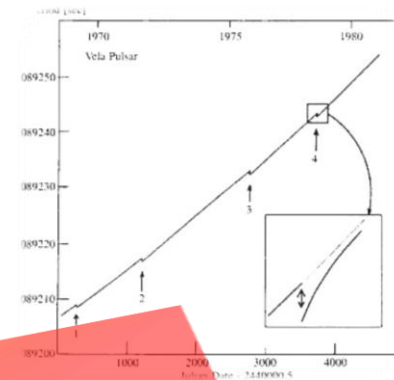
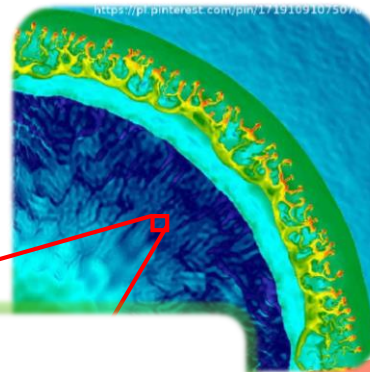
- dynamics of *vortices* in a lattice of *nuclei* (e.g. filament model)

Provide model ingredients

Microscopic

Nuclear Physics!!

- vortex-nucleus dynamics from *neutrons and protons*



10^4 m

Macroscopic

- observations
- hydrodynamics

Time-
Dependent
Band Theory
for the
Inner Crust of
Neutron Stars



The rest of the talk is based on one of my most recent publications:

PHYSICAL REVIEW C **105**, 045807 (2022)

**Time-dependent extension of the self-consistent band theory for neutron star matter:
Anti-entrainment effects in the slab phase**

Kazuyuki Sekizawa ^{1,2,*} Sorataka Kobayashi,³ and Masayuki Matsuo ^{4,†}

¹Center for Transdisciplinary Research, Institute for Research Promotion, Niigata University, Niigata 950-2181, Japan

²Nuclear Physics Division, Center for Computational Sciences, University of Tsukuba, Ibaraki 305-8577, Japan

³Graduate School of Science and Technology, Niigata University, Niigata 950-2181, Japan

⁴Department of Physics, Faculty of Science, Niigata University, Niigata 950-2181, Japan



(Received 28 December 2021; accepted 4 April 2022; published 25 April 2022)

in collaboration with



Sorataka Kobayashi

(Finished MSc in Mar. 2019)



Masayuki Matsuo

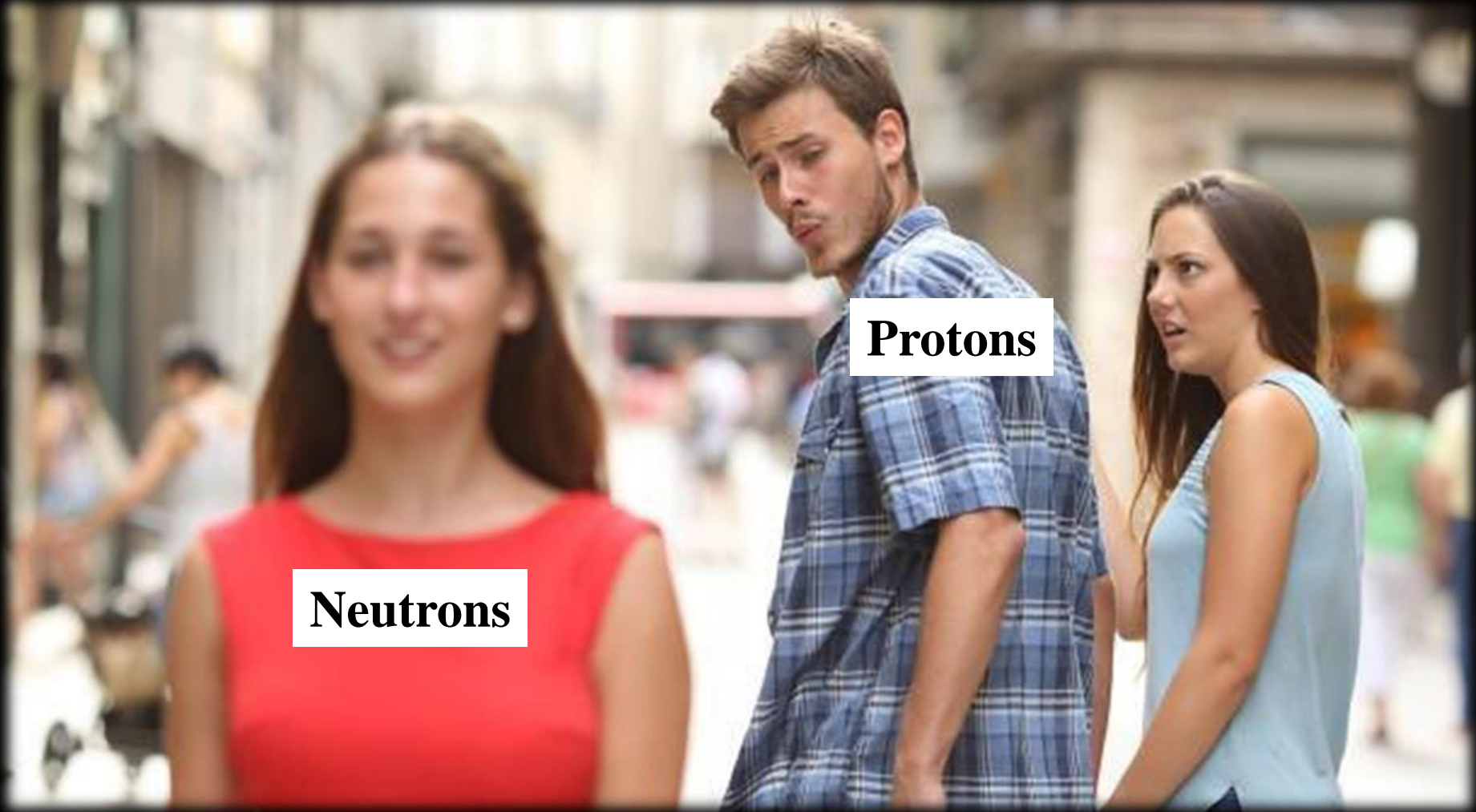


Kenta Yoshimura (M1)

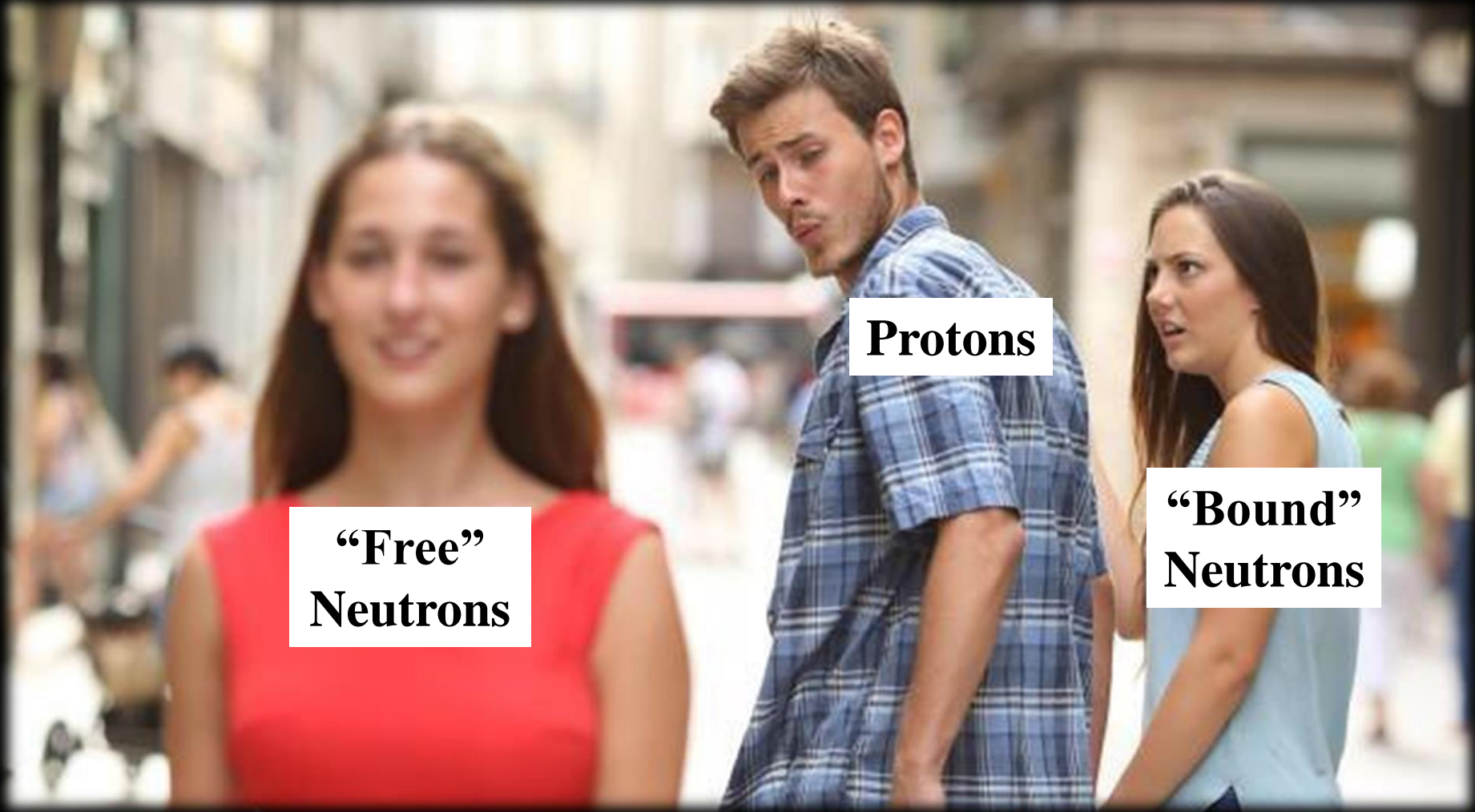


What is the “entrainment” effect?

“Entrainment” is a phenomenon between two species (particles, gases, fluids, etc.), where a motion of one component attracts the other.



“Entrainment” is a phenomenon between two species (particles, gases, fluids, etc.), where a motion of one component attracts the other.



**“Free”
Neutrons**

Protons

**“Bound”
Neutrons**

“Entrainment” in the inner crust

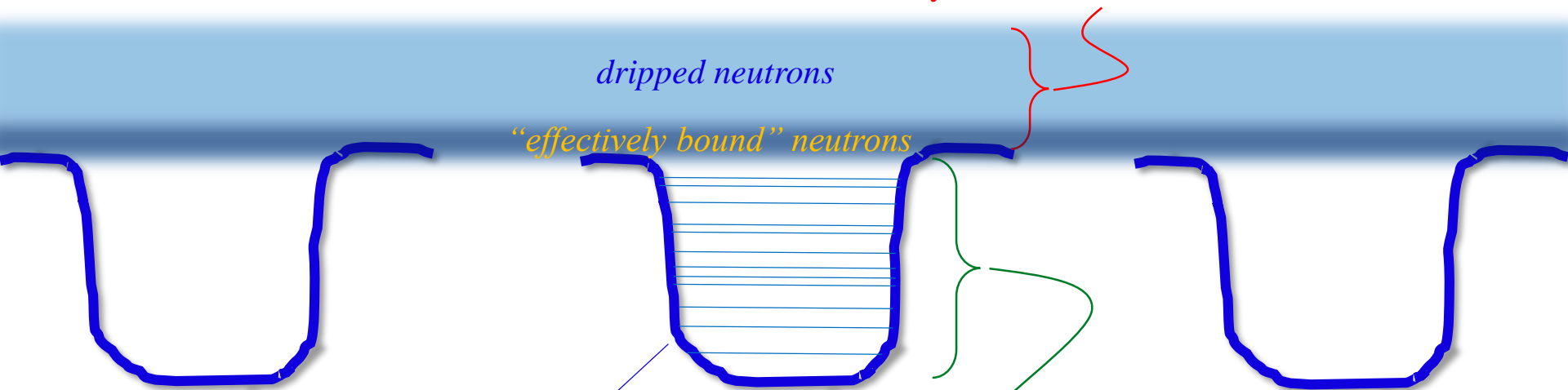
- Part of dripped neutrons are “effectively bound” (immobilized) by the periodic structure (due to Bragg scatterings), resulting in a larger effective mass

$$m_n n_n^f = m_n^* n_n^c$$

n_n^c : Conduction neutron number density
(neutrons that can actually flow)

m_n^* : (Macroscopic) Effective mass

Dripped neutrons extend spatially
→ Affected by the lattice, and a band structure is formed!



Entrainment leads:

→ reduction of n_c

→ enhancement of m^*

Potential for neutrons

Bound orbitals are well **localized**
→ Not affected by the lattice

The “entrainment effect” is still a debatable problem

- The first consideration for 1D, square-well potential

K. Oyamatsu and Y. Yamada, NPA**578**(1994)184

- Band calculations for slab (1D) and rod (2D) phases

B. Carter, N. Chamel, and P. Haensel, NPA**748**(2005)675

➔ Entrainment effects are **weak** for the slab & rod phases:

$$\frac{m^*}{m} \sim \begin{cases} 1.02 - 1.03 & \text{for the slab phase} \\ 1.11 - 1.40 & \text{for the rod phase} \end{cases}$$

- Band calculations for cubic-lattice (3D) phases

N. Chamel, NPA**747**(2005)109 (2005); NPA**773**(2006)263; PRC**85**(2012)035801; J. Low Temp. Phys. **189**, 328 (2017)

➔ **Significant** entrainment effects were found in a low-density region:

$$\frac{m^*}{m} \gtrsim 10 \text{ or more! for the cubic lattice}$$

- The first *self-consistent* band calculation for the slab (1D) phase (based on DFT with a BCPM EDF)

➔ “**Reduction**” of the effective mass was observed for the slab phase:

$$\frac{m^*}{m} \sim 0.65 - 0.75 \text{ for the slab phase}$$

Yu Kashiwaba and T. Nakatsukasa, PRC**100**(2019)035804

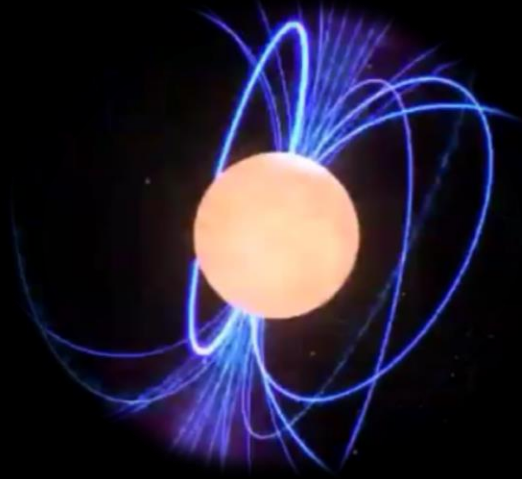
- **Time-dependent extension of the self-consistent band theory (based on TDDFT with a Skyrme EDF)**

➔ “**Reduction**” was observed, consistent with the Tsukuba group.

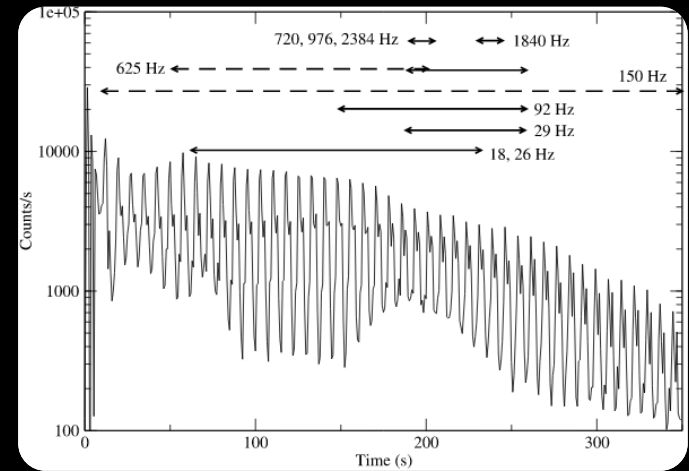
K. Sekizawa, S. Kobayashi, and M. Matsuo, PRC**105**(2022)045807

It may affect interpretation of various phenomena, e.g.:

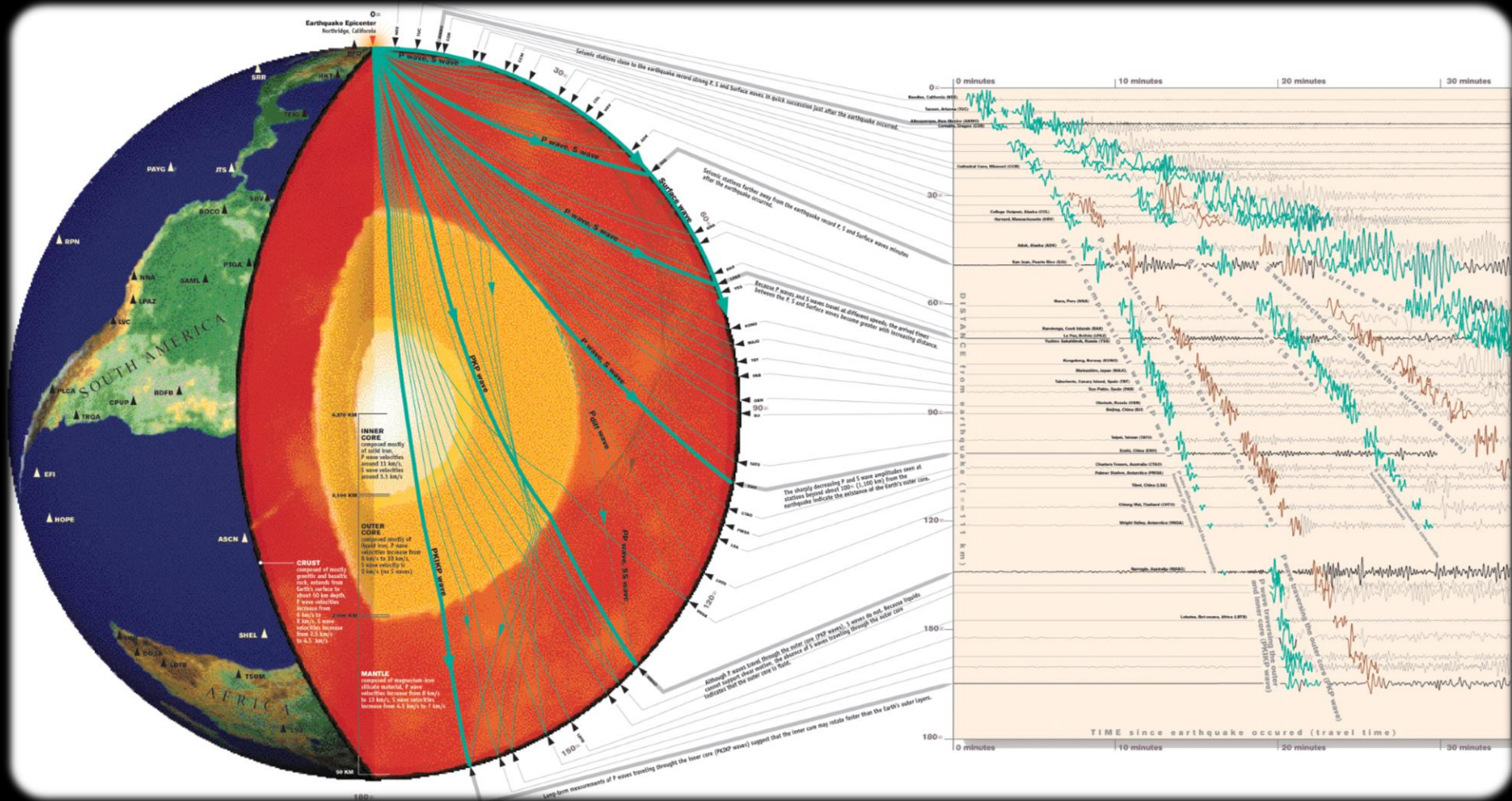
Neutron-star glitch



Quasi-periodic oscillation



Seismology (地震学): Studying inside of the Earth from earthquakes and their propagation





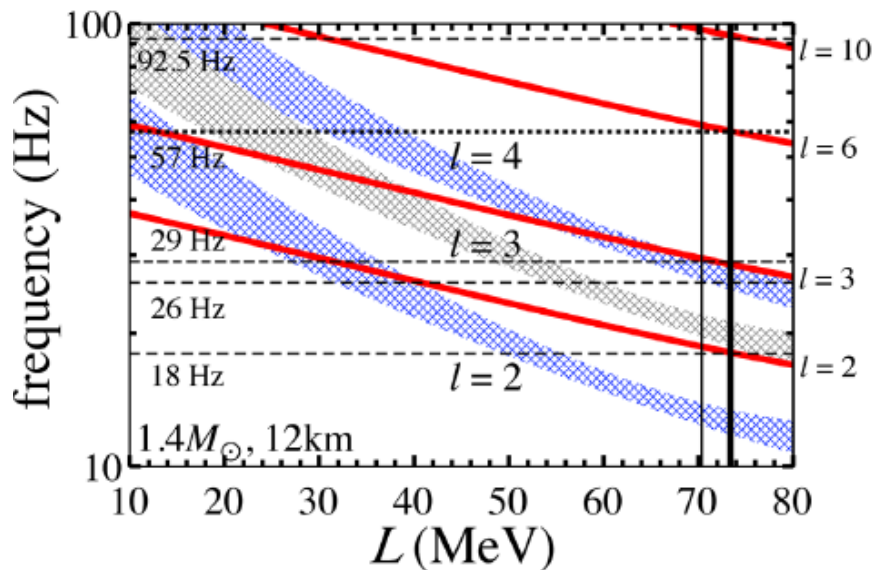
Astrophysical implications of double-layer torsional oscillations in a neutron star crust as a lasagna sandwich

Hajime Sotani¹,¹★ Kei Iida² and Kazuhiro Oyamatsu³

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²Department of Mathematics and Physics, Kochi University, 2-5-1 Akebono-cho, Kochi 780-8520, Japan

³Department of Human Informatics, Aichi Shukutoku University, 2-9 Katahira, Nagakute, Aichi 480-1197, Japan



- Many (~30) observed QPO frequencies, and prediction by a Bayesian analysis, have been nicely explained by torsional oscillations of tube–bubble or sphere–cylinder layer

Astrophysical implications of double-layer torsional oscillations in a neutron star crust as a lasagna sandwich

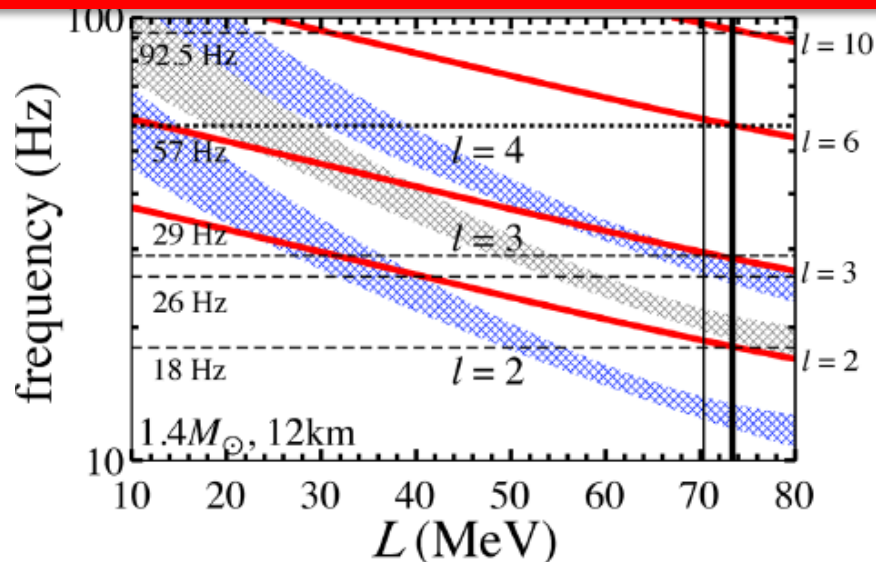
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The interpretation could be affected by the entrainment effects!



- Many (~30) observed QPO frequencies, and prediction by a Bayesian analysis, have been nicely explained by torsional oscillations of tube–bubble or sphere–cylinder layer

Recently we have developed:

Time-Dependent Band Theory based on TDDFT



Structure and dynamics of infinite neutron-star matter
can be described **microscopically** taking full account of
periodicity of crystalline structure
(*i.e.* band structure effects)

As the first step, it has been applied to the Slab phase!

We employ the Skyrme-Kohn-Sham DFT with the Bloch boundary condition

✓ The Bloch boundary condition for single-particle orbitals

$$\psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r}) = \frac{1}{\sqrt{V}} u_{\alpha\mathbf{k}}^{(q)}(z) e^{i\mathbf{k}\cdot\mathbf{r}} \quad u_{\alpha\mathbf{k}}^{(q)}(z + na) = u_{\alpha\mathbf{k}}^{(q)}(z)$$

Periodicity of the slabs

α : Band index \mathbf{k} : Bloch wave vector q : Isospin (n or p) a : Period of the slabs

✓ Skyrme EDF

$$\frac{E}{A} = \frac{1}{N_b} \int_0^a \left(\frac{\hbar^2}{2m} \tau(z) + \sum_{t=0,1} \left[C_t^p [n] n_t^2(z) + C_t^{\Delta\rho} n_t(z) \partial_z^2 n_t(z) + C_t^T (n_t(z) \tau_t(z) - \mathbf{j}_t^2(z)) \right] + \mathcal{E}_{\text{Coul}}^{(p)}(z) \right) dz$$

Number density:

$$n_q(z) = 2 \sum_{\alpha,\mathbf{k}}^{\text{occ.}} |\psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r})|^2$$

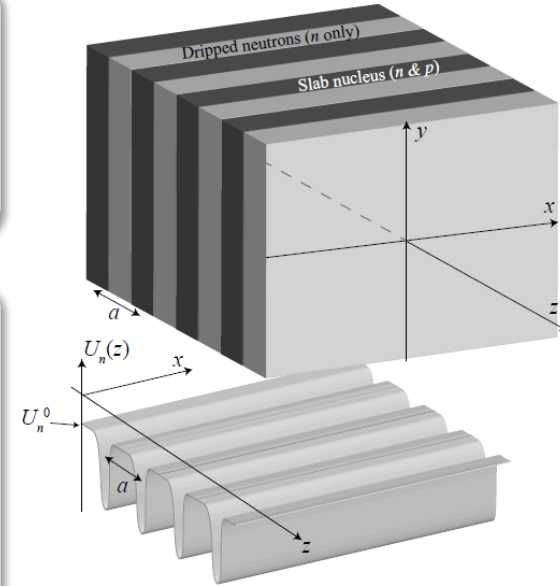
Kinetic density:

$$\tau_q(z) = 2 \sum_{\alpha,\mathbf{k}}^{\text{occ.}} |\nabla \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r})|^2$$

Current (momentum) density:

$$\mathbf{j}_q(z) = 2 \sum_{\alpha,\mathbf{k}}^{\text{occ.}} \text{Im} [\psi_{\alpha\mathbf{k}}^{(q)*}(\mathbf{r}) \nabla \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r})]$$

*Uniform background electrons are assumed for the charge neutrality condition: $n_e = \bar{n}_p$



Picture from PRC100(2019)035804

✓ Skyrme-Kohn-Sham equations

$$\hat{h}^{(q)}(z) \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r}) = \varepsilon_{\alpha\mathbf{k}}^{(q)} \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r}) \quad \rightarrow \quad \left(\hat{h}^{(q)}(z) + \hat{h}_{\mathbf{k}}^{(q)}(z) \right) u_{\alpha\mathbf{k}}^{(q)}(z) = \varepsilon_{\alpha\mathbf{k}}^{(q)} u_{\alpha\mathbf{k}}^{(q)}(z)$$

Ordinary single-particle Hamiltonian:

$$\hat{h}^{(q)}(z) = -\nabla \cdot \frac{\hbar^2}{2m_q^\oplus(z)} \nabla + U^{(q)}(z) + \frac{1}{2i} [\nabla \cdot \mathbf{I}^{(q)}(z) + \mathbf{I}^{(q)}(z) \cdot \nabla]$$

Additional (k -dependent) term:

$$\hat{h}_{\mathbf{k}}^{(q)}(z) = \frac{\hbar^2 \mathbf{k}^2}{2m_q^\oplus(z)} + \hbar \mathbf{k} \cdot \hat{\mathbf{v}}^{(q)}(z)$$

Velocity operator:

$$\hat{\mathbf{v}}^{(q)}(z) \equiv \frac{1}{i\hbar} [\mathbf{r}, \hat{h}^{(q)}(z)]$$

Note: While we deal with 3D slabs, the equations to be solved are 1D!

Results: Band structure ($Y_p = 0.1$)

Proton fraction:

$$Y_p = \frac{\bar{n}_p}{\bar{n}_n + \bar{n}_p}$$

Average nucleon density:

$$\bar{n}_q = \frac{1}{a} \int_0^a n_q(z) dz$$

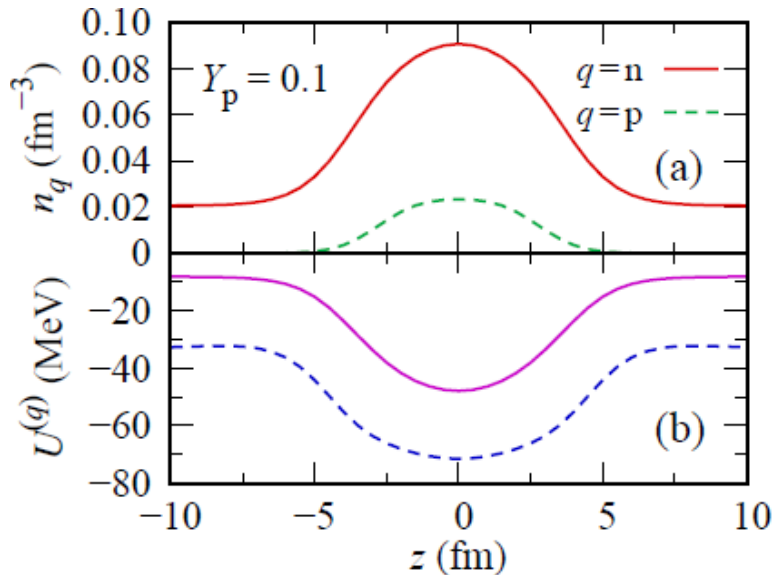
Single-particle energy:

$$\varepsilon_{\alpha\mathbf{k}}^{(q)} = \underbrace{e_{\alpha\mathbf{k}}^{(q)}}_{z\text{-component}} + \underbrace{\varepsilon_{\text{kin-xy},\alpha\mathbf{k}}^{(q)}}_{\approx \frac{\hbar^2 k_{\parallel}^2}{2m}} \quad k_{\parallel} = \sqrt{k_x^2 + k_y^2}$$

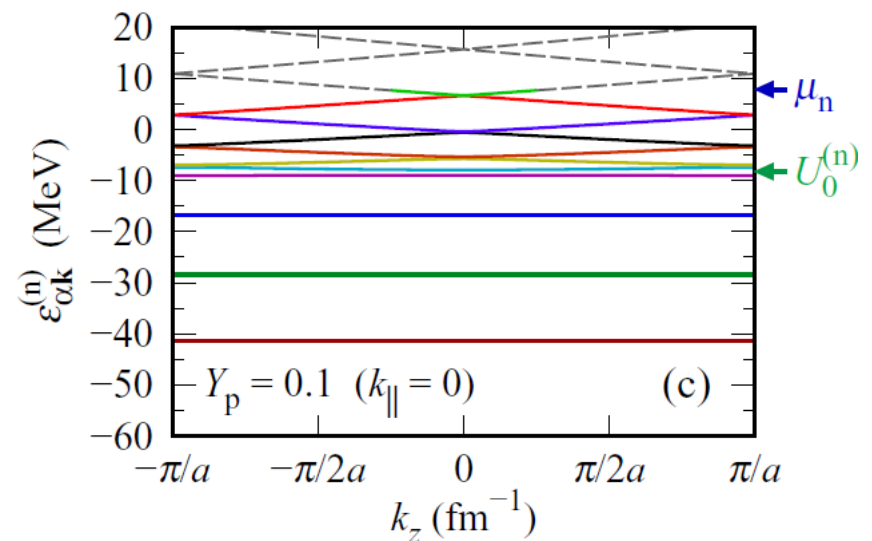
✓ Dripped neutrons show band structure (k_z dependence)

$Y_p = 0.1, n_B = 0.4 \text{ fm}^{-3}$: Neutron-dripped slab

Density and potential



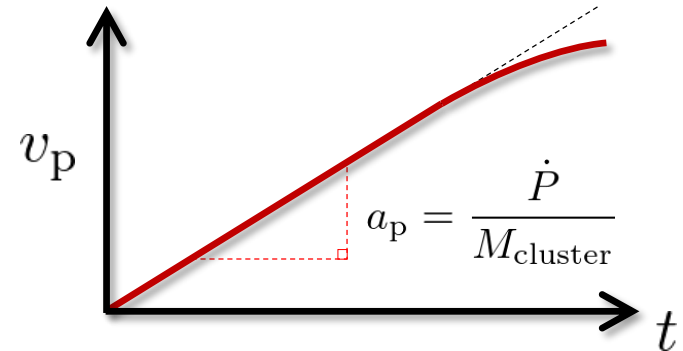
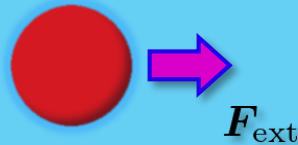
Neutron single-particle energies



- ✓ The collective mass is extracted from **acceleration motion under constant force**

The real-time method: Idea

Dripped neutrons



How to introduce spatially-uniform electric field

- ✓ TDKS equation in a “velocity gauge”

$$i\hbar \frac{\partial \tilde{u}_{\alpha\mathbf{k}}^{(q)}(z, t)}{\partial t} = \left(\hat{h}^{(q)}(z, t) + \hat{h}_{\mathbf{k}(t)}^{(q)}(z, t) \right) \tilde{u}_{\alpha\mathbf{k}}^{(q)}(z, t)$$

Spatially-uniform
Vector potential

$$\mathbf{k}(t) = \mathbf{k} + \frac{e}{\hbar c} A_z(t) \hat{\mathbf{e}}_z$$

Gauge transformation for the Bloch orbitals:

$$\tilde{u}_{\alpha\mathbf{k}}^{(q)}(z, t) = \exp\left[-\frac{ie}{\hbar c} A_z(t) z\right] u_{\alpha\mathbf{k}}^{(q)}(z, t)$$

Electric field:

$$E_z(t) = -\frac{1}{c} \frac{dA_z}{dt}$$

k -dependent term:

$$\hat{h}_{\mathbf{k}}^{(q)}(z) = \frac{\hbar^2 \mathbf{k}^2}{2m_q^\oplus(z)} + \hbar \mathbf{k} \cdot \hat{\mathbf{v}}^{(q)}(z)$$

Velocity operator:

$$\hat{\mathbf{v}}^{(q)}(z) \equiv \frac{1}{i\hbar} [\mathbf{r}, \hat{h}^{(q)}(z)]$$

cf. K. Yabana and G.F. Bertsch, Phys. Rev. B **54**, 4484 (1996); G.F. Bertch *et al.*, Phys. Rev. B **62**, 7998 (2000)

Results: The collective mass

Acceleration:

$$a_p = \frac{d^2 Z}{dt^2}$$

C.m. position of protons:

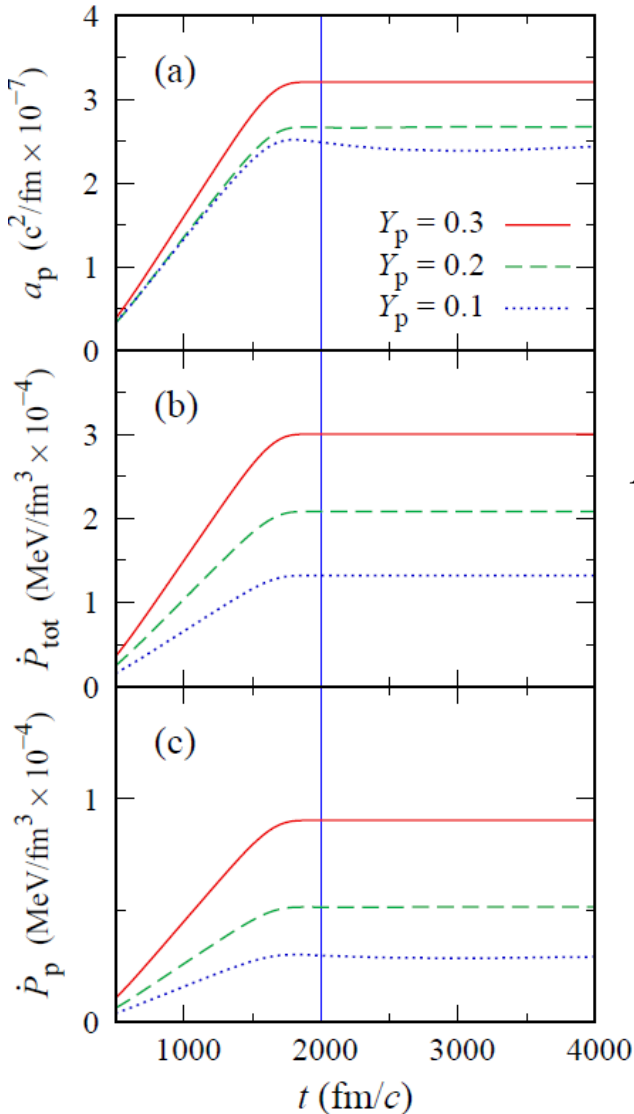
$$Z(t) = \frac{1}{a} \int_0^a z n_p(z, t) dz$$

Momentum of nucleons:

$$P_q(t) = \hbar \int_0^a j_q(z, t) dz$$

Total momentum:

$$P_{\text{tot}}(t) = P_n(t) + P_p(t)$$



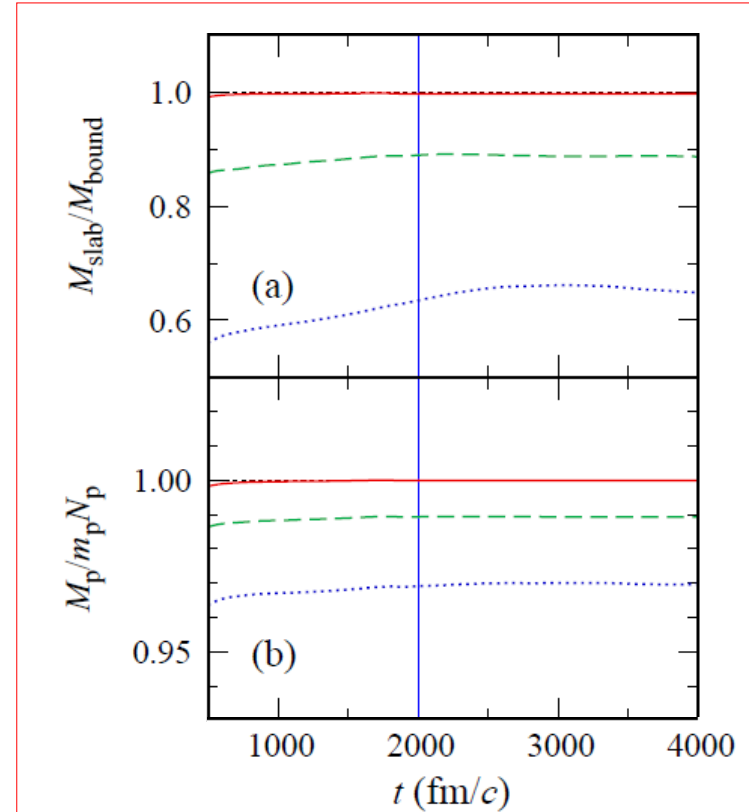
$$M_{\text{slab}} = \dot{P}_{\text{tot}}/a_p$$



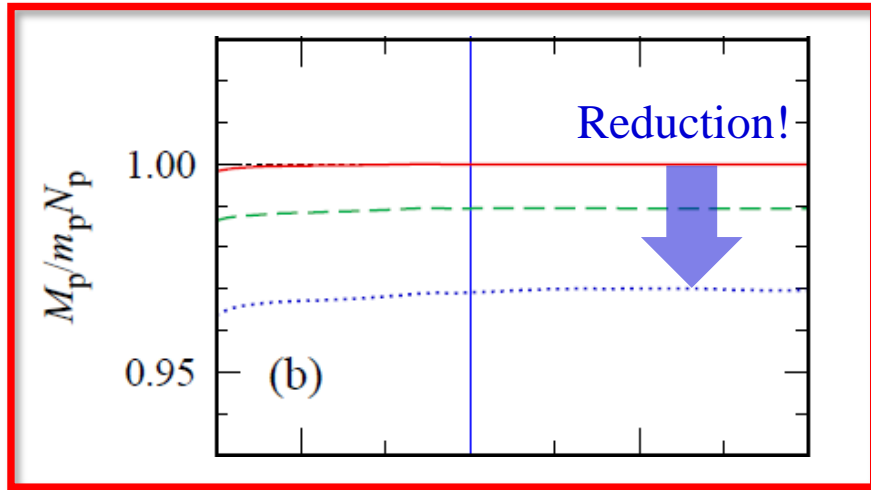
$$M_p = \dot{P}_p/a_p$$

✓ For neutron-dripped slabs, we find significant **reduction** of the collective mass!

➤ What is the origin of the reduction?



- ✓ Cause of the reduction of the collective mass of protons:
the density-dependent “microscopic” effective mass



Collective mass of protons

$$M_p \leq m_p N_p$$

$$\approx m_p^\oplus [n_n^{\text{b.g.}}] N_p$$

Protons and bound neutrons move together



There must be a velocity lag between protons and background neutrons!

The continuity equation within Skyrme TDDFT reads:

$$\frac{\partial \rho_q(\mathbf{r}, t)}{\partial t} + \hbar \nabla \cdot \mathbf{p}_q(\mathbf{r}, t) = 0$$

where

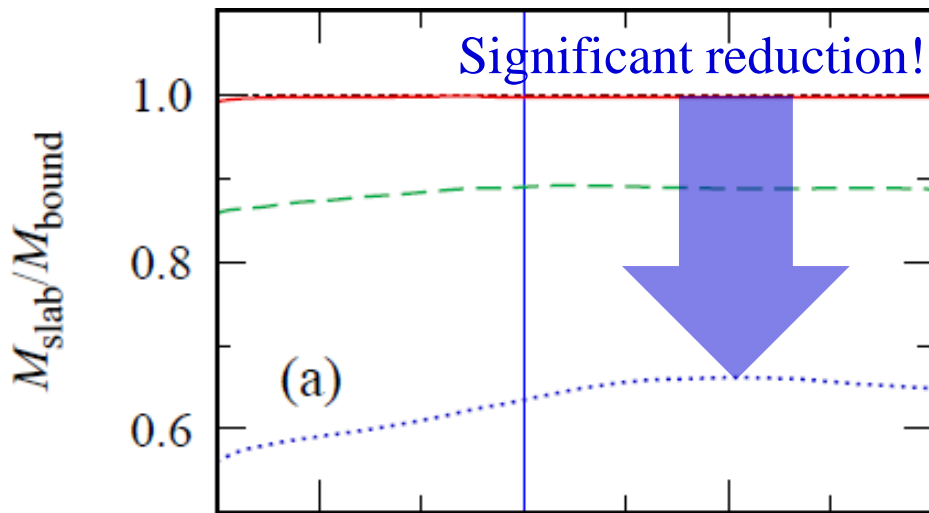
$$\mathbf{p}_q(\mathbf{r}, t) = \mathbf{j}_q(\mathbf{r}, t) + q \frac{2m_q}{\hbar^2} (C_0^\tau - C_1^\tau) n_n(\mathbf{r}, t) n_p(\mathbf{r}, t) \left(\frac{\mathbf{j}_p(\mathbf{r}, t)}{n_p(\mathbf{r}, t)} - \frac{\mathbf{j}_n(\mathbf{r}, t)}{n_n(\mathbf{r}, t)} \right)$$

+1 for protons
-1 for neutrons

velocity difference

Then, what is the cause of the reduction
of the collective mass of the slab?

→ an “anti-entrainment” effect!



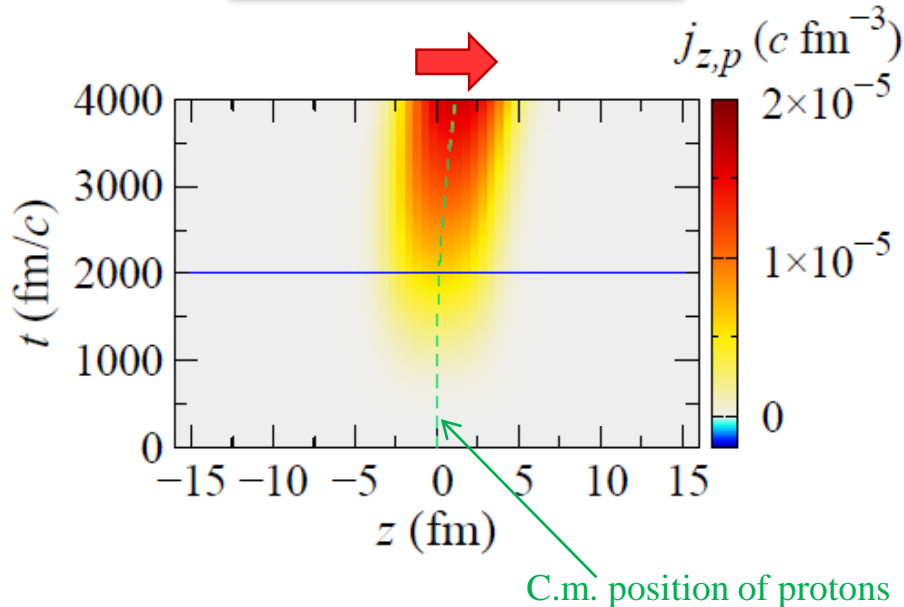
It can **not** be explained solely by
the microscopic effective mass.

Current density:

$$j_{z,q}(z,t) = \frac{\hbar}{m_q} \sum_{\alpha,\mathbf{k}}^{\text{occ.}} \text{Im}[\psi_{\alpha\mathbf{k}}^{(q)*}(\mathbf{r},t) \nabla \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r},t)] = \frac{\hbar}{m_q} \frac{1}{aN_{k_z}} \sum_{\alpha,k_z} \int \frac{k_{\parallel}}{\pi} \text{Im}[u_{\alpha\mathbf{k}}^{(q)*}(z,t)(\partial_z + ik_z)u_{\alpha\mathbf{k}}^{(q)}(z,t)] \theta(\mu_q - \varepsilon_{\alpha\mathbf{k}}^{(q)}) dk_{\parallel}$$

- ✓ Protons inside the slab move toward the direction of the external force, as expected.

Proton current density



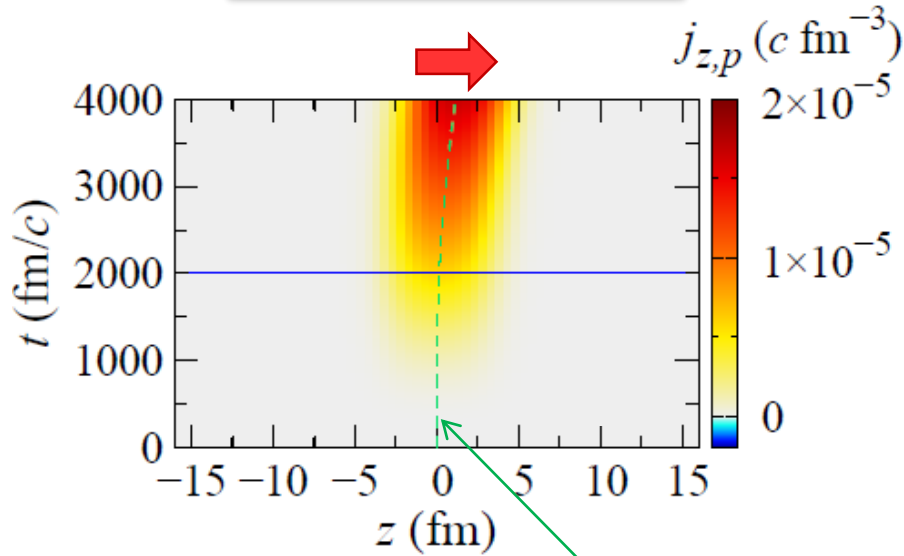
Current density:

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✓ Dripped neutrons outside the slab move toward the opposite direction!

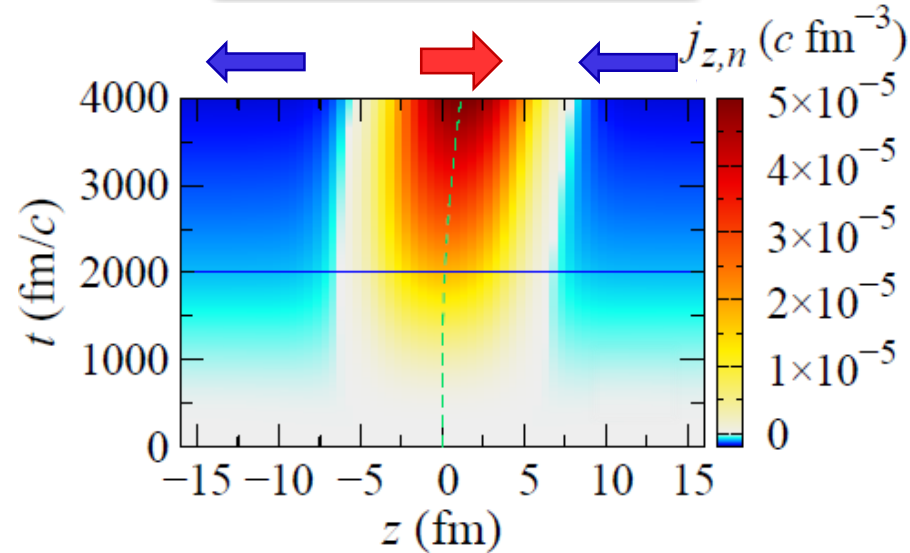
Since it reduces P_{tot} and \dot{P}_{tot} , $M_{\text{slab}} = \dot{P}_{\text{tot}}/a_p$ is reduced

Proton current density



C.m. position of protons

Neutron current density



$$(m_{n,\alpha\mathbf{k}}^{*-1})_{\mu\nu} = \frac{1}{\hbar^2} \frac{\partial^2 \varepsilon_{\alpha\mathbf{k}}^{(n)}}{\partial k_{\mu} \partial k_{\nu}}$$

Current density:

$$j_{z,q}(z,t) = \frac{\hbar}{m_q} \sum_{\alpha, \mathbf{k}}^{\text{occ.}} \text{Im}[\psi_{\alpha\mathbf{k}}^{(q)*}(\mathbf{r}, t) \nabla \psi_{\alpha\mathbf{k}}^{(q)}(\mathbf{r}, t)] = \frac{\hbar}{m_q} \frac{1}{aN_{k_z}} \sum_{\alpha, k_z} \int \frac{k_{\parallel}}{\pi} \text{Im}[u_{\alpha\mathbf{k}}^{(q)*}(z,t) (\partial_z + ik_z) u_{\alpha\mathbf{k}}^{(q)}(z,t)] \theta(\mu_q - \varepsilon_{\alpha\mathbf{k}}^{(q)}) dk_{\parallel}$$

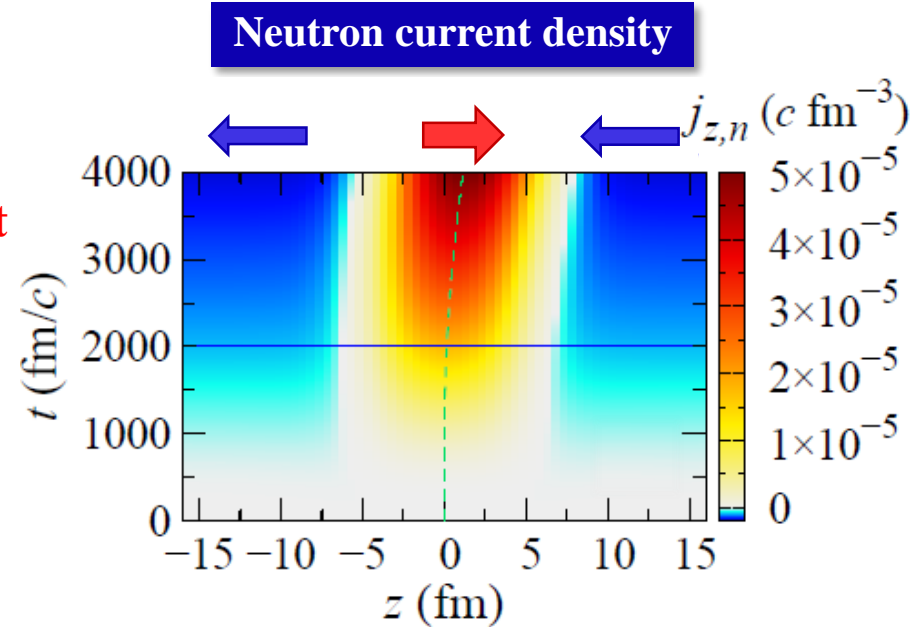
✓ Dripped neutrons outside the slab move toward the opposite direction!

Since it reduces P_{tot} and \dot{P}_{tot} , $M_{\text{slab}} = \dot{P}_{\text{tot}}/a_p$ is reduced

Reduction of M_{slab}
 → enhancement of n_c
 → reduction of m^*

We interpret it as an “anti-entrainment” effect

Y_p	n_n^f/\bar{n}_n	Static		Dynamic
		n_n^c/\bar{n}_n	m_n^*/m_n	n_n^c/\bar{n}_n
0.3	2.09×10^{-4}	0.005	0.040	0.005
0.2	0.127	0.256	0.496	0.229
0.1	0.362	0.630	0.574	0.586



$$(m_{n,\alpha\mathbf{k}}^{*-1})_{\mu\nu} = \frac{1}{\hbar^2} \frac{\partial^2 \varepsilon_{\alpha\mathbf{k}}^{(n)}}{\partial k_{\mu} \partial k_{\nu}}$$

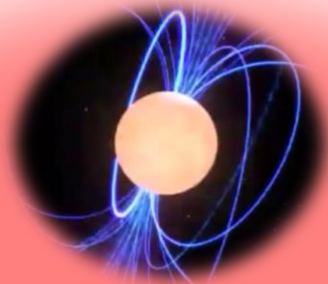


Summary

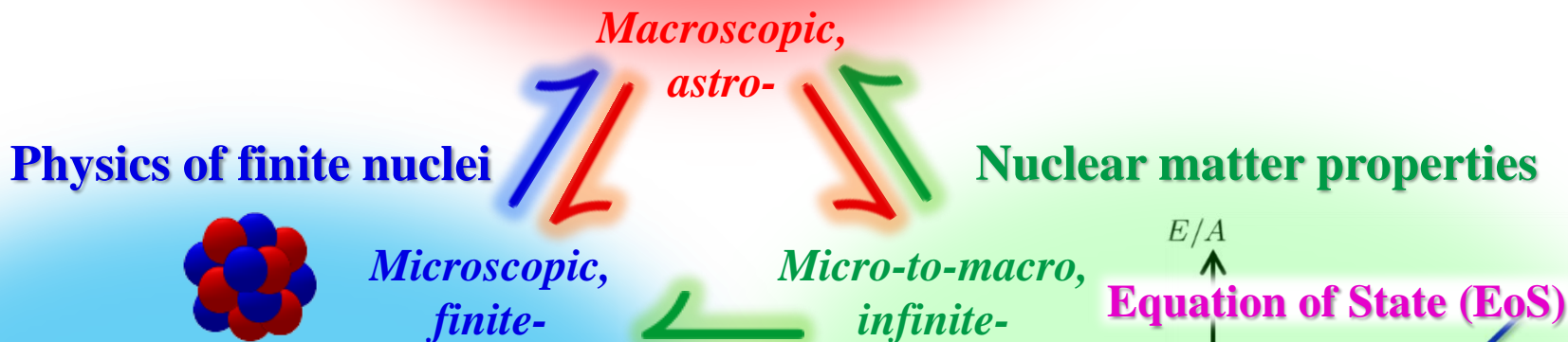
➤ Definitely, all are rooted with the wonder of nuclear physics which is basically a quantum many-body problem! ;)

Takeaway message

Physics of neutron stars

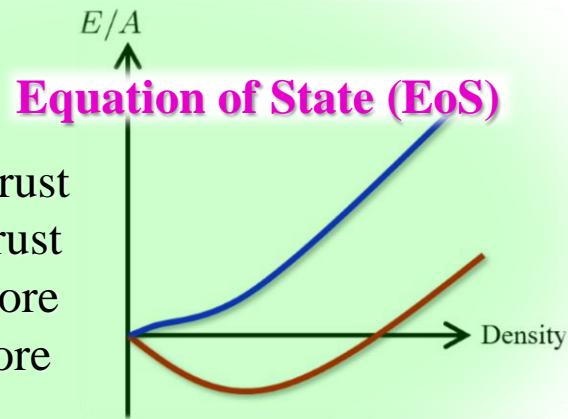


- Mass, radius, deformability, ... (\Leftrightarrow EoS, GW, ...)
- Torsional oscillations (\Leftrightarrow QPO, Pasta, GW, ...)
- Superfluidity, cooling, glitches, ...



- Quantum many-body problem
- Nuclear force (\Leftrightarrow EoS, Structure & Reactions)
- Mass/binding energy (\Leftrightarrow Crust compositions)
- Excitation properties (\Leftrightarrow GMR, GDR, EoS, ...)
- Nuclear reactions (\Leftrightarrow Stellar evolution, SNe, ...)

- Outer crust
- Inner crust
- Outer core
- Inner core



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About us: <https://nuclphystitech.wordpress.com/>

See also:

