Webinar at Department of Physics, Ankara University December 23rd, 2021

Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars

Kazuyuki Sekizawa

Department of Physics, School of Science Tokyo Institute of Technology



Brief personal history



Main research field: Nuclear Theory

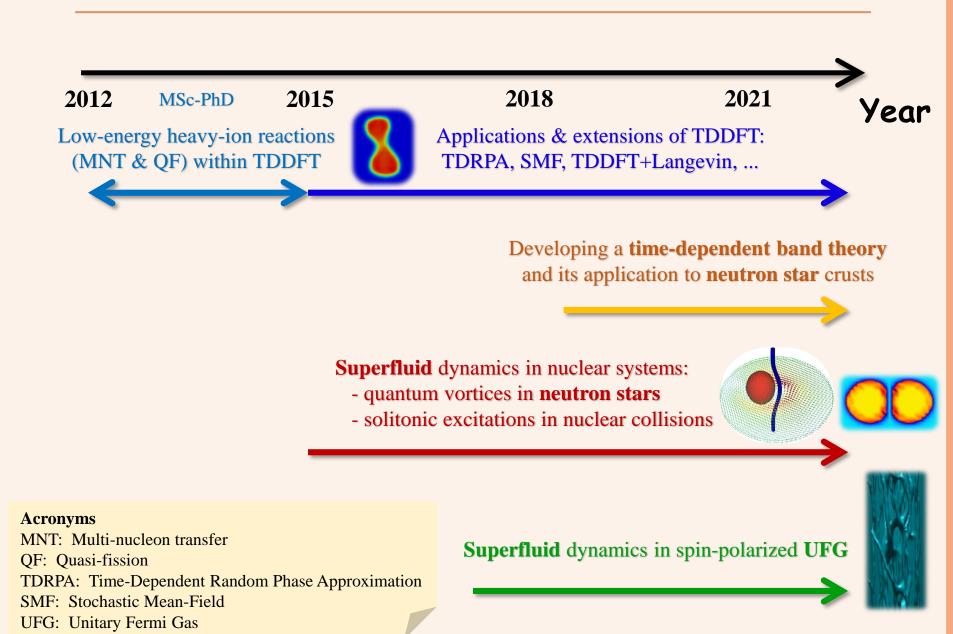








My brief research history



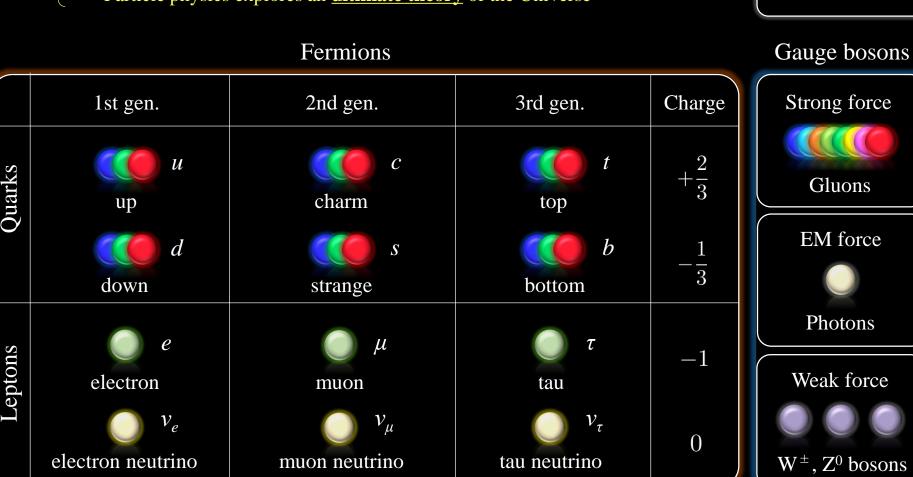
From quarks to atomic nuclei

Standard model of the elementary particles

Higgs boson

- Elementary particles: fundamental particles without structure
- Four forces: strong, weak, electromagnetic, and gravitational forces \checkmark
- Particle physics explores an **<u>ultimate theory</u>** of the Universe

Quarks



From quarks to hadrons



The QCD phase diagram

Early universe

LHC

Exploring the evolution of the Universe through high-energy nuclear experiments

Quark Gluon Plasma (QGP)

Atomic

Nuclei

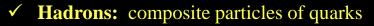
RHIC

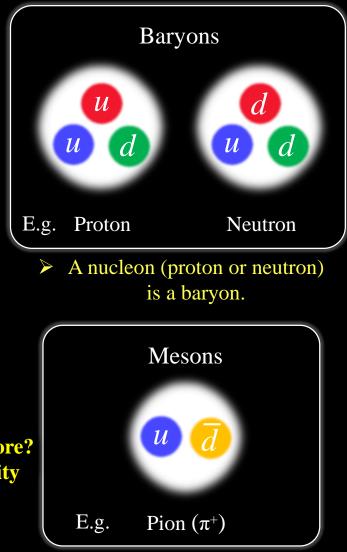
Hadrons

Quark matter in NS core? → color superconductivity

Density

Neutron Stars





Left figure: https://www.bnl.gov/newsroom/news.php?a=24281

Energy scales and degrees of freedom

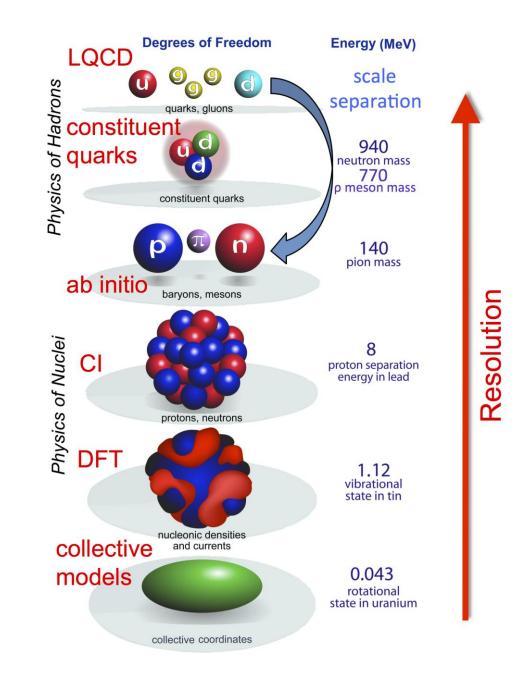


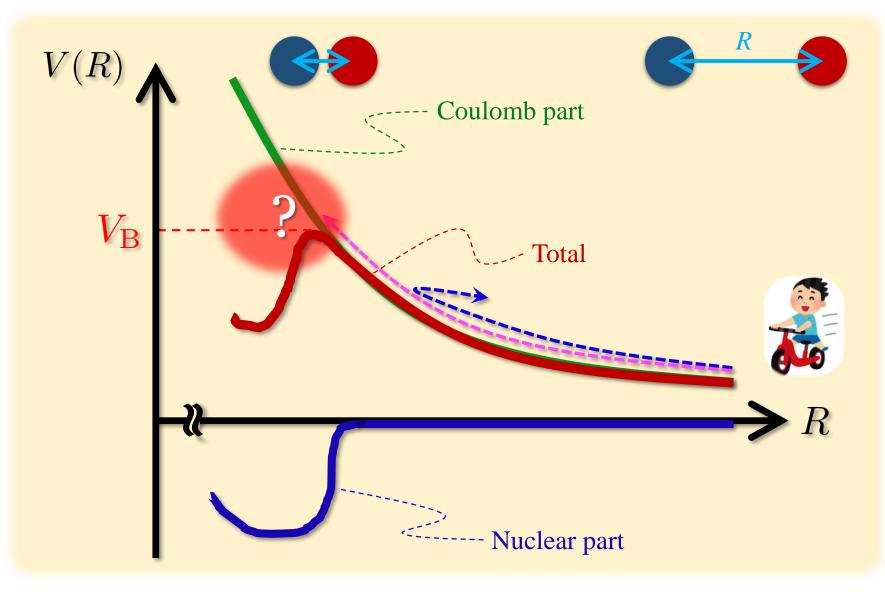
Figure: https://www.asc.ohio-state.edu/physics/ntg/6805/slides/6805_overview_slides.php#Dofs

Not this "high energy" !!

Animation: Brookhaven National Laboratory (http://www.bnl.gov/rhic/); https://youtu.be/Vyq_AYWctSo

We collide two nuclei "gently"

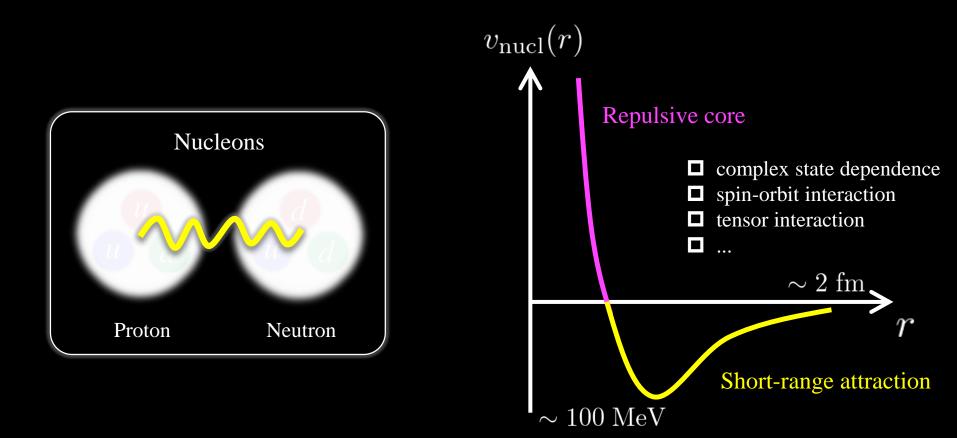
and study quantum many-body dynamics of neutrons and protons

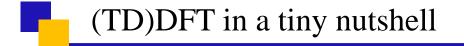


In "low-energy" nuclear physics, we treat neutrons and protons as building blocks

What we study is:

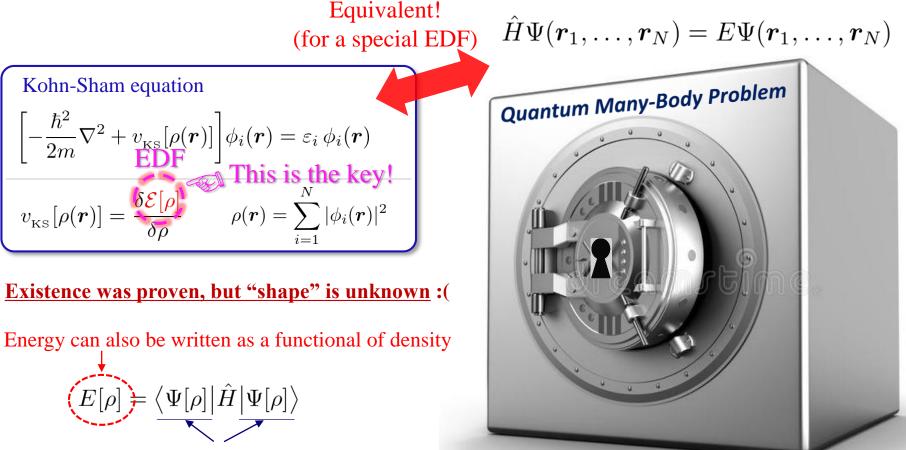
A quantum many-body problem of fermions interacting through the nuclear force







A theory which gives us access to the *exact* solution

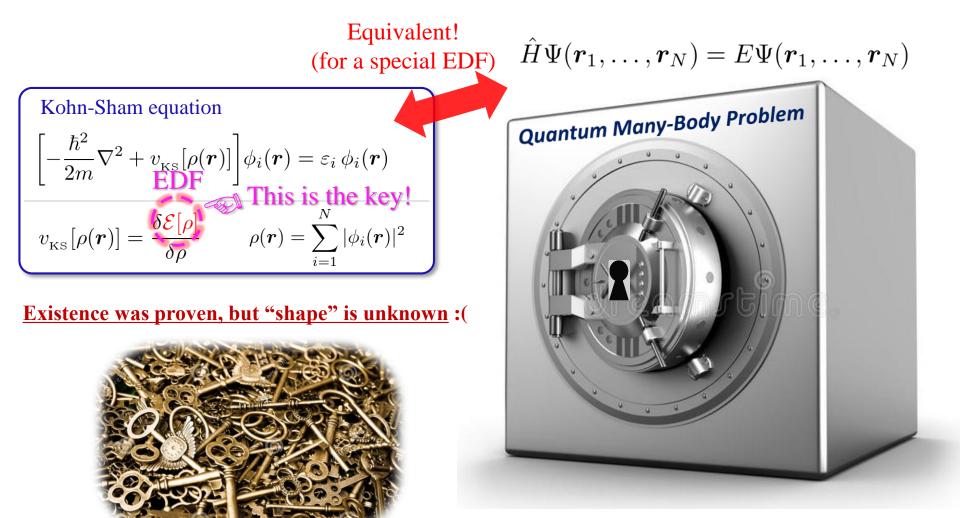


w.f. is a functional of density

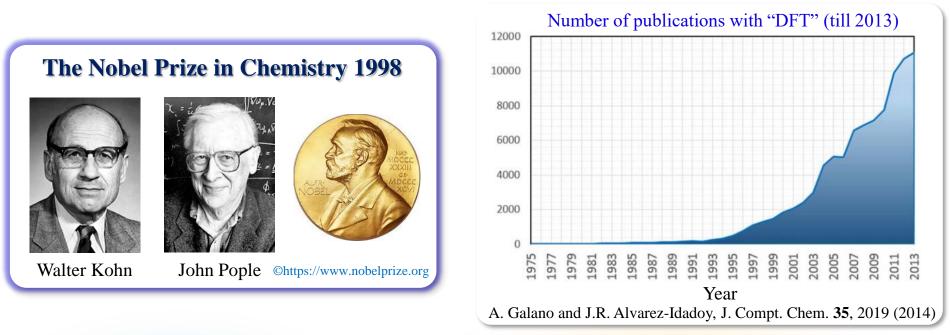
P. Hohenberg and W. Kohn, Phys. Rev. B 136, 864 (1964)

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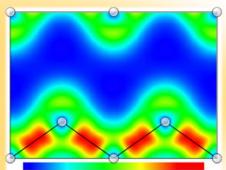
A theory which gives us access to the *exact* solution



Great Success of the Density Functional Theory



Si crystal



0.0e+000 2.5e-002 5.0e-002 7.5e-002 1.0e-001

Y. Shinohara, K. Yabana, Y. Kawashita, J.-I. Iwata, T. Otobe, and G. F. Bertsch, Phys. Rev. B 82, 155110 (2010)

The seminal papers on DFT

C-Z. Gao et al.,

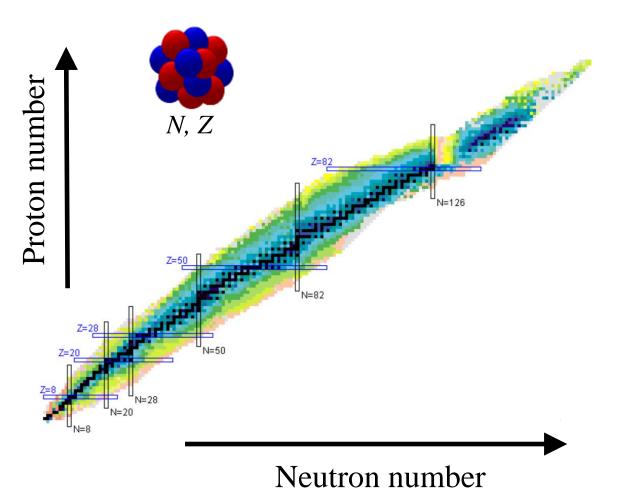
Fullerene: C₆₀

J. Phys. B: At. Mol. Opt. Phys. 48, 105102 (2015)

P. Hohenberg and W. Kohn, Phys. Rev. 136, B864 (1964) 19,015 citations!
W. Kohn and L.J. Sham, Phys. Rev. 140, A1133 (1965) 24,384 citations!

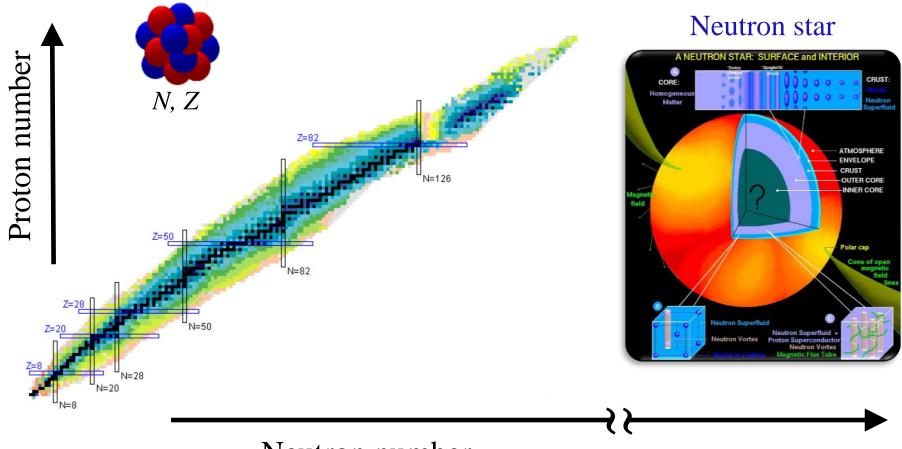
Then, how it works?

All nuclei can be described with a single EDF



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All nuclei can be described with a single EDF

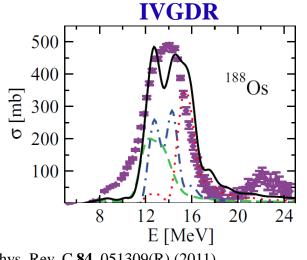


Neutron number

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TDDFT in Nuclear Physics

TDDFT is a versatile tool!!



Phys. Rev. C 84, 051309(R) (2011) I. Stetcu, A. Bulgac, P. Magierski, and K.J. Roche

Vortex-nucleus dynamics

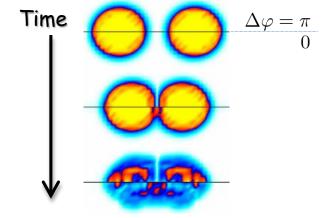
Phys. Rev. Lett. **117**, 232701 (2016) G. Wlazłowski, K.S., P. Magierski, A. Bulgac, and M.M. Forbes

Induced fission of ²⁴⁰Pu



Phys. Rev. Lett. **116**, 122504 (2016) A. Bulgac, P. Magierski, K.J. Roche, and I. Stetcu

Low-energy heavy-ion reactions



Phys. Rev. Lett. **119**, 042501 (2017) P. Magierski, K.S., and G. Wlazłowski

Voyage towards the limit of existence

The continent of stability has been well explored..

The Great Wall of China by Stacy Funderburke @shutterstock

Now we are sailing towards the edge of the nuclear landscape..

Stable nuclei: 288 Experiment: ~3300 Theory: ~7000-10000

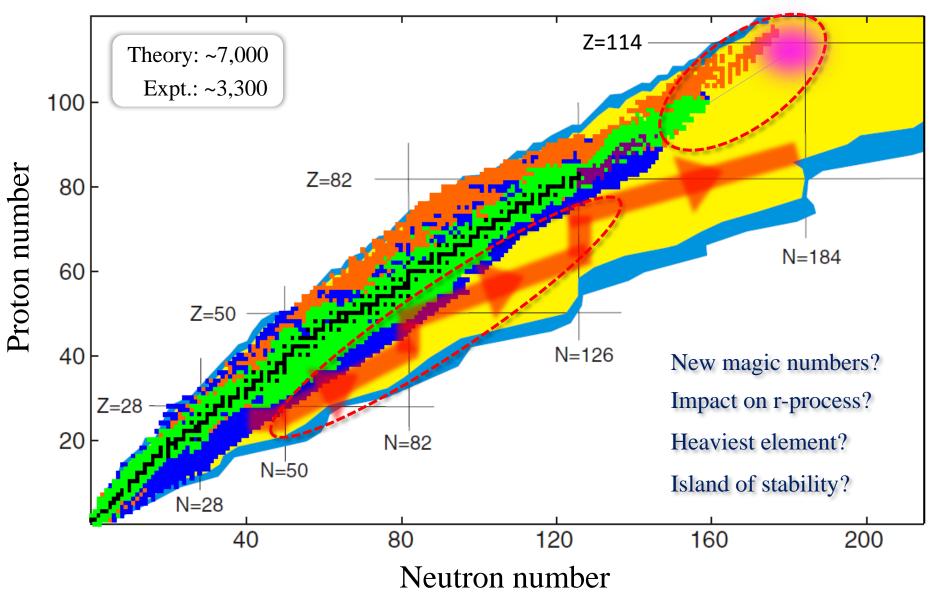
□ drip lines

- shell structure
- deformation
- skin, halo

- □ nuclear matter properties
- nucleosynthesis

"Flat Earth" by iStock; "Fanciful view of ship sailing over edge of Earth" by Georgia Studies Images

How can we create yet-unknown neutron-rich nuclei?



M. Thoennessen, Rep. Prog. Phys. 76, 056301 (2013)

At the frontiers in nuclear physics:

Synthesis of superheavy elements

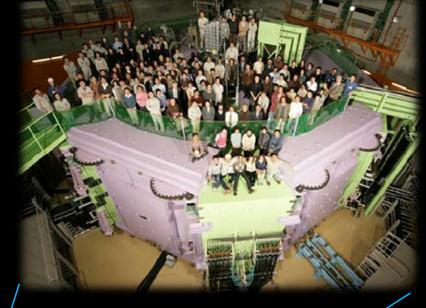
→ also relevant to **chemistry**

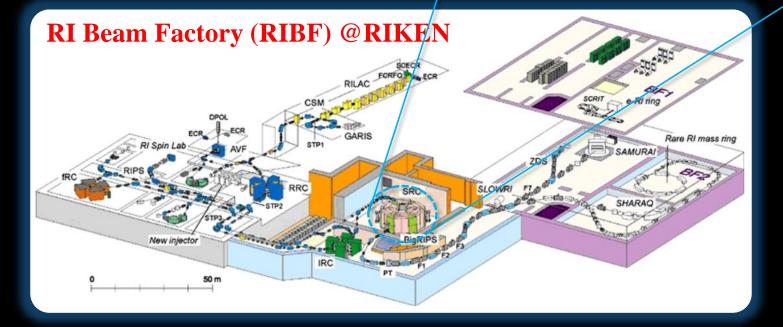
Nucleosynthesis at accelerator facilities



The world-leading factory of unstable nuclei!

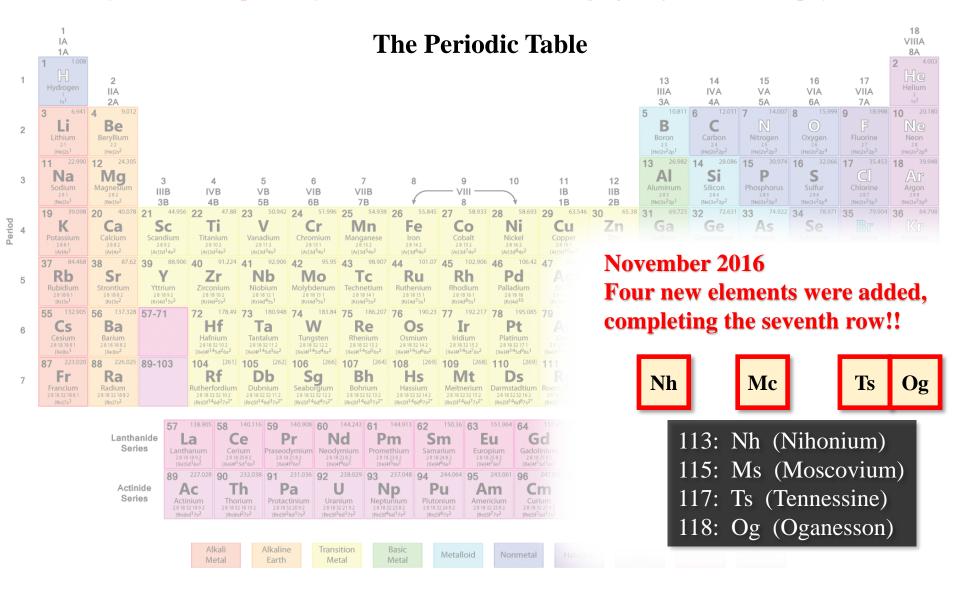
Superconducting Ring Cyclotron (SRC)





The periodic table is growing!

Synthesis of "superheavy element" is the most challenging subject in nuclear physics





Nihonium is the only one discovered by Asian country: The element 113

Nov. 30, 2016: Name confirmed June 8, 2016: Name announced Dec. 31, 2015: Naming rights given



2019 was the 150th anniversary of the Periodic Table!

2019

IYPT 2019 CLOSING CEREMONY

2621

m

n

2661

praseodymium

140.91

protactinium

neodymium

60

144.24

Iranium

5 December, Tokyo Prince Hotel

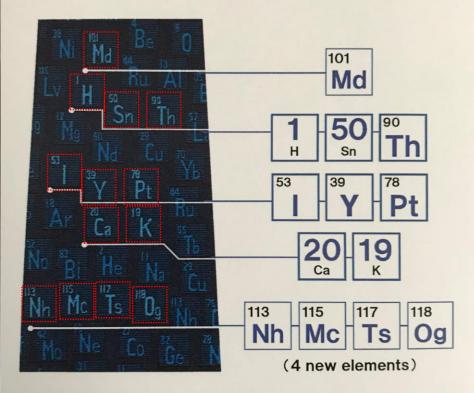
wifi D:ClesingCoremony1YPT2019 Password:20191205

Dec. 5, 2019 @Tokyo Prince Hotel



IYPT2019 Special Tie of the Elements

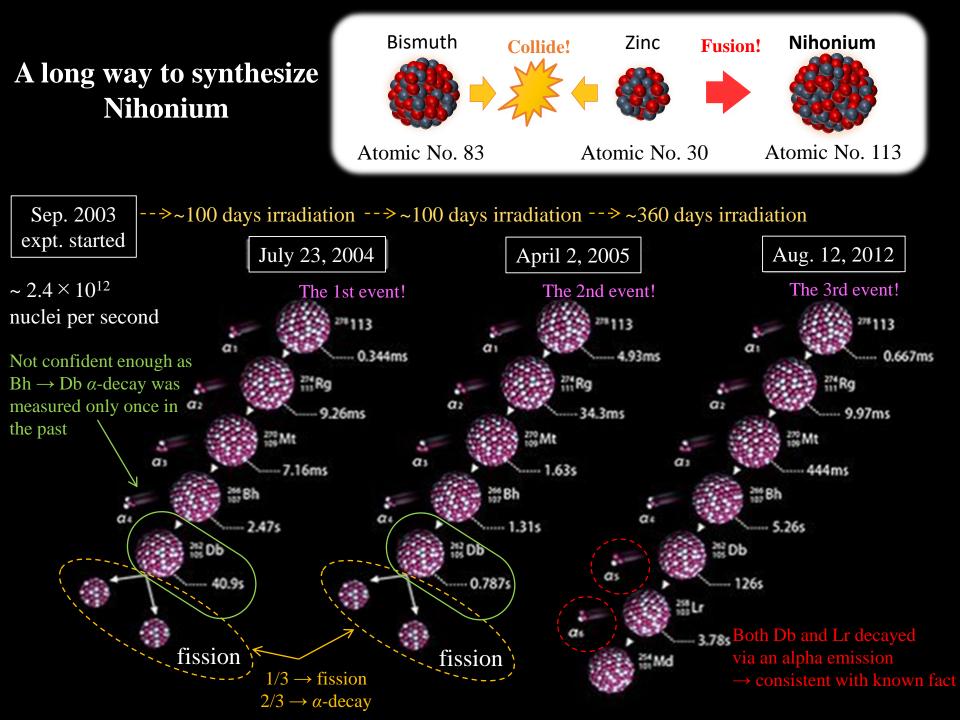
How special?



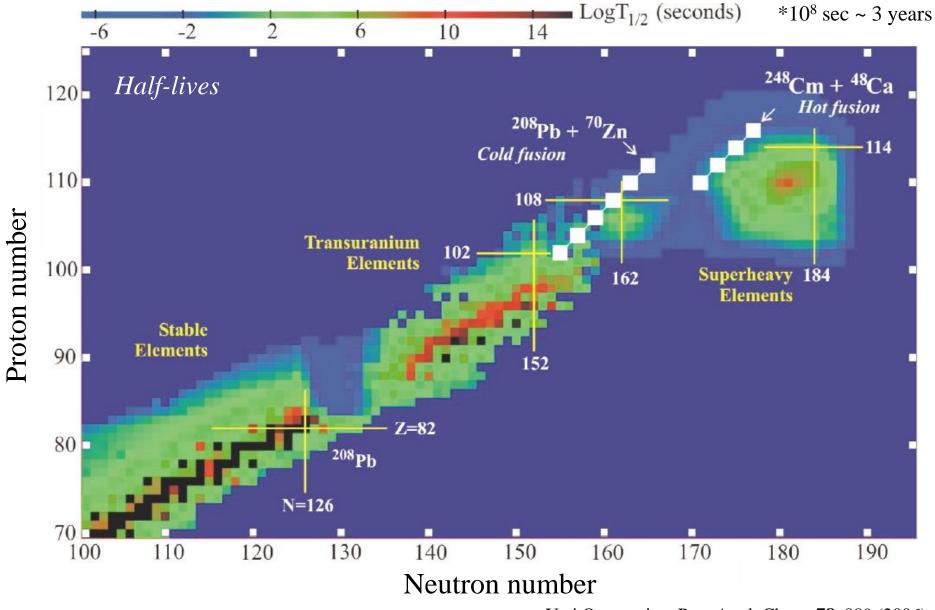


Designed by GINZA TAYA, Tokyo & Kohei Tamao

◄ VISIT OUR WEBSITE



The north-east part of the nuclear map



Yuri Oganessian, Pure Appl. Chem. 78, 889 (2006)

The north-east part of the nuclear map

We do need theoretical predictions!

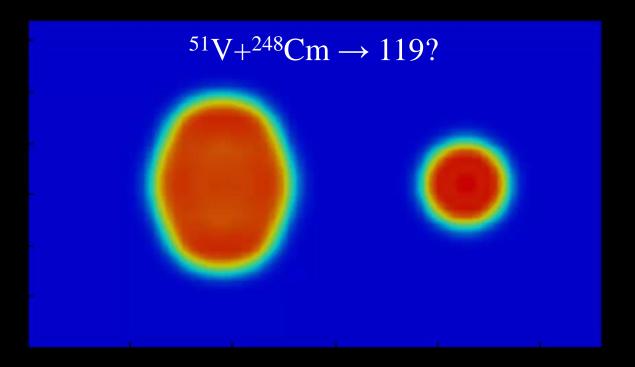






The next is: 119 or 120

TDDFT simulation for SHE synthesis on a supercomputer



Collaborative works are in progress with Kyoto U., Kindai U., and Kyushu U. cf. K.S. and K. Hagino, Phys. Rev. C **99**, 051602(R) (2019)

Now we are sailing towards the edge of the nuclear landscape..

Let's leave the planet of finite nuclei!

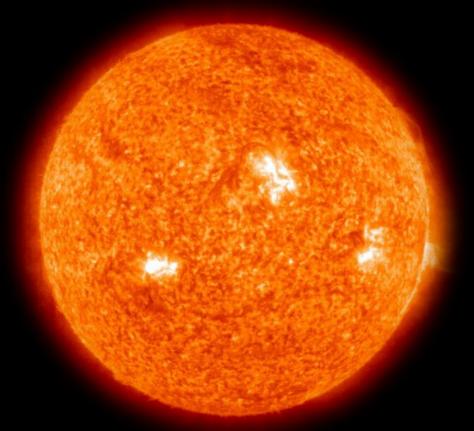
"Flat Earth" by iStock; "Star Wars Spaceship" by Valérian Pierret @Artstation

From nuclei to neutron stars

"3d hyperspace background with warp tunnel effect Free Photo" @freepik

The Sun

- Radius: $\sim 7 \times 10^8$ m (~109 times bigger than Earth)
- Mass: $\sim 2 \times 10^{30}$ kg (~330 thousands times heavier than Earth)
- Central temp.: ~10 million °C
- Surface temp.: ~5000 °C

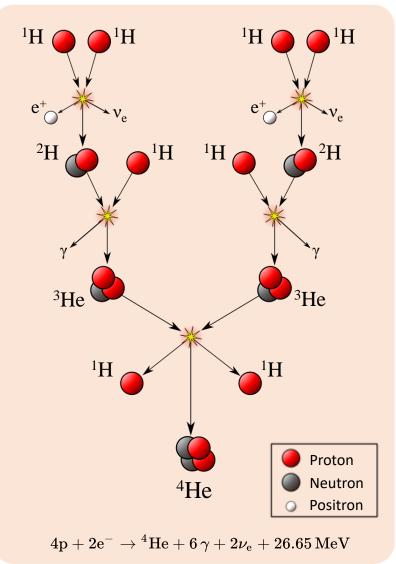


Stars shine due to nuclear fusion reactions

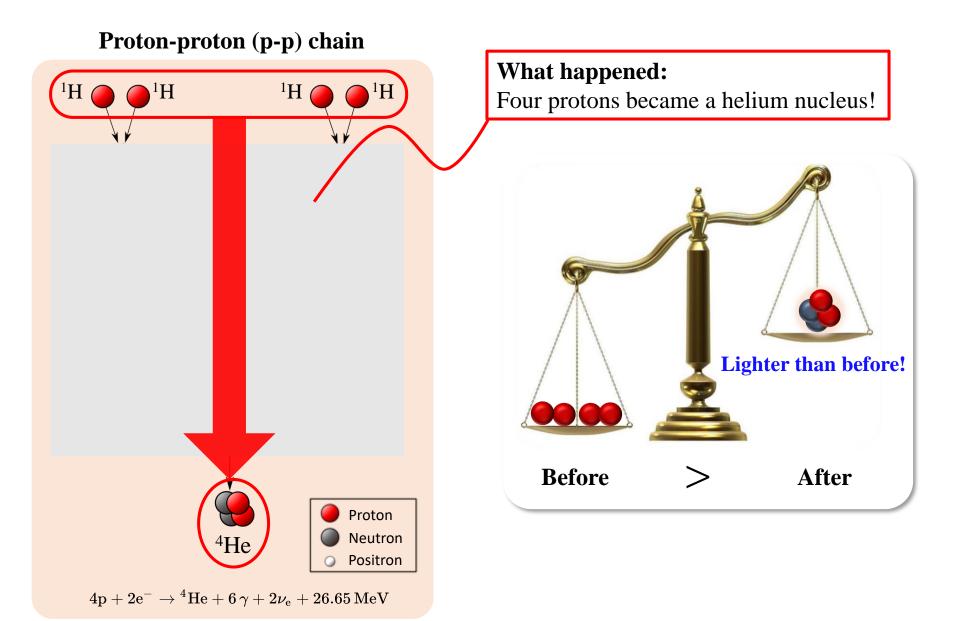
Picture: https://en.wikipedia.org/wiki/Sun

Energy source of the Sun: Nuclear fusion

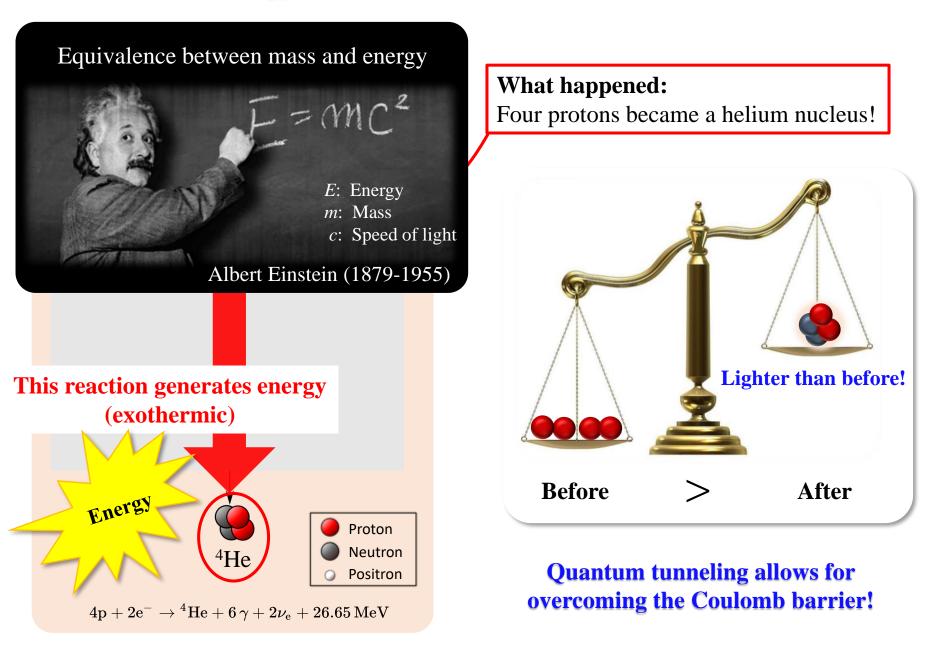
Proton-proton (p-p) chain

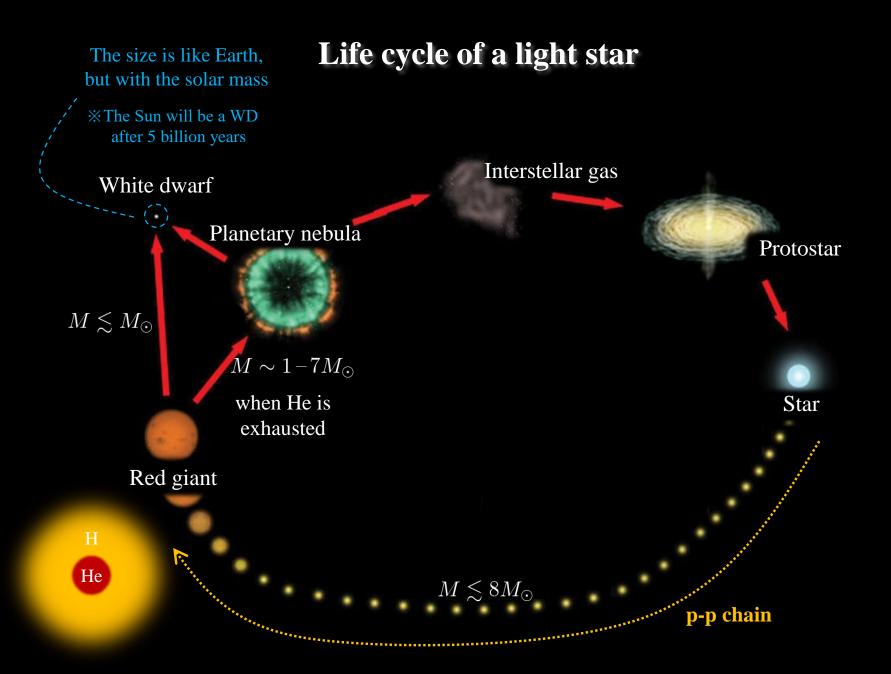


Energy source of the Sun: Nuclear fusion

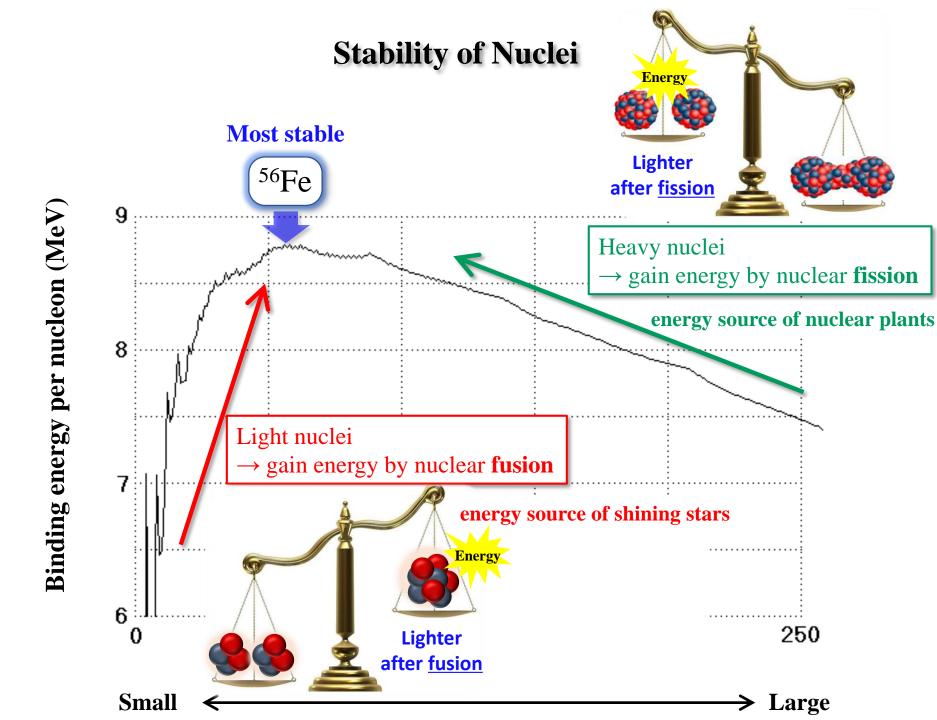


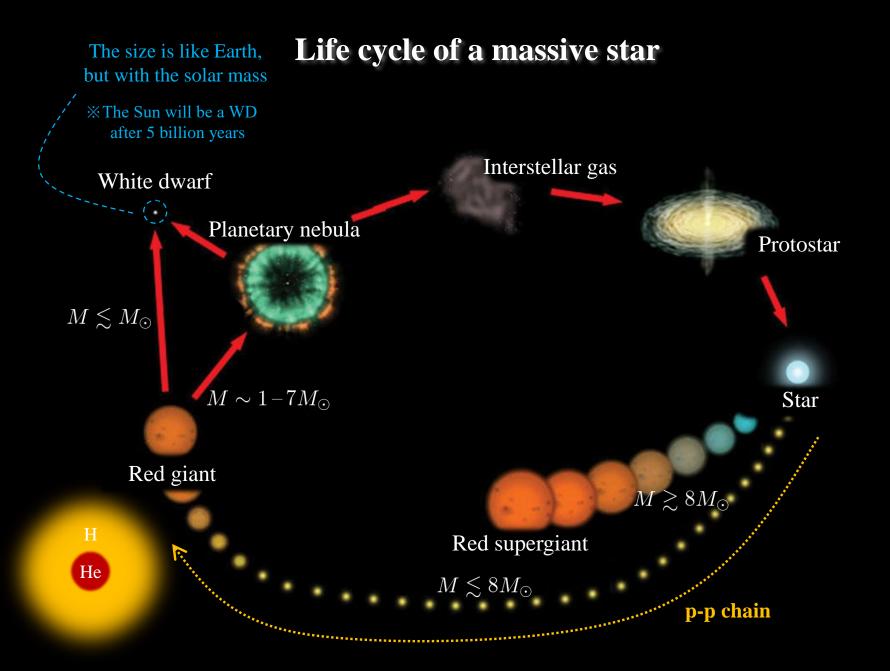
Energy source of the Sun: Nuclear fusion





Picture: https://www.astroarts.co.jp/products/stlnav7/spec/plane_list-j.shtml





Picture: https://www.astroarts.co.jp/products/stlnav7/spec/plane_list-j.shtml

The fate of a massive star

Nuclear reactions: ${}^{1}H \rightarrow {}^{4}He \rightarrow {}^{12}C \rightarrow {}^{16}O \rightarrow {}^{20}Ne \rightarrow {}^{24}Mg \rightarrow {}^{28}Si \rightarrow ... \rightarrow {}^{56}Fe$

"Onion structure"

He C, O O, Ne, Mg Si

Fe



After forming the iron core...

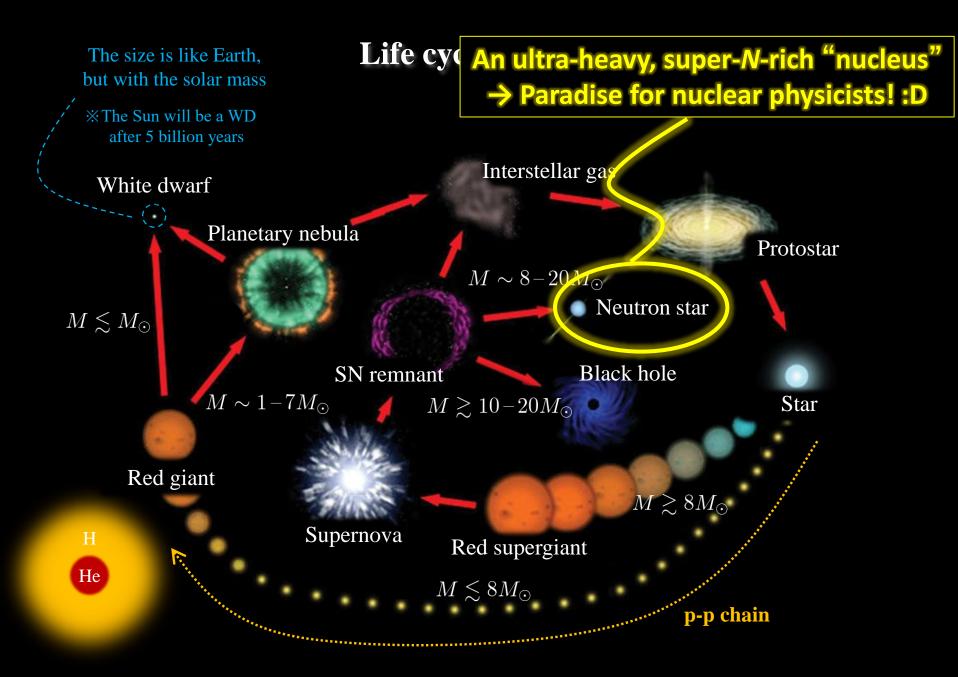
 \rightarrow no more fuel

 \rightarrow gravitational collapse

 \rightarrow supernova explosion

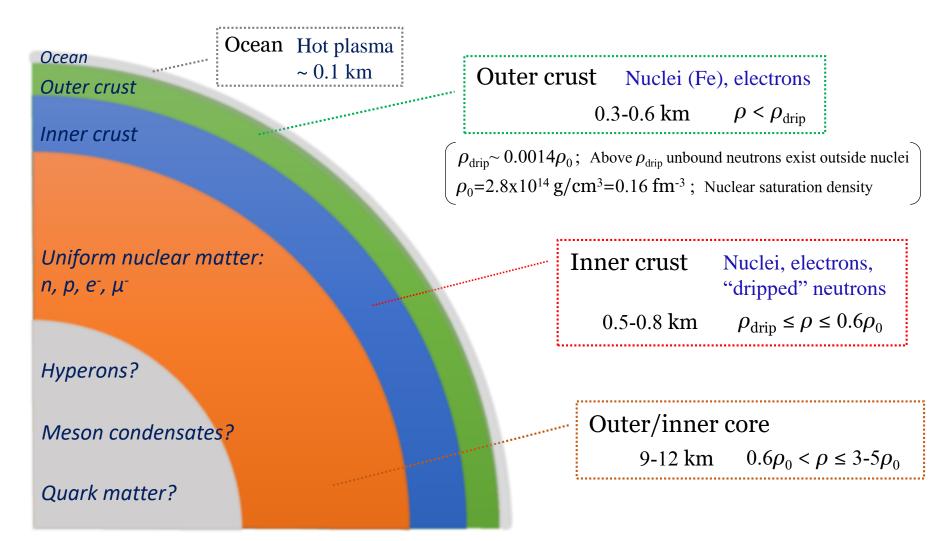
The Crab Nebula Remnant of the SN in 1054

Picture: https://en.wikipedia.org/wiki/Crab_Nebula



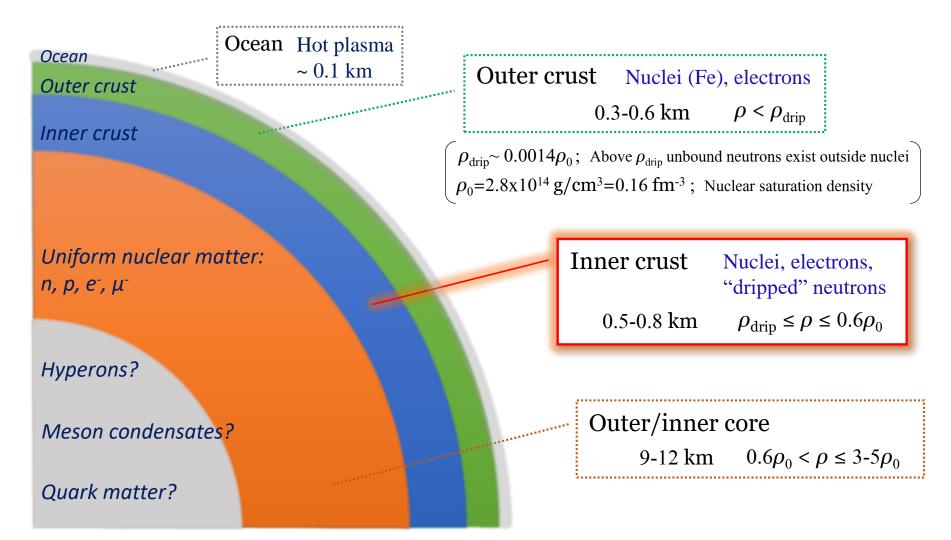
Picture: https://www.astroarts.co.jp/products/stlnav7/spec/plane_list-j.shtml

Neutron star is a great playground for nuclear physicists



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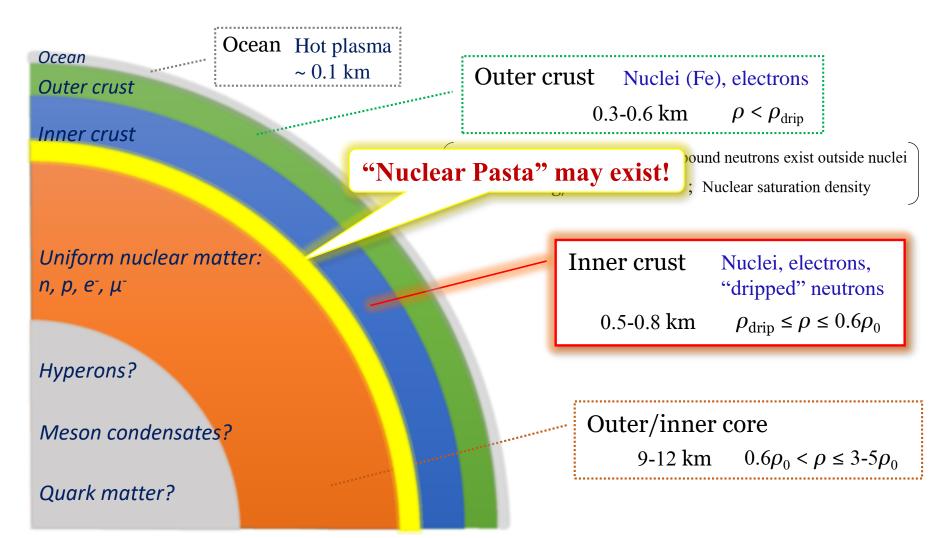
Neutron star is a great playground for nuclear physicists



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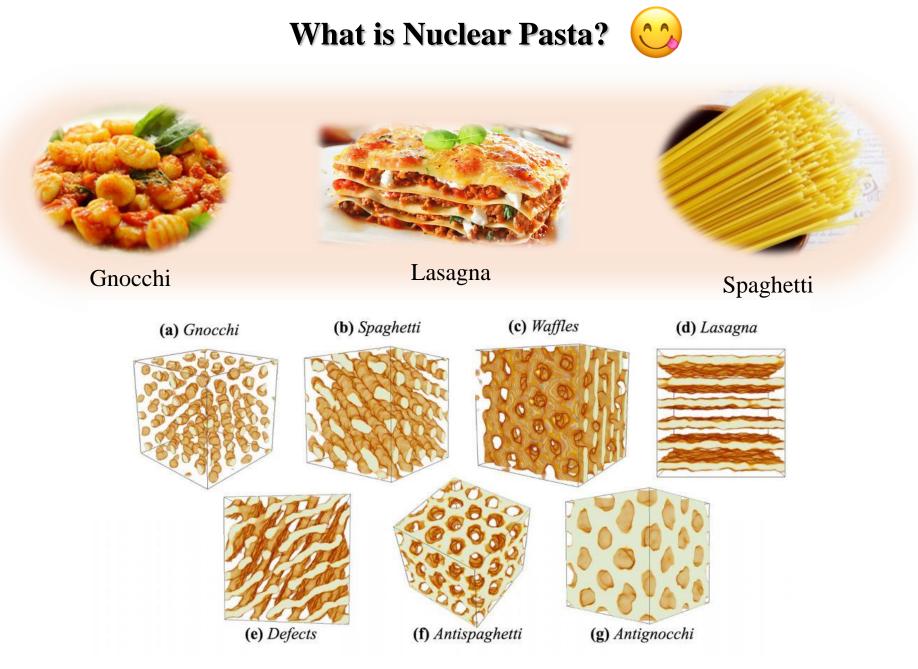
Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars

Neutron star is a great playground for nuclear physicists



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Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars



M. E. Caplan and C. J. Horowitz, Rev. Mod. Phys. 89, 041002 (2017)







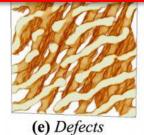
We have developed

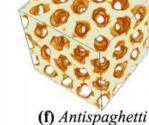
a time-dependent <u>band theory</u> based on TDDFT

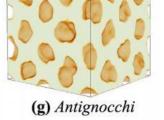
for nuclear dynamics in crystalline pasta phases

\rightarrow relevant to <u>solid-state physics</u>

K.S., S. Kobayashi, and M. Matsuo, in preparation.







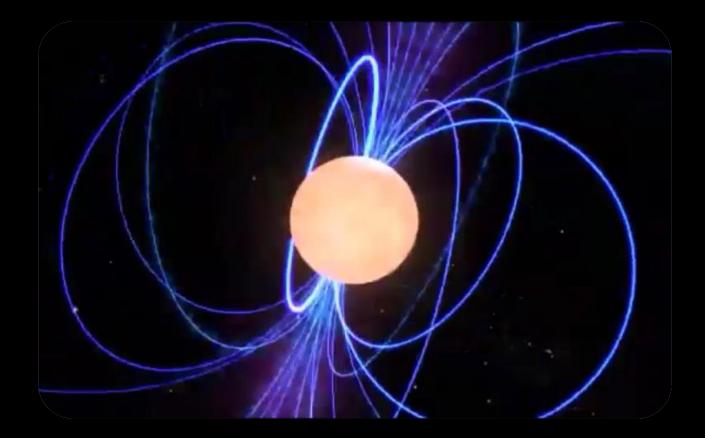
M. E. Caplan and C. J. Horowitz, Rev. Mod. Phys. 89, 041002 (2017)

Neutron-star "glitch"

Picture: https://astronomy.com/magazine/ask-astro/2017/12/stellar-magnets

Pulsar - a rotating neutron star

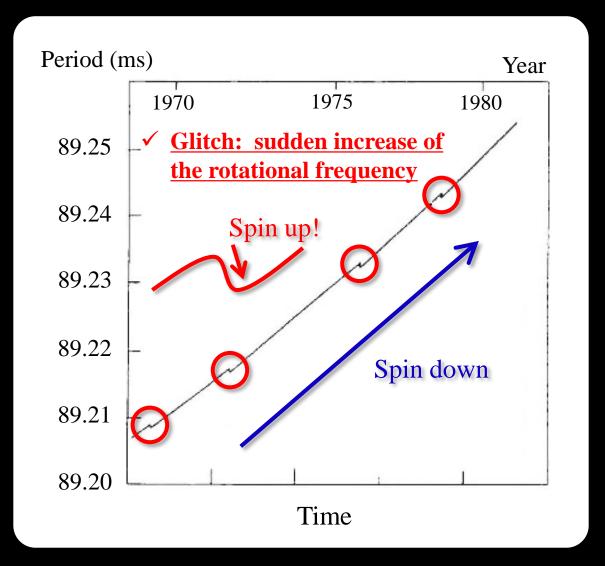
- ✓ First discovery in August 1967 → "Little Green Man" LGM-1 → PSR B1919+21
- ✓ Since then, more than 2650 pulsars have been observed
- ✓ It gradually <u>spins down</u> due to the EM radiation



What is the glitch?

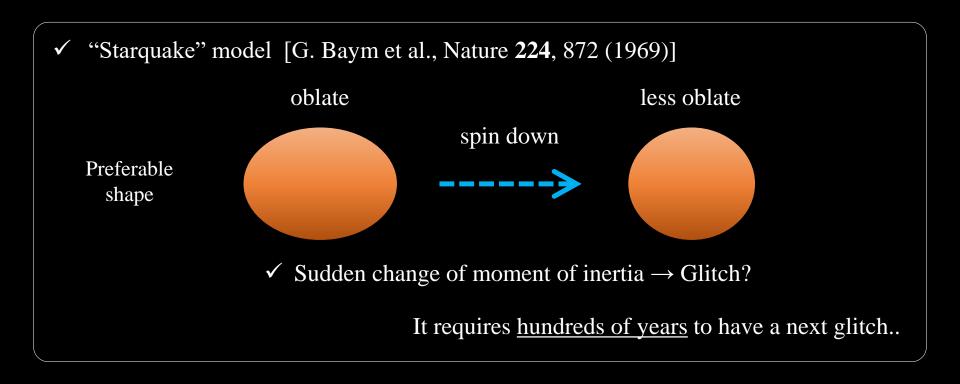
Typical example: the Vela pulsar

Irregularity has been observed from continuous monitoring of the pulsation period



What happened?

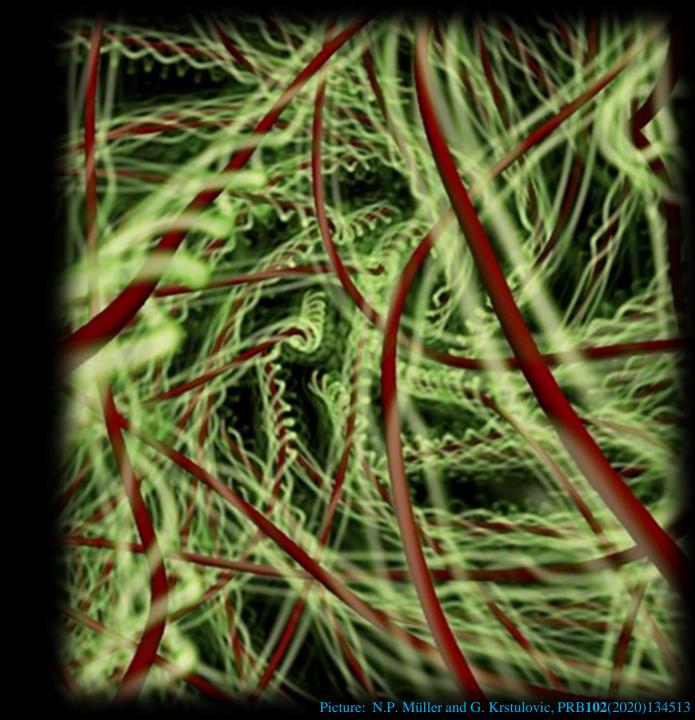
Something must happen inside the neutron star!



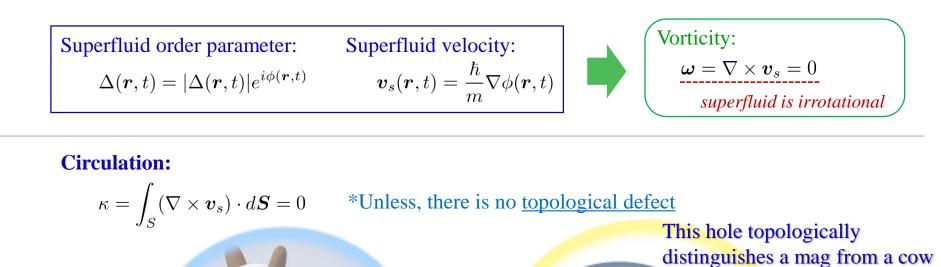
Vortex mediated glitch [P.W. Anderson and N. Itoh, Nature 256, 25 (1975)]

Dynamics of superfluid "quantized vortices" play a key role!

Quantum vortices



In superfluid, vortices are quantized!



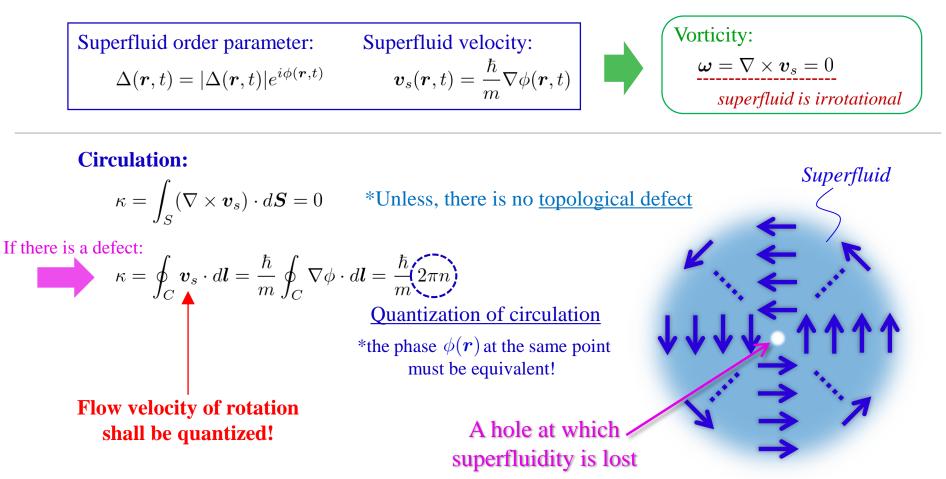
Pictures: https://en.wikipedia.org/wiki/Topology

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Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars

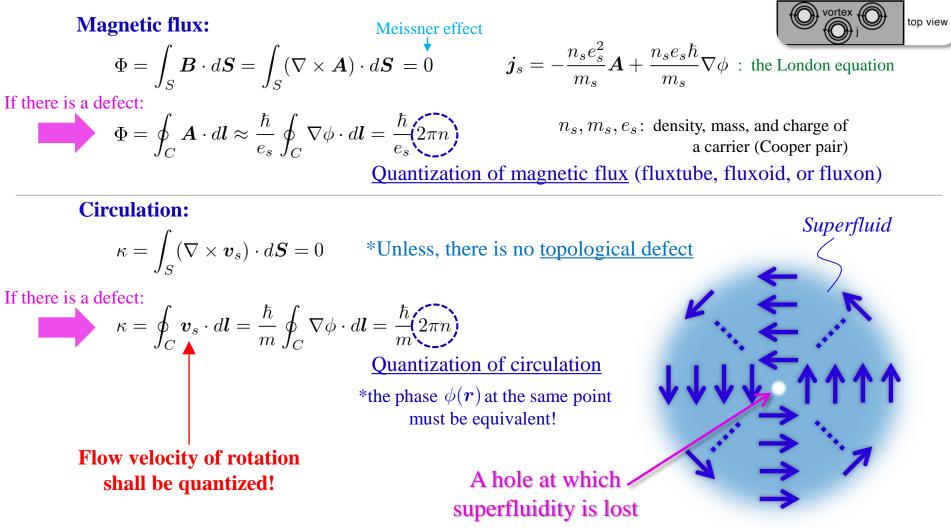
Thu., Dec. 23, 2021

In superfluid, vortices are quantized!



Quantum vortex

In superconductor, magnetic flux is quantized!



 Φ_{B}

side view

In daily life, a vortex is continuous..

In superfluid, vortices are quantized!!

G.P. Bewley, D.P. Lathrop, and K.R. Sreenivasan, Nature 441, 588 (2006)

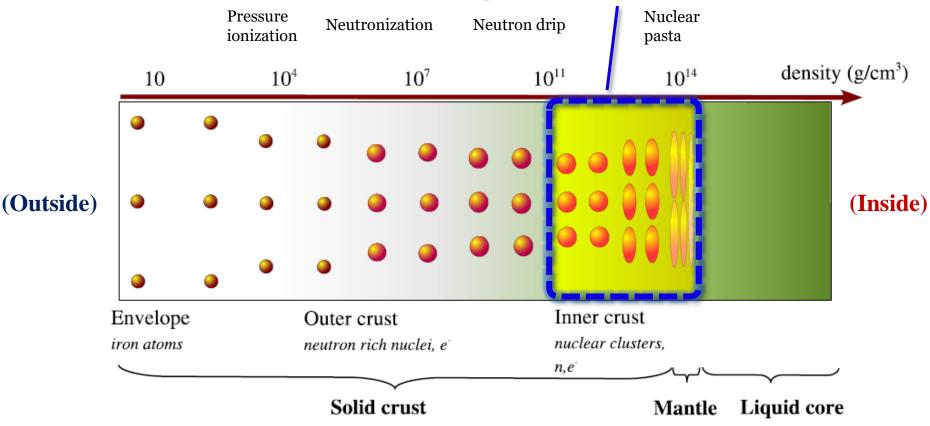
A movie from a talk by W. Guo (available from <u>https://youtu.be/P2ckefSAN20</u>) at INT Program 19-1a "Quantum Turbulence: Cold Atoms, Heavy Ions, and Neutron Stars" March 18 - April 19, 2019

Direct visualization of quantized vortices



Hydrogen particles were trapped in the vortex core, then worked as a tracer

A lattice of neutron-rich nuclei are immersed in a neutron superfluid



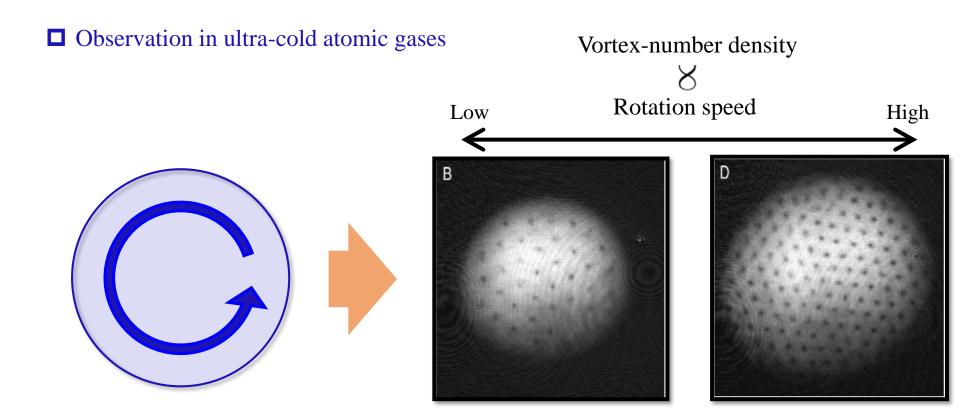
Quantum vortices can exist!

Fig.4 in N. Chamel and P. Haensel, Living Rev. Relativity 11, 10 (2008)

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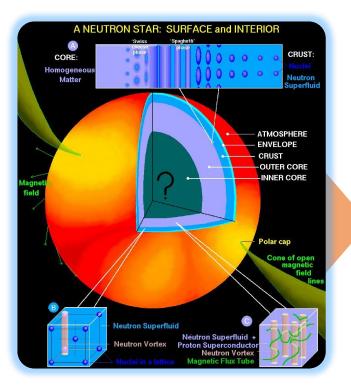
Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars Thu., Dec. 23, 2021

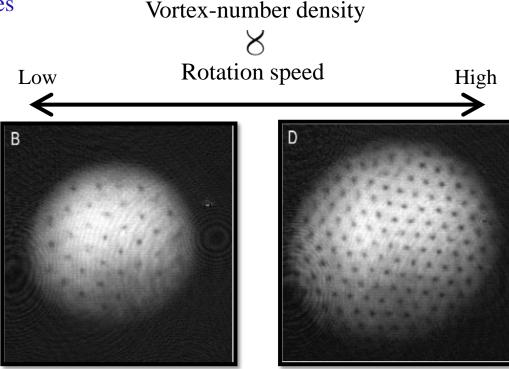
In rotating superfluid, an array of quantum vortices is generated



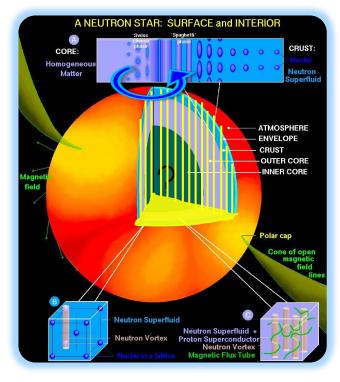
In rotating superfluid, an array of quantum vortices is generated

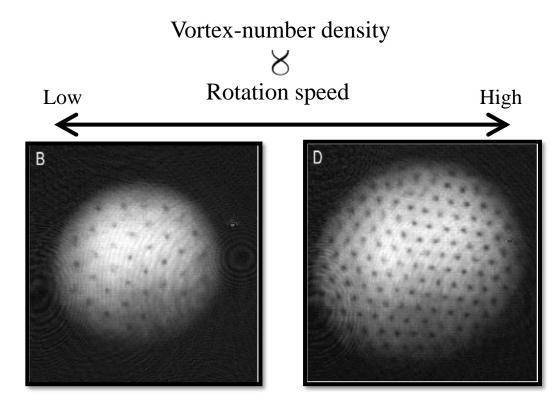
Observation in ultra-cold atomic gases



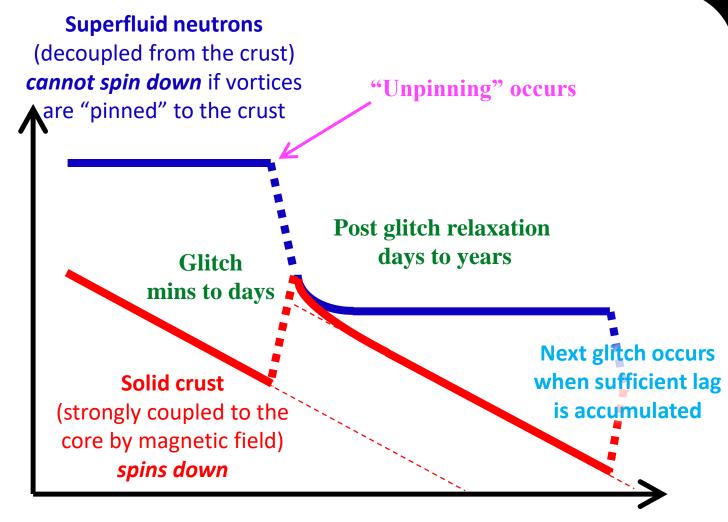


There must be a huge number (~10¹⁸) of vortices inside a neutron star!!





The vortex mediated glitch: Naive picture



Rotation frequency

Time

Recent progress:

inclusion of superfluidity/superconductivity in (TD)DFT

We can now study dynamics of quantum vortices microscopically from nucleonic degrees of freedom!

TDSLDA (Time-Dependent Superfluid Local Density Approximation)

TDSLDA: TDDFT with local treatment of pairing

Kohn-Sham scheme is extended for non-interacting quasiparticles

TDSLDA equations (formally equivalent to TDHFB or TD-BdG equations)

$$i\hbar\frac{\partial}{\partial t}\begin{pmatrix}u_{k,\uparrow}(\boldsymbol{r},t)\\u_{k,\downarrow}(\boldsymbol{r},t)\\v_{k,\uparrow}(\boldsymbol{r},t)\\v_{k,\downarrow}(\boldsymbol{r},t)\end{pmatrix} = \begin{pmatrix}h_{\uparrow\uparrow}(\boldsymbol{r},t) & h_{\uparrow\downarrow}(\boldsymbol{r},t) & 0 & \Delta(\boldsymbol{r},t)\\h_{\downarrow\uparrow}(\boldsymbol{r},t) & h_{\downarrow\downarrow}(\boldsymbol{r},t) & -\Delta(\boldsymbol{r},t) & 0\\0 & -\Delta^{*}(\boldsymbol{r},t) & -h_{\uparrow\uparrow}^{*}(\boldsymbol{r},t) & -h_{\uparrow\downarrow}^{*}(\boldsymbol{r},t)\\\Delta^{*}(\boldsymbol{r},t) & 0 & -h_{\downarrow\uparrow}^{*}(\boldsymbol{r},t) & -h_{\downarrow\downarrow}^{*}(\boldsymbol{r},t)\end{pmatrix} \begin{pmatrix}u_{k,\uparrow}(\boldsymbol{r},t)\\u_{k,\downarrow}(\boldsymbol{r},t)\\v_{k,\uparrow}(\boldsymbol{r},t)\\v_{k,\downarrow}(\boldsymbol{r},t)\end{pmatrix}$$

$$h_{\sigma} = \frac{\delta E}{\delta n_{\sigma}} \quad : \text{ s.p. Hamiltonian} \qquad \qquad n_{\sigma}(\boldsymbol{r}, t) = \sum_{E_k < E_c} |v_{k,\sigma}(\boldsymbol{r}, t)|^2 \quad : \text{ number density} \\ \Delta = -\frac{\delta E}{\delta \nu^*} \quad : \text{ pairing field} \qquad \qquad \nu(\boldsymbol{r}, t) = \sum_{E_k < E_c} u_{k,\uparrow}(\boldsymbol{r}, t) v_{k,\downarrow}^*(\boldsymbol{r}, t) \quad : \text{ anomalous density} \\ \boldsymbol{j}_{\sigma}(\boldsymbol{r}, t) = \hbar \sum_{E_k < E_c} \operatorname{Im}[v_{k,\sigma}^*(\boldsymbol{r}, t) \boldsymbol{\nabla} v_{k,\sigma}(\boldsymbol{r}, t)] \quad : \text{ current} \end{cases}$$

A large number (10⁴-10⁶) of 3D coupled non-linear PDEs have to be solved!! # of qp orbitals ~ # of grid points

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TDSLDA (Time-Dependent Superfluid Local Density Approximation)

TDSLDA: TDDFT with local treatment of pairing

Kohn-Sham scheme is extended for non-interacting quasiparticles

TDSLDA equations (formally equivalent to TDHFB or TD-BdG equations)

A large number (10⁴-10⁶) of 3D coupled non-linear PDEs have to be solved!! # of qp orbitals ~ # of grid points

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*The number indicates the rank according to the TOP500 list (November 2021)

Piz Daint, CSCS, Switzerland (No. 20)

TITAN, ORNL, USA

TSUBAME3.0, Japan (No. 59)

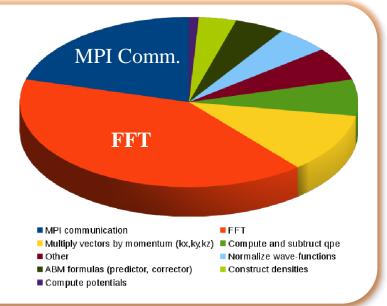




Summit, ORNL, USA (No. 2) GPU, 200 PFlops/s

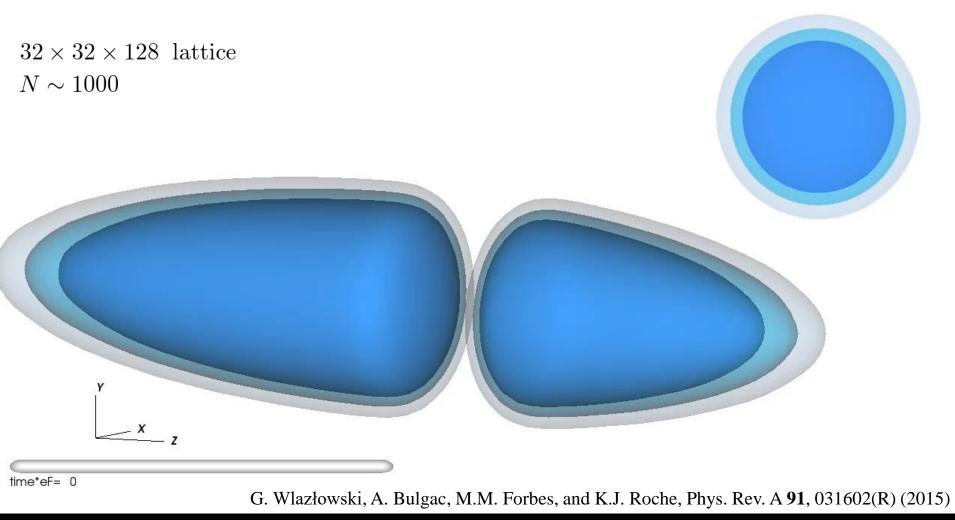
Present computing capabilities:

- ✓ Full 3D (w/o symmetry restrictions)
- \checkmark Volume as large as 100³ lattice points
- ✓ Evolution up to 10^6 time steps (as long as 10^{-19} sec)



Result of TDSLDA simulation:

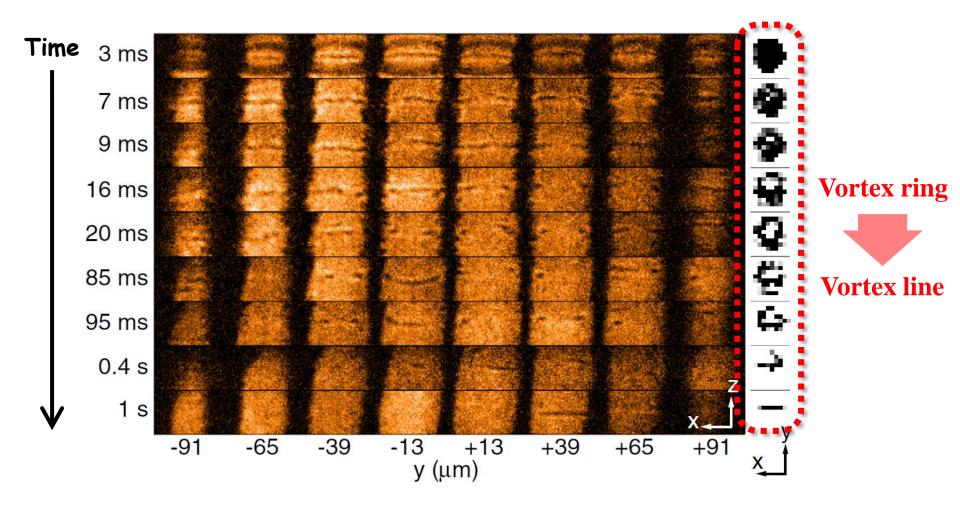
Phase discontinuity creates a vortex ring which decays into a vortex line



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The cascades of solitonic excitations have been identified experimentally

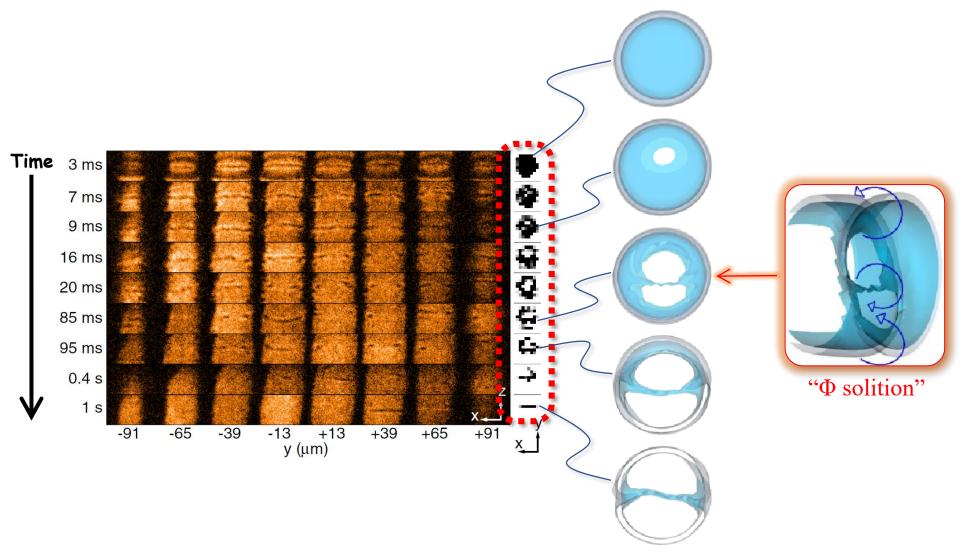


M.J.H. Ku, B. Mukherjee, T. Yefsah, and M.W. Zwierlein, Phys. Rev. Lett. 116, 045304 (2016)

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Exploring Quantum Many-Body Dynamics in Nuclear Systems: From Nuclei to Neutron Stars

Each stage of solitonic cascade could be reproduced with TDSLDA!



G. Wlazłowski, K.S. M. Marchwiany, and P. Magierski, Phys. Rev. Lett. 120, 253002 (2018)

To fully understand the glitches, we need to clarify:

Glitch dynamics

How do vortices move?

and, of course, details of NS matter..

Pinning mechanism

How are vortices pinned?

Trigger mechanism

How are vortices unpinned?

We attacked this problem using the state-of-the-art microscopic nuclear theory

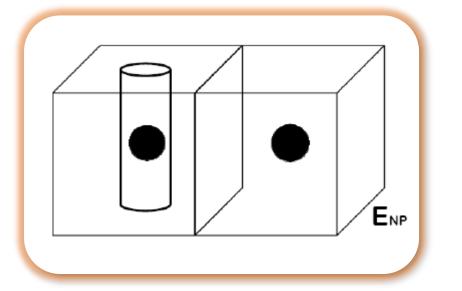
A key to understand the glitches is: <u>Vortex pinning mechanism in the inner crust of neutron stars</u>

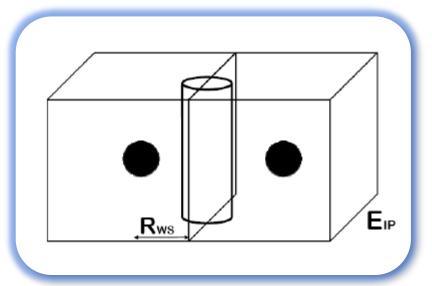
Q. Is the vortex-nucleus interaction

Attractive?

or



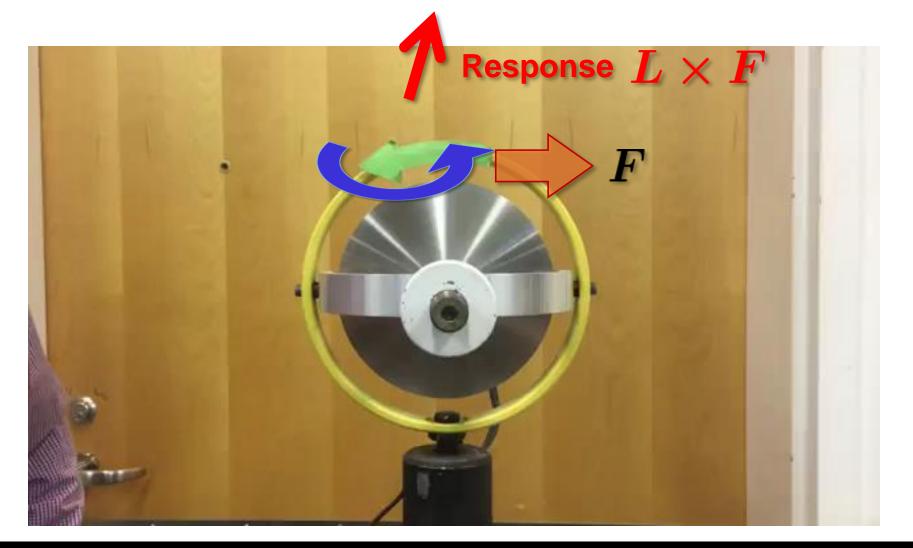




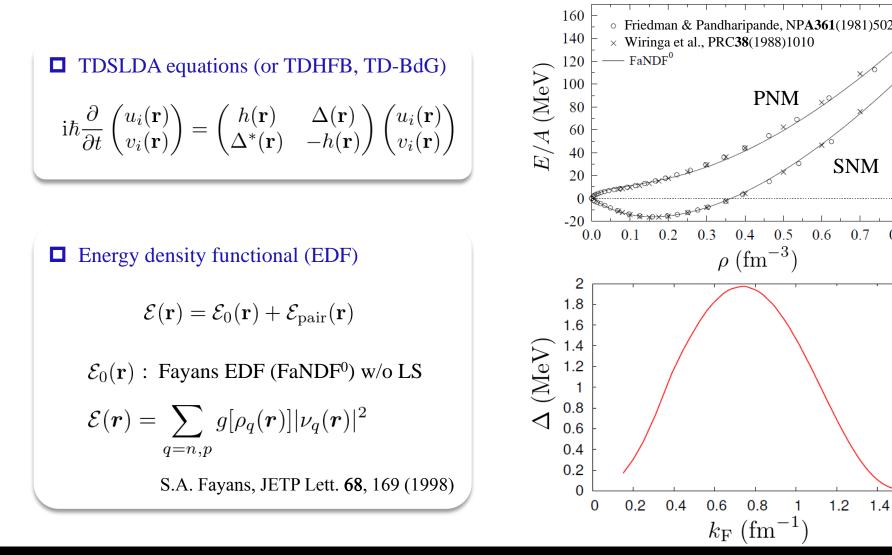
"Nuclear pinning"

"Interstitial pinning"

Response of a spinning gyroscope when pushed



K. Sekizawa



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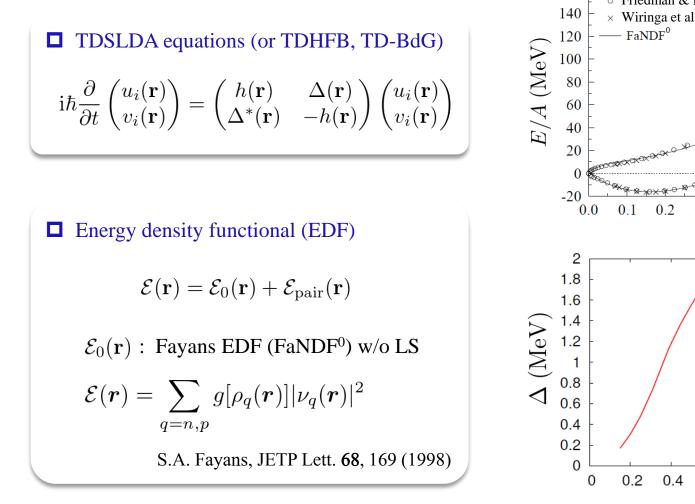
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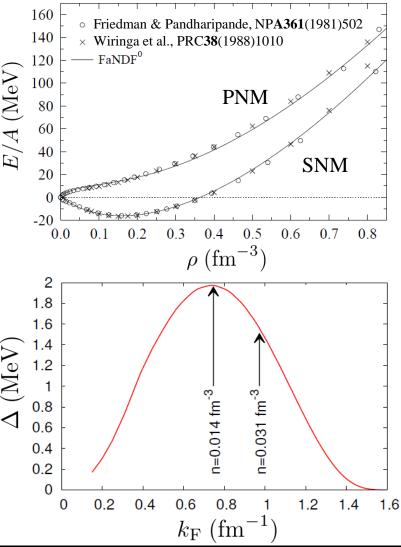
0.7

1.4

1.6

0.8





K. Sekizawa

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix} = \begin{pmatrix} h(\mathbf{r}) & \Delta(\mathbf{r}) \\ \Delta^*(\mathbf{r}) & -h(\mathbf{r}) \end{pmatrix} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix}$$

Computational details

75 fm × 75 fm × 60 fm $(50 \times 50 \times 40, \ \Delta x = 1.5 \text{ fm})$ $k_{\rm c} = \pi/\Delta x > k_{\rm F}$ $k_{\rm F} = (3\pi^2 \rho_n)^{1/3}$ Nuclear impurity: Z = 50 $\rho_n \simeq 0.014 \text{ fm}^{-3} (N \simeq 2,530)$ $\rho_n \simeq 0.031 \text{ fm}^{-3} (N \simeq 5,714)$ # of quasi-particle w.f. $\approx 100,000$

20 30 R=30fm 50 60 55 45 70 Z=50 $\rho(\mathbf{r})$ 50 40 30 20 10 $\rho_n \simeq 0.014 \, \mathrm{fm}^{-3}$

a vortex line exists here

Thu., Dec. 23, 2021

TDSLDA equations (or TDHFB, TD-BdG)

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix} = \begin{pmatrix} h(\mathbf{r}) & \Delta(\mathbf{r}) \\ \Delta^*(\mathbf{r}) & -h(\mathbf{r}) \end{pmatrix} \begin{pmatrix} u_i(\mathbf{r}) \\ v_i(\mathbf{r}) \end{pmatrix}$$

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MPI+GPU → 48h w/ 200GPUs for 10,000 fm/c



TITAN, Oak Ridge

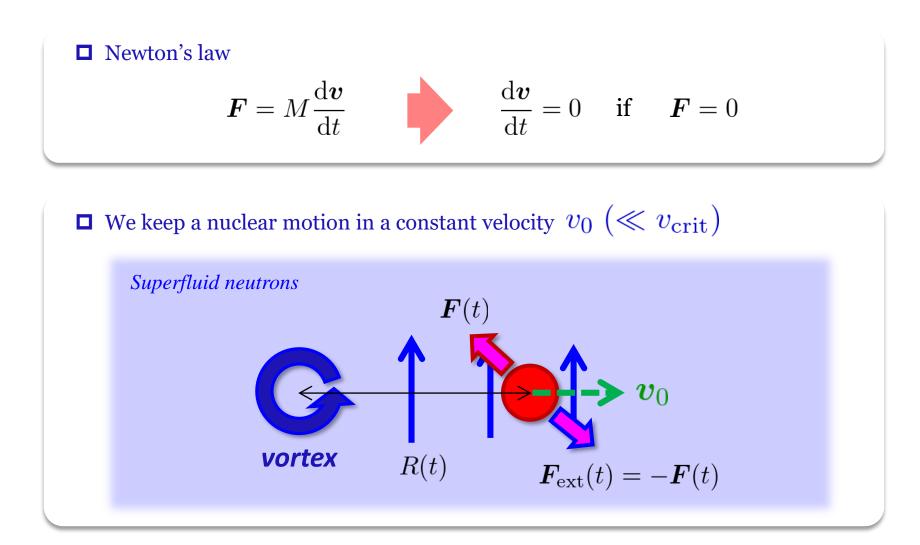


NERSC Edison, Berkeley



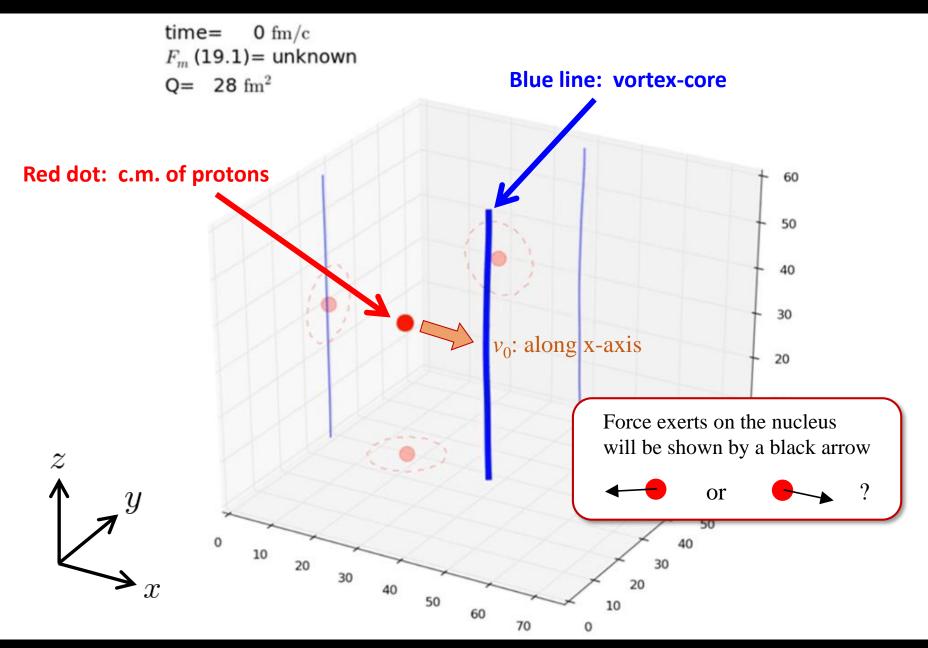
HA-PACS, Tsukuba

We directly measure the force F(R) in dynamical simulation



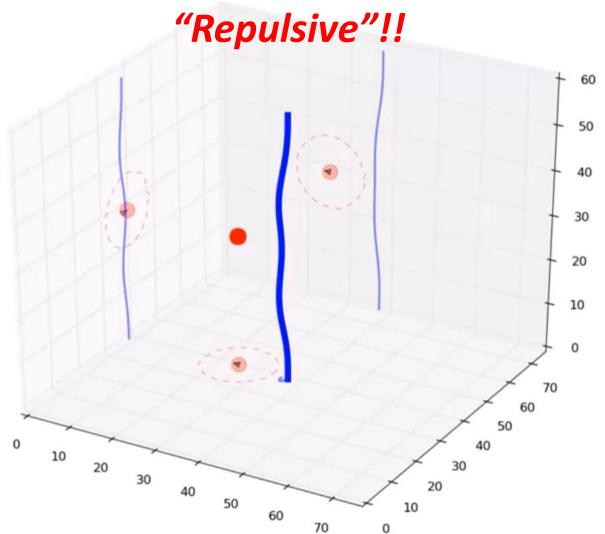
K. Sekizawa

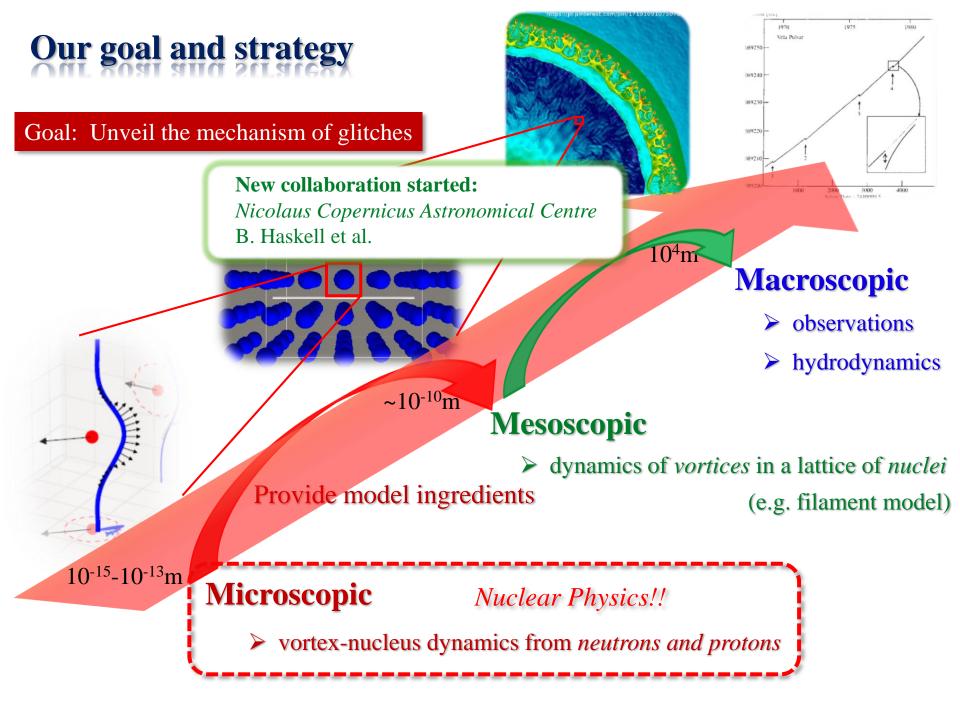
Results of TDSLDA calculation: $\rho_n \simeq 0.014 \text{ fm}^{-3}$



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time= 8032 fm/c F_m (10.6)= 0.17 MeV/fm Q= 13 fm²





Fe is the most stable!



Well, then how were elements heavier than iron produced!?

One of the unsolved problems in Physics

but, we have learned a lot!

Neutron Star Merger!!

Relevant to gravitational waves, nucleosynthesis, as well as neutron stars

At the frontier in Physics!!

Gravitational waves and gamma rays were detected on Aug. 17, 2017

NASA's Goddard Space Flight Center





I hope you enjoyed our exciting adventure!



with a magic of DFT and TDDFT..

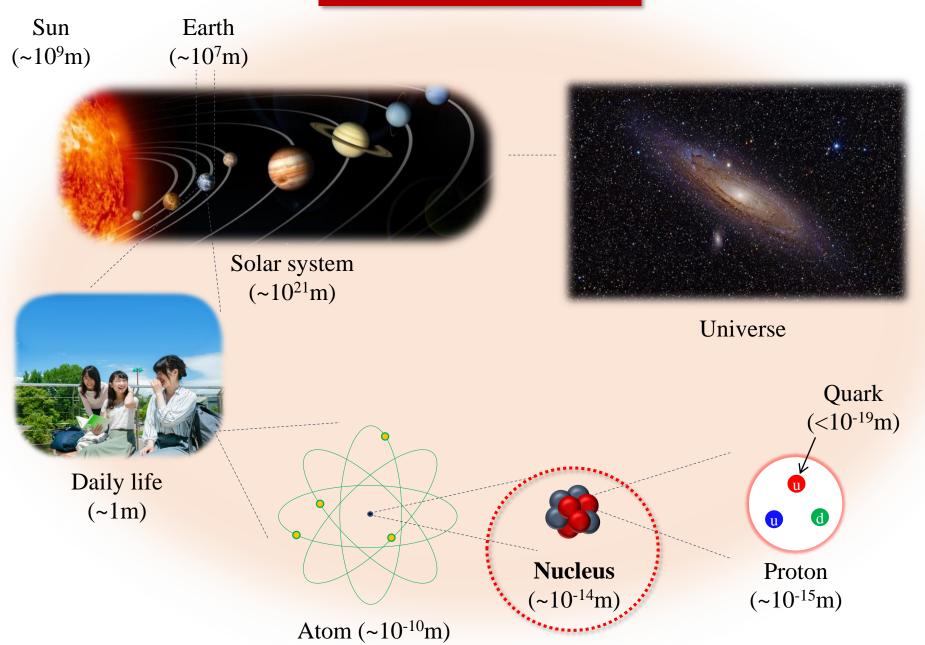


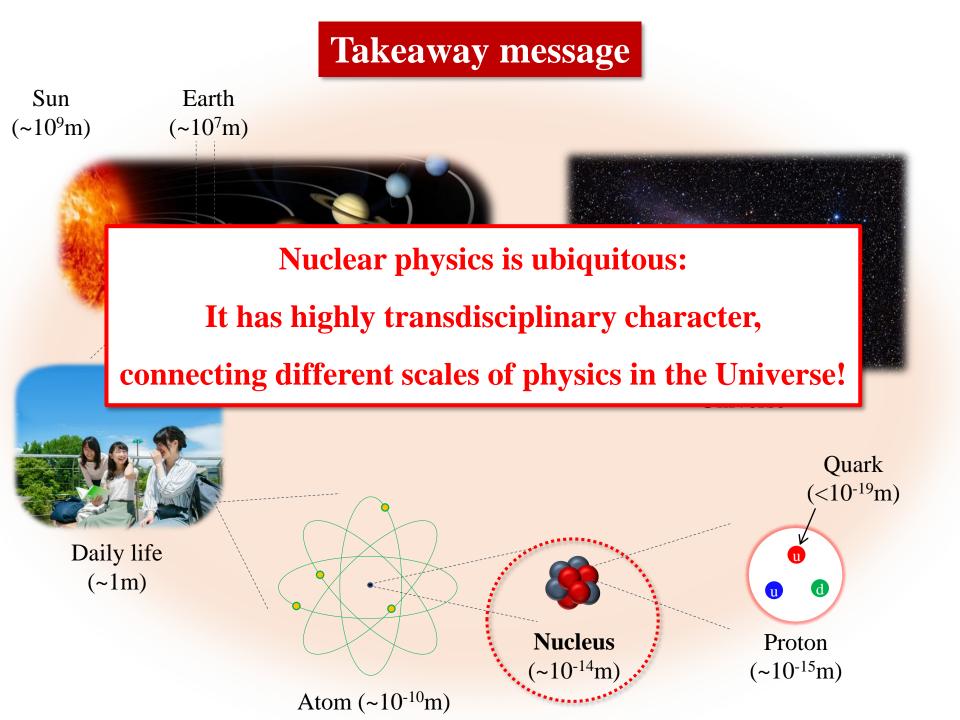
"Flat Earth" by iStock; "Fanciful view of ship sailing over edge of Earth" by Georgia Studies Images; "Star Wars Spaceship" by Valérian Pierret @Artstation

The <u>transdisciplinary character</u> is one of the fascinating points of Nuclear Physics

What I showed in my lectures are only a tiny part of the huge field! :)

Takeaway message





Message #1 for undergraduate course (BSc) students

All the subjects you learn (e.g., classical mechanics, electromagnetism, analytical mechanics, thermodynamics, statistical mechanics, quantum mechanics, etc.) are indispensable to explore the wonderful world of physics in the universe. It's like equipment for climbing. When completed, you'll see the breathtaking beauty of the nature that may change the rest of your life! :)

Study hard, be ambitious, and have fun!

Photo by Patrick Hendry on Unsplash

Message #2 for graduate course (MSc and PhD) students

Study what interests you the most. Dig it deeper and deeper. It doesn't matter what others say (of course it can be useful as an "opinion" though). Try to reach the deepest ever achieved and find a way to dig in further. This is your thesis work. It will form your "roots." Then, a thick and high trunk, a wide variety of branches, and abundant fruits of your research "tree" will grow up. The experience to grow up the tree allows you to plant other trees as well.

Be confident, anything can be interesting!

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